

Bandpass Calibration: Hands-on



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Atacama Large Millimeter/submillimeter Array
Expanded Very Large Array
Robert C. Byrd Green Bank Telescope
Very Long Baseline Array



What is Bandpass Calibration?

In general, the goal of calibration is to find the relationship between the observed visibilities, V_{obs} , and the true visibilities, V :

$$V_{ij}(t, \nu)_{\text{obs}} = V_{ij}(t, \nu) G_{ij}(t) B_{ij}(t, \nu)$$

where t is time, ν is frequency, i and j refer to a pair of antennas (i, j) (i.e., one baseline), G is the complex "continuum" gain, and B is the complex frequency-dependent gain (the "bandpass").

Bandpass calibration is the process of measuring and correcting the *frequency-dependent* part of the gains, $B_{ij}(t, \nu)$.

B_{ij} may be constant over the length of an observation, or it may have a slow time dependence.

Why is BP Calibration important?

Good bandpass calibration is a key to detection and accurate measurement of spectral features, especially weak, broad features.

Bandpass calibration can also be the limiting factor in dynamic range of continuum observations.

- Bandpass amplitude errors may mimic changes in line structure with ν
- ν -dependent phase errors may lead to spurious positional offsets of spectral features as a function of frequency, mimicking doppler motions
- ν -dependent amplitude errors limit ability to detect/measure weak line emission superposed on a continuum source. Consider trying to measure a weak line on a strong continuum with $\sim 10\%$ gain variation across the band.

Bandpass Calibration

- Determine the variations of phase and amplitude with frequency
- Account for slow time-dependency of the bandpass response
- We will arrive at antenna-based solutions against a reference antenna
 - In principle, could use autocorrelation data to measure antenna-based amplitude variations, but not phase
 - Most bandpass corruption is antenna-based, yet we are measuring $N(N-1)/2$ baseline-based solutions
 - Amounts to channel-by-channel self-cal

Bandpass Calibration: What makes good calibrators?

- Best targets are bright, flat-spectrum sources with featureless spectra
 - Although point-source not absolutely required, beware frequency dependence of resolved sources
 - If necessary, can specify a spectral index using *setjy*
- Don't necessarily need to be near science target on the sky

CASA Tasks for Bandpass Calibration

- First we will examine the bandpass structure with *plotms*
- Next use *gaincal* to measure time variation of phase
- Then use *bandpass* task
 - We will calibrate channel-to-channel variation (preferred method)
 - Alternatively, could fit a smooth function
 - Pay close attention to solutions; e.g. bright calibrators are rare, esp. at Band 9
- Use *applycal* to apply the bandpass solution to other sources

