## ALMA Data and Archives what to expect when you're expecting (ALMA data)!



### **Cassie Reuter**

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Associated Universities. Inc. Atacama Large Millimeter/submillimeter Array Expanded Very Large Array



## **The Condensed Version**



- Data delivered after passing Quality Assurance (QA)
- Download data from Archive Query and Request Handler tools on the ALMA Science Portal
- Delivered data include:
  - Calibration tables and diagnostics
  - Preliminary images (better products may be possible with more careful continuum identification & interactive cleaning)
- Sections 11, 12, 14, and Appendix C of ALMA Technical Handbook

# This talk will be available online for reference after this workshop.



An approximate\* ALMA timeline



## Goals of Quality Assurance (QA) Process



- Ensure reliable final data product
  - Desired sensitivity (as specified by PI)
  - Desired resolution (as specified by PI)
- Ensure calibration and QA imaging free from major artifacts
- Warning: Errors in PI-supplied parameters are outside scope of QA process, including:
  - Incorrect source coordinates
  - Inadequate frequency specification
  - Inadequate sensitivity limits



See <u>ALMA Technical Handbook</u> for details.



## **During Observations – QA0**

- Monitoring of on-the-fly calibration and system performance
- Rapidly-varying parameters (~EB timescales)
  - Atmospheric effects
  - Antenna issues
  - Front/back-end issues
  - Connectivity issues
- Tolerances for each are explicitly laid out
  - No fewer than 40 antennas in 12m array
  - Bandpass calibrator is strong enough
- Quick reduction may be run to check flux measurements and phase stability



**Between Observations – QAI** 

- "Regular array maintenance" timescales
- Slowly Varying Parameters (~>week timescales)
- General array calibration
  - Baseline measurements
  - Delays
- Antenna Calibrations
  - All-sky pointing
  - Focus curves
  - Beam patterns, etc.
- Observatory Calibrator Surveys
  - Solar-system and quasar flux monitoring



## After Observations – QA2

- Calibration by pipeline (~70%), checked by NRAO staff.
- Final QA calibration checks include:
  - Bandpass quality

**Review period** 

- Flux scale calibration
- Phase transfer and astrometry
- Final data characterization includes:
  - RMS noise in target images
  - Spatial resolution for specified weighting scheme
  - Coverage and time on target (after flagging)
  - And more! (see technical handbook)
- Information about QA review is aggregated for delivery in the QA2 Report



## After Delivery – QA3

- Additional QA stage possibly triggered by PI reporting any issues underlying:
  - Data, observing procedure, calibration
- Re-evaluation of calibrated data products
  - Only occurs if QA0 -> QA2 miss something
- Proprietary period extension (within two months of delivery)



• After two months, extension only until fix is delivered





## **Optional emails**



#### Science Highlights - Molecular Gas Within the Supernebula of the Dwarf Galaxy NGC 5253

by <u>Portal Admin</u> — last modified Nov 30, 2017 09:38 PM



One of the areas of extragalactic research which makes great dwarf galaxies. In a <u>recent study</u> by Dr. Jean Turner and her co and <sup>13</sup>CO(3-2) emission (Cloud D1) from the core of a giant st third of the galaxy's infrared luminosity and is in proximity to c analysis, the region is estimated to have 1400-1800 O stars..

Full Summary...

If you want to receive emails about the status of your data before its delivered:



In case of problems with the registration, please use this Web form to contact us You may find a solution to your problem in the Support Center/Knowledgebase

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## **Data Delivery!**

- Pls get two delivery emails:
  - 1. Sent when an individual observation passes QA2

- Link to data archive with products
- 2. From NAASC staff
  - Fully-calibrated MS
  - Calibration and Imaging Report
  - ADMIT products
- Triggers Start of Proprietary Period
- Publication Requirements:

**Review period** 

- ALMA acknowledgement
- NRAO specific acknowledgement



## How to find the archive

Go to the science portal: https://almascience.nrao.edu

- Click on "Data" and select "Archive"

Atacama Large Millimeter/submillimeter Array In search of our Cosmic Daigins	NRAO Associated Universities Inc. Log in
C About Science Proposing Observing Data Processing Tools Documentation Help	Search Site P
Archive Archive Query	
Documentation We provide a comprehensive ALMA Science Archive Manual.	
Data delegation	
Pls can grant access to one of their projects to a registered ALMA user by logging into the Science Portal, going to the user prof delegees in the "Project delegation" tab.	ile page in the top right corner and then adding
Curle 0 content	

Please go here to see the content of the Cycle 0 deliveries.



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eso nrao naoj



#### ALMA Science Archive Query

#### http://almascience.nrao.edu/aq/

Query Form Results Table			
Search Reset			Query Help
Position	Energy	Time	Polarisation
Source name (Sesame) Source name (ALMA) RA Dec	Frequency Bandwidth Spectral resolution Band	Observation date Integration time	Polarisation type
<b>Observation</b> Water vapour	Project code Project title Project title PI name	Project code Project code. Description Project code, in the form YYYY.NNNNN.C.AAA, where: Example 2010.2.00010.N 2010.* 2010.* 2010.?.*.CSV *.CSV !(*.CSV   *.SIM)	Options View: ⊙raw data ⊙ project ✓ public data only ✓ science observations only





## **Archive Query**

Query Form

Results Table

Submit download request

Results Bookmark Export Table Results Help

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	Project code	Source name	RA	Dec	Band	Integration	Release date 🔺	Velocity resolution	Frequency support
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☑	2012.1.00090.S	S2CLS_UDS156	02:18:24.23	-05:22:53.4	7	8.836	2014-11-07T09:35:00.000	27236.96	336.00351.99GHz
☑	2012.1.00090.S	S2CLS_UDS160	02:18:23.86	-05:11:36.2	7	8.842	2014-11-07T09:35:00.000	27236.96	336.00351.99GHz
	2012.1.00090.S	S2CLS_UDS168	02:18:20.34	-05:31:41.6	7	8.843	2014-11-07T09:35:00.000	27236.96	336.00351.99GHz
☑	2012.1.00090.S	S2CLS_UDS199	02:18:07.38	-04:44:11.7	7	8.812	2014-11-07T09:35:00.000	27236.96	336.00351.99GHz
	2012.1.00090.S	S2CLS_UDS204	02:18:03.01	-05:28:39.8	7	8.873	2014-11-07T09:35:00.000	27236.96	336.00351.99GHz
	2012.1.00090.S	S2CLS_UDS216	02:17:56.80	-04:52:39.6	7	8.82	2014-11-07T09:35:00.000	27236.96	336.00351.99GHz
	2012.1.00090.S	S2CLS_UDS252	02:17:37.79	-05:20:10.2	7	8.827	2014-11-07T09:35:00.000	27236.96	336.00351.99GHz
	2012.1.00090.S	S2CLS_UDS286	02:17:25.76	-05:25:36.5	7	9.657	2014-11-07T09:35:00.000	27236.96	336.00351.99GHz
	2012.1.00090.S	S2CLS_UDS292	02:17:21.85	-05:19:03.3	7	8.815	2014-11-07T09:35:00.000	27236.96	336.00351.99GHz
	2012.1.00090.S	S2CLS_UDS298	02:17:19.90	-05:09:36.4	7	9.55	2014-11-07T09:35:00.000	27236.96	336.00351.99GHz
	2012.1.00090.S	S2CLS_UDS334	02:17:02.81	-04:57:24.9	7	8.856	2014-11-07T09:35:00.000	27236.96	336.00351.99GHz
	2012.1.00090.S	S2CLS_UDS408	02:16:22.59	-05:11:06.0	7	8.819	2014-11-07T09:35:00.000	27236.96	336.00351.99GHz
	2012.1.00090.S	S2CLS_UDS421	02:16:17.62	-05:09:02.0	7	8.803	2014-11-07T09:35:00.000	27236.96	336.00351.99GHz
	2012.1.00090.S	S2CLS_UDS47	02:19:24.97	-05:09:19.9	7	8.785	2014-11-07T09:35:00.000	27236.96	336.00351.99GHz