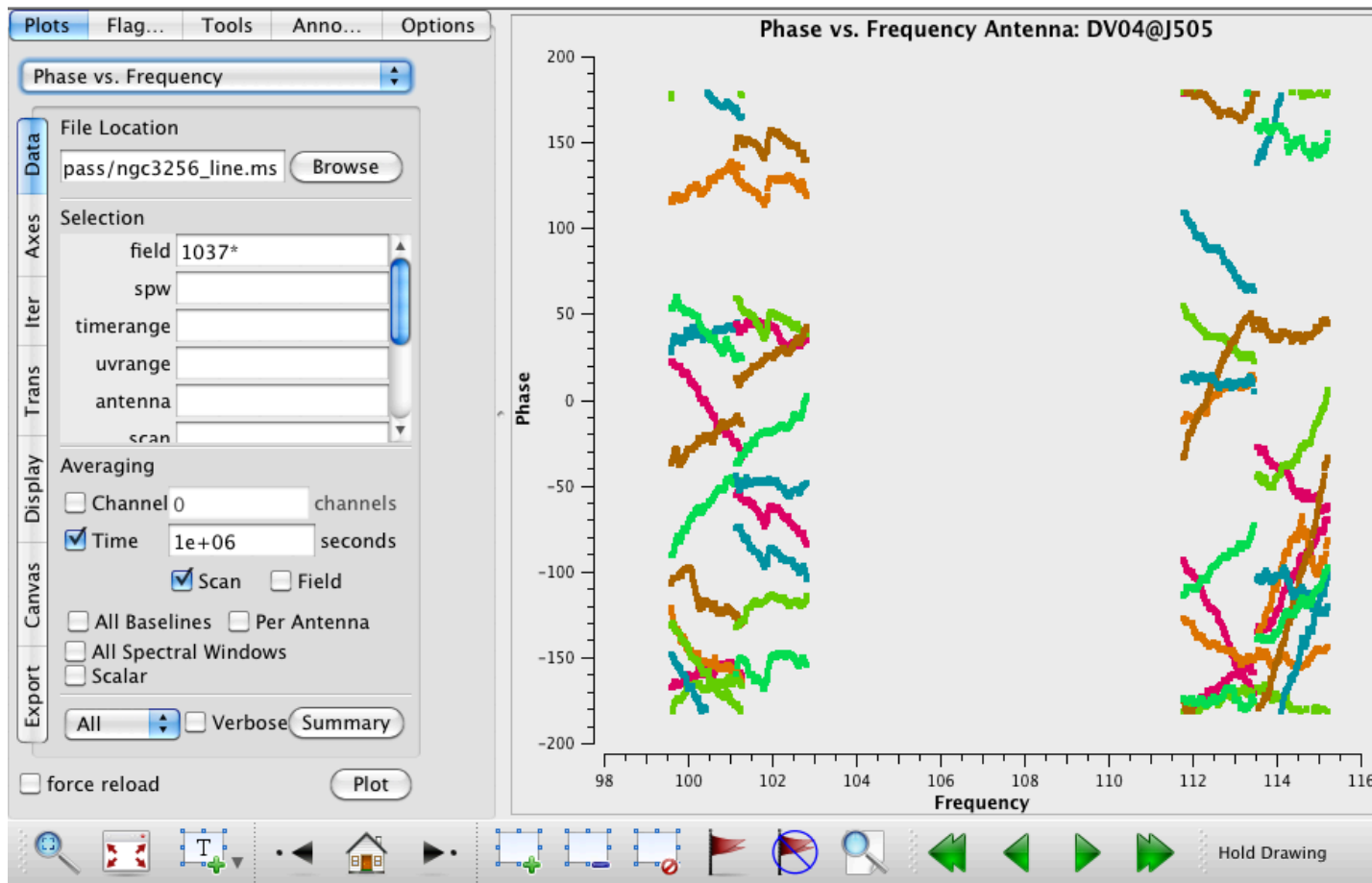
Bandpass Calibration

From here we will follow the commands in the script:

```
~/distrib/ngc3256/bandpass/bandpass.py
```