## Archive Query: more columns

	Project code	1	Show all columns Reset colum	n orde	order alphabetically		Frequency support	
r:			Project code		Project code, in the form YYYY.NNNNN.C.AAA, where:			
	2012.1.00090.S		Source name		Name of the source as registered in the ASDM. Partial matches through wildcards (?, *), and boolean OR expressions (" "), can be used.	1	335.99351.99GHz	
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	2012.1.00090.S		🔽 Dec	deg	Declination of the field pointing.	ł	335.99351.99GHz	
	2012.1.00090.S		Band		ALMA receiver band.	1	335.99351.99GHz	
	2012.1.00090.S		Integration	s	Aggregated integration time for the field in the ASDM.		335.99351.99GHz	
	2012.1.00090.S		Release date				335.99351.99GHz	
	2012.1.00090.S		Velocity resolution	m/s	Estimated velocity resolution from all the spectral windows, from frequency resolution.	1	335.99351.99GHz	
	2012.1.00090.S		Frequency support	GHz	All frequency ranges used by the field	1	335.99351.99GHz	
	2012.1.00090.S					- 1	335.99351.99GHz	
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	2012.1.00090.S 2012.1.00090.S 2012.1.00090.S		Spatial resolution Frequency resolution	kHz	Estimated frequency resolution from all the spectral windows, using median values of		335.99351.99GHz 335.99351.99GHz 335.99351.99GHz	
	2012.1.00090.S 2012.1.00090.S 2012.1.00090.S 2012.1.00090.S		Spatial resolution Frequency resolution	kHz	Estimated frequency resolution from all the spectral windows, using median values of channel widths.		335.99351.99GHz 335.99351.99GHz 335.99351.99GHz 335.99351.99GHz	
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## **Archive Query**

Query Form

Results Table

Submit download request

Results Bookmark Export Table Results Help

Showing 3	lowing 30 rows (30 before filtering).								
	Project code	Source name	RA	Dec	Band	Integration	Release date 🔺	Velocity resolution	Frequency support
Filter:								 	
☑	2012.1.00090.S	S2CLS_UDS110	02:18:48.44	-05:18:05.0	7	9.326	2014-11-07T09:35:00.000	27236.96	336.00351.99GHz
☑	2012.1.00090.S	S2CLS_UDS156	02:18:24.23	-05:22:53.4	7	8.836	2014-11-07T09:35:00.000	27236.96	336.00351.99GHz
☑	2012.1.00090.S	S2CLS_UDS160	02:18:23.86	-05:11:36.2	7	8.842	2014-11-07T09:35:00.000	27236.96	336.00351.99GHz
	2012.1.00090.S	S2CLS_UDS168	02:18:20.34	-05:31:41.6	7	8.843	2014-11-07T09:35:00.000	27236.96	336.00351.99GHz
2	2012.1.00090.S	S2CLS_UDS199	02:18:07.38	-04:44:11.7	7	8.812	2014-11-07T09:35:00.000	27236.96	336.00351.99GHz
	2012.1.00090.S	S2CLS_UDS204	02:18:03.01	-05:28:39.8	7	8.873	2014-11-07T09:35:00.000	27236.96	336.00351.99GHz
	2012.1.00090.S	S2CLS_UDS216	02:17:56.80	-04:52:39.6	7	8.82	2014-11-07T09:35:00.000	27236.96	336.00351.99GHz
	2012.1.00090.S	S2CLS_UDS252	02:17:37.79	-05:20:10.2	7	8.827	2014-11-07T09:35:00.000	27236.96	336.00351.99GHz
	2012.1.00090.S	S2CLS_UDS286	02:17:25.76	-05:25:36.5	7	9.657	2014-11-07T09:35:00.000	27236.96	336.00351.99GHz
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	2012.1.00090.S	S2CLS_UDS421	02:16:17.62	-05:09:02.0	7	8.803	2014-11-07T09:35:00.000	27236.96	336.00351.99GHz
	2012.1.00090.S	S2CLS_UDS47	02:19:24.97	-05:09:19.9	7	8.785	2014-11-07T09:35:00.000	27236.96	336.00351.99GHz



## **Download requests**



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## QA2 Data Products Package: the processed data



After un-tarring the processed data we have a directory tree: Science goal Project code



Data delivery products...

QA2 Data Products Package: the processed data



#### **Calibration Directory:**

Calibration tables generated by the pipeline

Contains manual flagging commands, continuum selection, flux measurements for calibrators

#### |-- calibration | |- member.uid\_\_\_A001\_X1299\_X39.hifa\_calimage.auxproducts.tgz | |-- member.uid\_\_\_A001\_X1299\_X39.session\_1.auxcaltables.tgz | -- member.uid\_\_\_A001\_X1299\_X39.session\_1.caltables.tgz | -- uid\_\_\_A002\_Xc8ed15\_X1a9.ms.calapply.txt | -- uid\_\_\_A002\_Xc8ed15\_X1a9.ms.flagversions.tgz | -- uid\_\_\_A002\_Xc8ed15\_X1a9\_target.ms.auxcalapply.txt