Bandpass Calibration

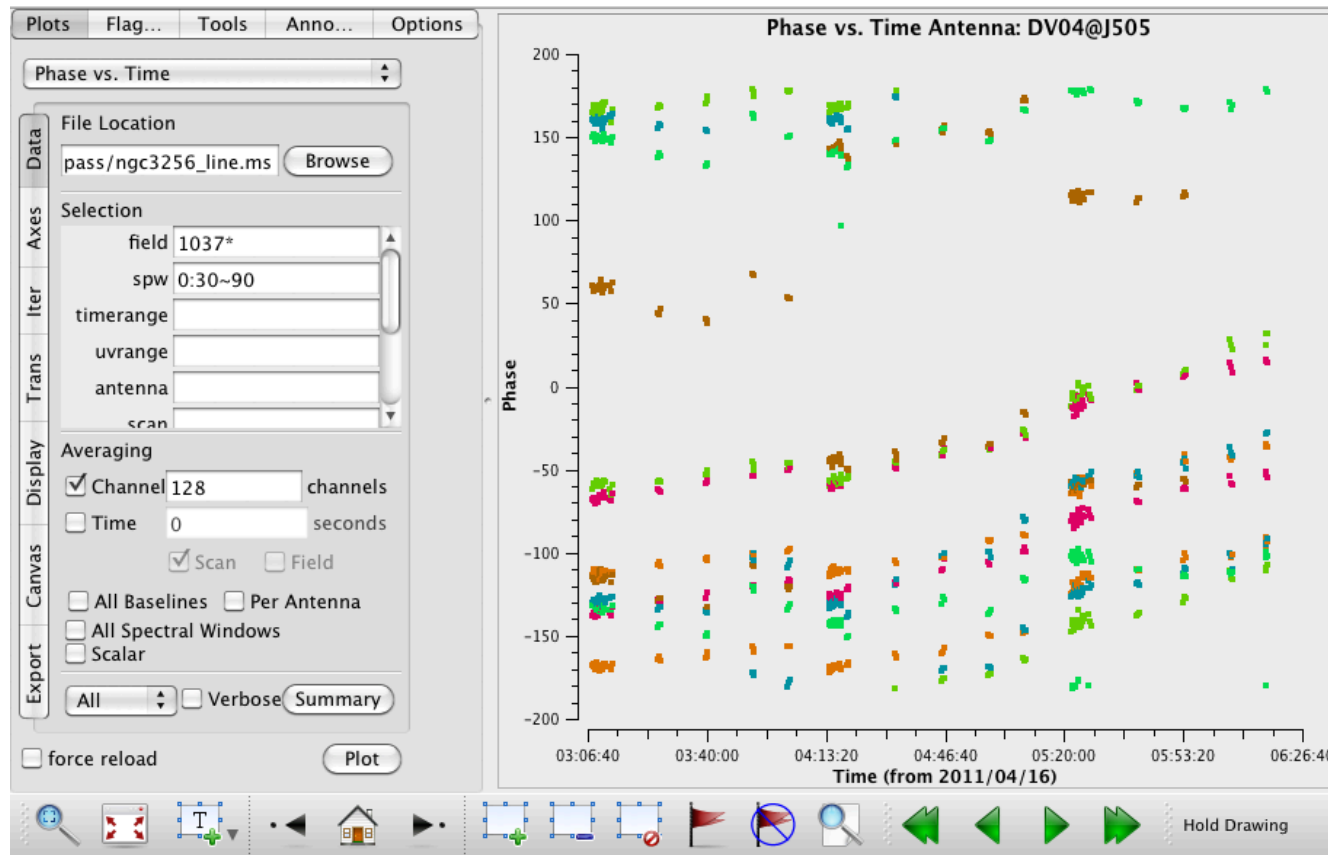
```
plotms(vis='ngc3256_line.ms', xaxis='freq', yaxis='phase', selectdata=True,  
       field='1037*', avgtime='1E6', avgscan=T, coloraxis='baseline', iteraxis='antenna')
```



Phase Before Bandpass Calibration

Bandpass Calibration

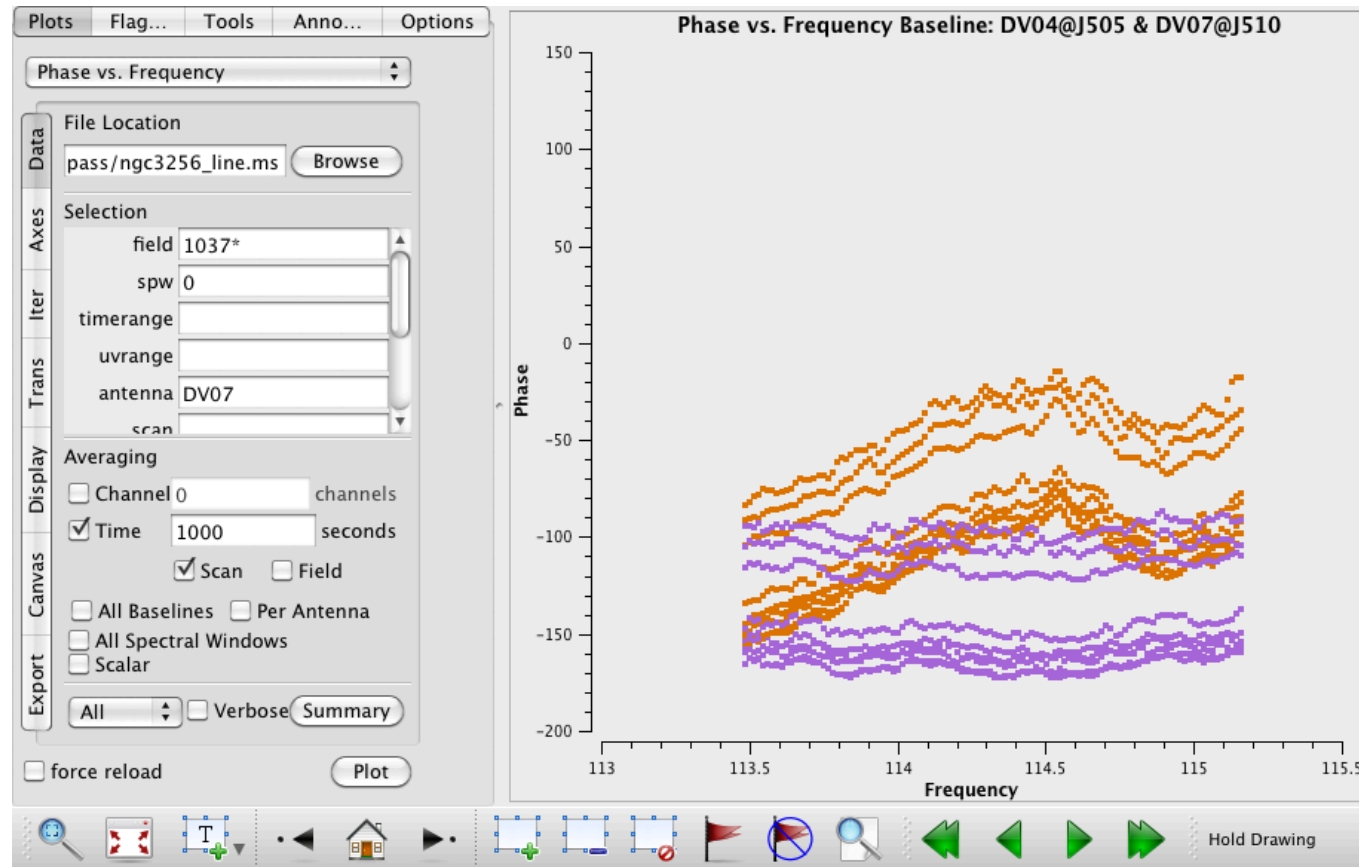
```
plotms(vis='ngc3256_line.ms', xaxis='time', yaxis='phase', selectdata=True, field='1037*',  
       spw='0:30~90', avgchannel='128', avgscan=T, coloraxis='baseline', iteraxis='antenna')
```



Phase variation with time for the BP calibrator before calibration.

Bandpass Calibration

```
plotms(vis='ngc3256_line.ms', xaxis='freq', yaxis='phase', selectdata=True, field='1037*',  
       avgtime='1E3', avgscan=T, spw='0', coloraxis='corr', antenna='DV07', iteraxis='baseline')
```