All flags will be restored during calibration



## QA2 Data Products Package: the processed data

#### Calibration Products: Log of equivalent CASA commands (non-executable)

#### log -- member.uid A001 X1299 X39.hifa calimage.casa commands.log member.uid A001 X1299 X39.README.txt product member.uid A001 X1299 X39.SOURCE sci.spw25 27 29 31.cont.I.pb.fits A001 X1299 X39.SOURCE sci.spw25 27 29 31.cont.I.pbcor.fits -- member.uid -- member.uid A001 X1299 X39.SOURCE sci.spw25.cube.I.mask.fits A001 X1299 X39.SOURCE sci.spw25.cube.I.pbcor.fits -- member.uid A001 X1299 X39.SOURCE sci.spw25.cube.I.pb.fits.gz -- member.uid A001 X1299 X39.J0117p1418 ph.spw31.mfs.I.pbcor.fits -- member.uid -- member.uid A001 X1299 X39.J0117p1418 ph.spw31.mfs.I.pb.fits.gz

# Directions to access QA comments and restoration instructions

Calibration and Target images produced during reduction (may be representative)





# QA2 Data Products Package: the processed data

## Calibration Scripts and Weblog:

# Weblog contains plots and images from reduction and imaging. Unpack this for lots of information! \

qa
 `-- member.uid\_\_\_A001\_X1299\_X39.hifa\_calimage.weblog.tgz

- -- script
  - -- member.uid \_\_\_\_ A001\_X1299\_X39.calimage.pipeline\_manifest.xml
  - -- member.uid \_\_\_\_\_ A001\_X1299\_X39.calimage.product\_rename.txt

  - -- member.uid \_\_\_\_\_A001\_X1299\_X39.hifa\_calimage.casa\_pipescript.py

  - -- member.uid \_\_\_A001\_X1299\_X39.scriptForPI.py

#### Run scriptForPI.py to restore calibration -

Commands to re-run the pipelihe





## **Resources After Delivery**

- HelpDesk
- Face to Face visits in Charlottesville: https://science.nrao.edu/ facilities/alma/visitors-shortterm

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ALMA in the P	ress	se	ets. To meet this need,	, the NAASC runs a visit	or	stance with the technical	acrosts of				
Workshops & Tutorials		program. In some cases short-term visits may be scheduled for assistance with the technical aspects of ALMA proposals or designing scheduling blocks (SBs). Scientists carrying out archival projects may also take advantage of the visitor program.									
News, Memos Outreach	&	The NAASC will provide assistance to up to three visitors from each scheduled or observed project, though funds for travel are more restricted. An experienced investigator must accompany any student who is new to radio interferometry.									







## For more info:

https://almascience.nrao.edu/

The Atacama Large Millimeter/submillimeter Array (ALMA), an international astronomy facility, is a partnership of Europe, North America and East Asia in cooperation with the Republic of Chile. ALMA is funded in Europe by the European Organization for Astronomical Research in the Southern Hemisphere (ESO), in North America by the U.S. National Science Foundation (NSF) in cooperation with the National Research Council of Canada (NRC) and the National Science Council of Taiwan (NSC), and in East Asia by the National Institutes of Natural Sciences (NINS) of Japan in cooperation with the Academia Sinica (AS) in Taiwan. ALMA construction and operations are led on behalf of Europe by ESO, on behalf of North America by the National Radio Astronomy Observatory (NRAO), which is managed by Associated Universities, Inc. (AUI), and on behalf of East Asia by the National Astronomical Observatory of Japan (NAOJ). The Joint ALMA Observatory (JAO) provides the unified leadership and management of the construction and operation of ALMA.



#### Monitor Project Status: SnooPl

## Listing of Pl'ed projects

JALMA	SnooP	PI	John Smith, EU Executive, EU ARC	<ul> <li>All proje</li> <li>Contact</li> </ul>	cts scientist	
PI	Co-I					
4	Projects			Sea	rch	Q
Proje	ect code 🛦		Project Title 🔺		Status 🔺	Grade 🔺
2015	5.1.09876.S		A most inspired project title		*	A
2013	.1.04567.S	Ø	Observing stars, planets, nebulae, open clusters, glo galaxies and galaxy clusters with ALMA	obular	•	с
2013	8.1.06789.S	0	Observing the centre of the galaxy with ALMA SgrB2_a_03_TE SgrB2_a_03_TC SgrB2_a_03_TC SgrB2_a_03_TC		$\mathbf{\hat{\mathbf{V}}}$	В
			SgrB2_a_03_TP All data taken 3c454.3_SgrB2_a_03_TP 🗸	Check of yo	observing ur projects	status for all at a glance