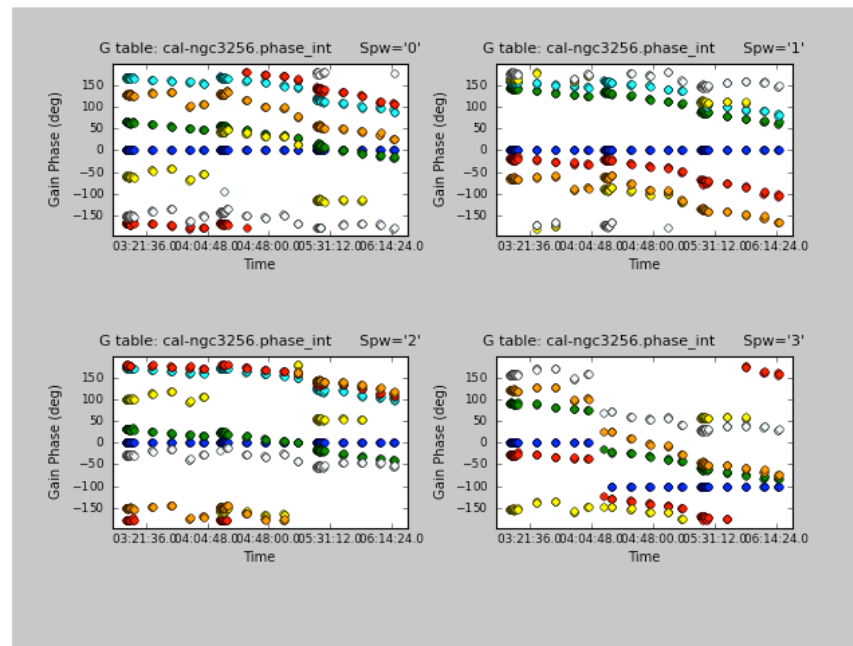


BP shape can change with time, but here it is stable over a few hours,
apart from a phase offset

Bandpass Calibration

Calculate the time-dependent phase variation:

```
os.system("rm -rf cal-ngc3256,phase_int")
gaincal(vis='ngc3256_line.ms',
        caltable='cal-ngc3256,phase_int',
        spw='*:40~80',
        field='1037*',
        selectdata=T,
        solint='int',
        refant='DV04',
        calmode='p')
```



Bandpass Calibration

```
CASA <B>: inp bandpass
-----> inp(bandpass)
# bandpass :: Calculates a bandpass calibration solution
vis = '' # Name of input visibility file
caltable = '' # Name of output gain calibration table
field = '' # Select field using field id(s) or field name(s)
spw = '' # Select spectral window/channels
intent = '' # Select observing intent
selectdata = False # Other data selection parameters
solint = 'inf' # Solution interval in time[,freq]
combine = '' # Data axes which to combine for solve (scan, spw, and/or field)
refant = '' # Reference antenna name(s)
minblperant = 4 # Minimum baselines _per antenna_ required for solve
minsnr = 3.0 # Reject solutions below this SNR (only applies for bandtype = B)
solnorm = False # Normalize average solution amplitudes to 1.0
bandtype = 'B' # Type of bandpass solution (B or BPOLY)
  fillgaps = 0 # Fill flagged solution channels by interpolation

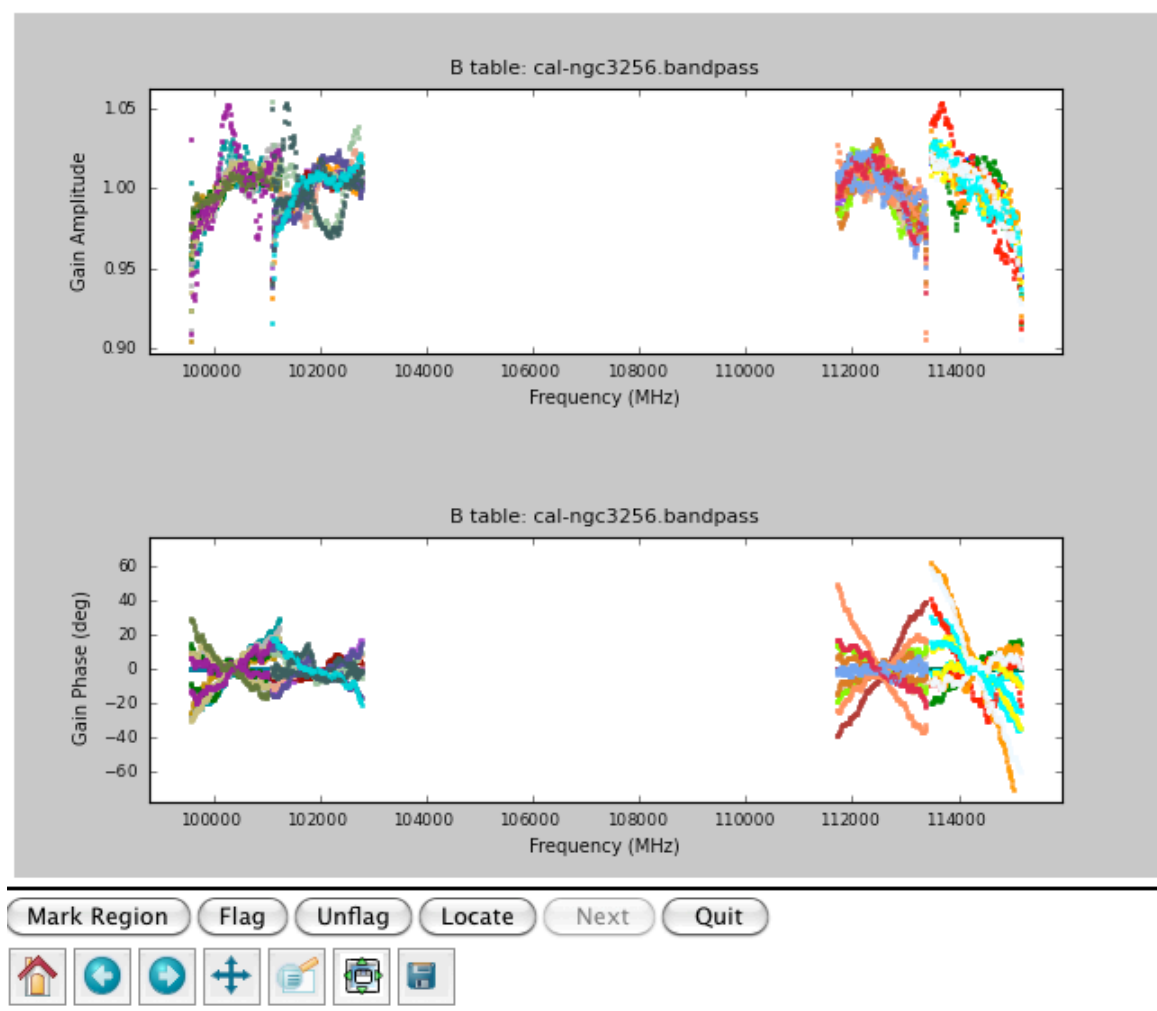
smodel = [] # Point source Stokes parameters for source model.
append = False # Append solutions to the (existing) table
gaintable = [''] # Gain calibration table(s) to apply on the fly
gainfield = [''] # Select a subset of calibrators from gaintable(s)
interp = [''] # Interpolation mode (in time) to use for each gaintable
spwmap = [] # Spectral windows combinations to form for gaintables(s)
gaincurve = False # Apply internal VLA antenna gain curve correction
opacity = [] # Opacity correction to apply (nepers), per spw
parang = False # Apply parallactic angle correction
async = False # If true the taskname must be started using bandpass(...)
```

Bandpass Calibration

```
bandpass(vis = 'ngc3256_line.ms',  
         caltable = 'cal-ngc3256.bandpass',  
         gaintable = 'cal-ngc3256.phase_int',  
         field = '1037*',  
         minblperant=3,  
         minsnr=2,  
         solint='inf',  
         combine='scan',  
         bandtype='B',  
         fillgaps=1,  
         refant = 'DWO4',  
         solnorm = T)
```