#### **Monitor Project Status: SnooPl**

Listing of Pl'ed projects



- Project has been submitted (Phase 1)
- 🔀 Rejected at proposal review stage
- Approved but SBs not yet prepared
- SBs prepared but are not yet in the observing queue
- SBs are in the observing queue but not yet taken
- 🔺 Some data has been taken
- 🖋 All the data has been taken
- Completed and delivered
- Partially completed and all data taken has been delivered
- Project is timed out
- × Project cancelled and not observed
- Project that are not to be observed



Onknown status

## Monitor Project Status: SnooPl Single Project View:

SnooPl	John Smith, EU Executive, EU ARC	<ul> <li>All project</li> <li>Contact set</li> </ul>	ts cientist
Project Code: 2013.1.06789.S	Full Proposal [pdf]. Grade B. ARC	node: <mark>Czech</mark> . Con	tact scientist: Jack Black Project report.
♣ 2013.1.06789.S		~	Exec.
Observing the centre of the	e galaxy with ALMA		
🗞 ObsUnitSet			
SG OUS (CH3CN 5-4 & is 0, HNC 1-0 map)	otopologue, H2CS 3-2, HCO+ 1-0, HC	N 1-	
& Group OUS			
🗞 🛛 Member OUS (SgrE	32)		
SgrB2_a_03_TP		×	41/40
🗞 Member OUS (SgrB	32)		
SgrB2_a_03_TC			4/3
🗞 Member OUS (SgrB	32)	<b>a</b>	
SgrB2_a_03_7M			4/4
🗞 Member OUS (que	ry)	×	
3c454.3_SgrB2_a	_03_TP		6/1
& Member OUS (SgrE	32)	¢ŝ	
SgrB2_a_03_TE	Click here to f	ind QA R	eport

INRAC

## **Monitor Project Status: SnooPl**

Single Project View:

- a set of gears indicate that the OUS is being processed;
- a truck indicates that the OUS has been delivered
- an ambulance indicates an OUS for which QA3 is in progress.





#### Monitor Project Status: Optional emails

- Subscribe to email notification for updates on changes to project status through your Science Portal user profile
  - ...
  - Phase2Submitted
  - Running
  - Partially Observed
  - Fully Observed
  - Pipeline Processing
  - ...
- With or without optional emails, PIs always receive notification when new data are available





## **Optional emails**



Atacama Large Millimeter/submillimeter Array

In search of our Cosmic Origi

About Science Processing Tools Documentation Proposing Observing Data

Help

Search Site

Observatory News	EU ARC News	Status		
Additional Information for Cycle 6 Proposals Feb 01, 2018	Researcher position available at the Nordic ARC node Jan 10, 2018	ALMA Cycle 5 Config Schedule		
New Science Verification data are now available for download Jan 22, 2018	Post-doc position available at the Italian ARC-node Dec 20, 2017	Refereed publications: 916		
Announcement of intent to release a new installment of Science Verification data	2017 European Radio Interferometry School May 11, 2017	Last observed source: W43-MM1 Current configuration: C43-5		
More	More	More		

#### Science Highlights - Molecular Gas Within the Supernebula of the Dwarf Galaxy NGC 5253



One of the areas of extragalactic research which makes great use of ALMA's resolution and sensitivity is the study of the molecular gas properties of dwarf galaxies. In a recent study by Dr. Jean Turner and her collaborators, they make use of Band 7 ALMA observations to detect warm <sup>12</sup>CO(3-2) and <sup>13</sup>CO(3-2) emission (Cloud D1) from the core of a giant star-forming region, in the dwarf galaxy NGC 5253. This "supernebula" is the source of onethird of the galaxy's infrared luminosity and is in proximity to optical clusters with measured stellar ages of ~ 1 Myr. From radio recombination line analysis, the region is estimated to have 1400-1800 O stars..

Full Summary ...





## **Optional emails**



### ALMA Central Authentication Service (CAS)

A L N A A	Enter your NetID and Password	For security reasons, please Log Out and Exit your web browser when you are done accessing services that require authentication!			
username an <u>d</u>	NetID:	If you don't have an account, you can create one in the following link: Registration web form			
password	Warn me before logging me into other	If you forgot you account ID, you can go to the following link: Forgot account ID page			
	sites.	If you want to reset your password, you can go to the following link: Reset password page			
	LOGIN clear	You may find a solution to your problem in the Support Center/Knowledgebase: Helpdesk			

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