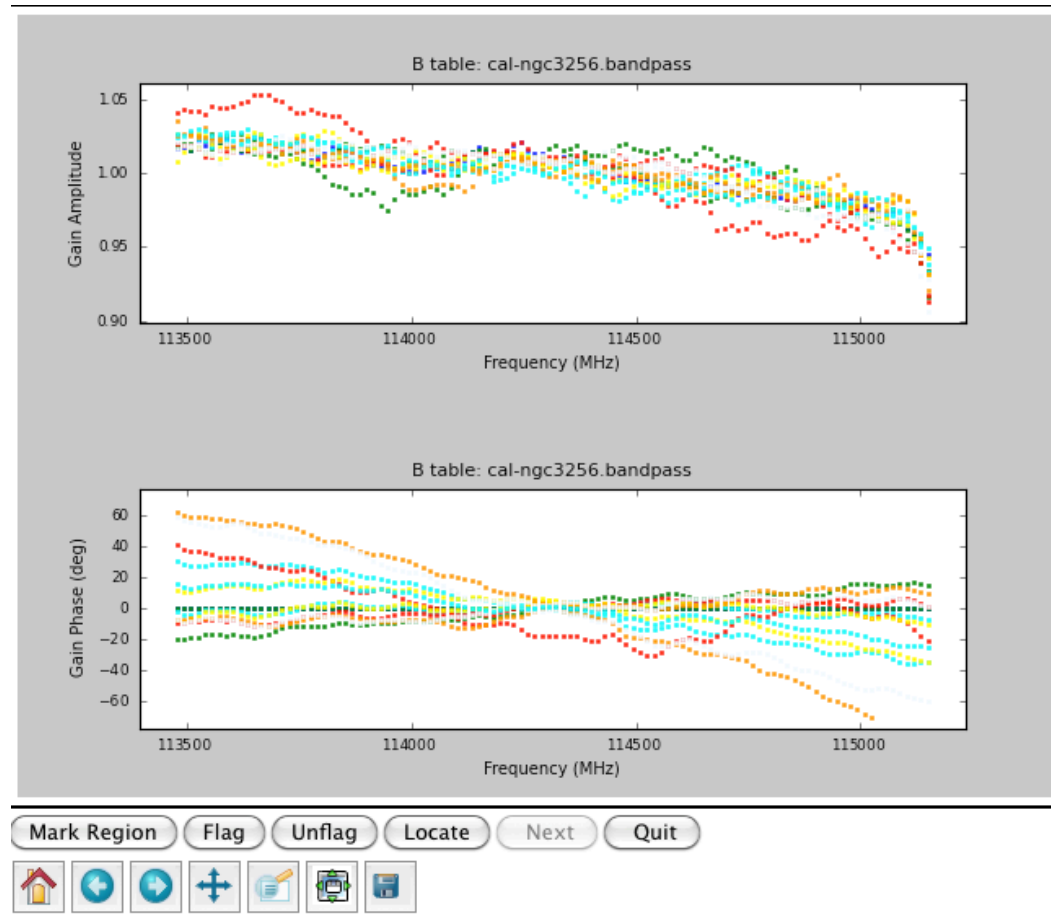
Bandpass Calibration

We determine solutions ...



Bandpass Calibration

Looking at only spw 0 ...



Bandpass Calibration

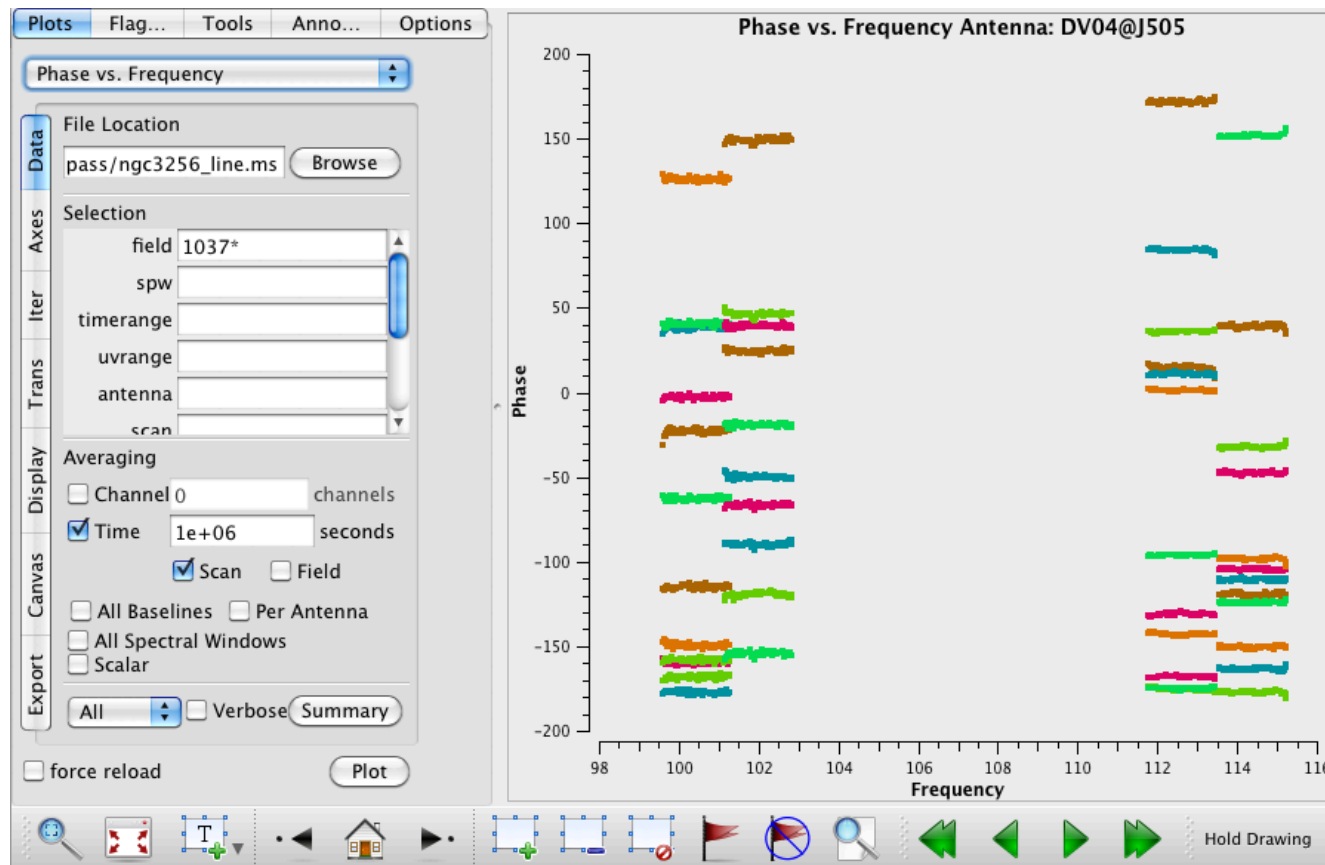
Apply the calibration ...

```
applycal(vis="ngc3256_line.ms",  
         field="",  
         spw="",  
         gaintable="cal-ngc3256.bandpass",  
         gainfield="1037*",  
         interp="nearest,nearest")
```


Bandpass Calibration

Apply the bandpass calibration and then examine the phase of BP calibrator with solutions applied

```
plotms(vis='ngc3256_line.ms', xaxis='freq', yaxis='phase', selectdata=True, field='1037*',  
       avgtime='1E6', avgscan=T, coloraxis='baseline', iteraxis='antenna', ydatacolumn="corrected")
```



Bandpass Calibration

... and amplitude of bandpass calibrator with solutions applied

