



Cycle 8 Capabilities

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ALMA: Extension of Proprietary Times

- Due to the COVID-19 crisis and its impact on the global astronomical community, the ALMA Director has made the following decision regarding proprietary times of ALMA data sets.
- The proprietary period of all MOUSs that have been delivered to principle investigators (PIs), but have not become public by 19 March 2020, will be extended by three months. This implies that the proprietary period of all these ALMA data will effectively be set to 15 months, or nine months for DDT projects. 19 March was the last day that ALMA was observing before Cycle 7 observations were suspended.
- For data sets that already reached the end of the proprietary time between March 19 and today, the proprietary time will be reinstated and extended for an additional 3 months. Please be aware that ALMA cannot guarantee that these data sets have not already been downloaded by archive users.
- PIs of data sets that do not fall in the category defined above, who feel that their ability to work on their data within the regular proprietary period is impacted by the COVID-19 crisis, are requested to submit a ticket to the “Proprietary Period Extension Requests” Department of the [ALMA Helpdesk](#). ALMA will handle these requests in a fair and lenient manner as we are committed to support our users as much as possible during this stressful time.

- Updates on this policy may be posted on the Science Portal in the future.

ALMA in Cycle 8

In Cycle 8 we continue to operate as what is been defined as “Steady State Operations”

- **In Cycle 8, the following technical capabilities will be available for the first time:**
 - Solar observations in Band 5.
 - VLBI observations of faint science targets (correlated flux density <500 mJy within an unresolved core on ALMA baselines up to 1 km). These observations will be done in passive phasing mode, where it is recommended to have a bright calibrator within 5 deg of the science target.
 - High-frequency observations (Bands 9 and 10) with the stand-alone 7-m Array
 - Mosaicking of continuum linear polarization observations (Bands 3 to 7)
 - Spectral scans with the 7-m Array

ALMA Capabilities

- **The Cycle 8 capabilities are fully described in Appendix A of the ALMA Proposers Guide available at:**
 - (<https://almascience.nrao.edu/documents-and-tools>)
- **Number of antennas**
 - At least forty-three (43) antennas in the 12-m Array
 - At least ten (10) 7-m antennas (for short baselines) and three (3) 12-m antennas (for making single-dish maps) in the ACA
- **Receiver bands**
 - Receiver Bands 3, 4, 5, 6, 7, 8, 9, and 10 (wavelengths of about 3.1, 2.1, 1.6, 1.3, 0.87, 0.74, 0.44, and 0.32 mm, respectively)
- **12-m Array Configurations**
 - **Maximum baselines for the antenna configurations will vary from 0.16 km to 8.5 km. Configurations C-9 and C-10 will not be offered in Cycle 8.**
 - Maximum baselines of 3.6 km for Bands 8, 9 and 10
 - Maximum baselines of 8.5 km for Bands 3 to 7
 - Files containing **representative** antenna configurations for the 12-m and 7-m arrays suitable for Common Astronomy Software Applications (CASA) simulations are available from the ALMA Science portal (<http://almascience.org/documents-and-tools/cycle8/alma-configuration-files>)

ALMA Capabilities

- **Spectral line, continuum, and mosaic observations**
 - Spectral line and continuum observations with the 12-m Array and the 7-m Array in all bands
 - Single field interferometry (all bands) and mosaics (Bands 3 to 9) with the 12-m Array and the 7-m Array
 - Single-dish spectral line observations in Bands 3 to 8
- **Polarization**
 - Single pointing, on-axis, full, linear and circular polarization for both continuum and full-spectral-resolution observations in Bands 3, 4, 5, 6, and 7 on the 12-m Array.
 - Linear polarization imaging of a compact source on-axis in both continuum and full spectral resolution modes is feasible at the level of 0.1% (3 sigma) fractional polarization for the very brightest calibrators, and 0.2% (3 sigma) level for a typical observation.
 - The minimum detectable degree of circular polarization is 1.8% of the peak flux for both continuum and full spectral resolution observations. **(NOTE that Zeeman observations have not been fully commissioned and should be discouraged from proposing.)**
 - **Mosaicking of continuum linear polarization observations (Bands 3 to 7).**
- **Observing Time:**
 - 4300 hours for successful proposals of PI programs expected on the 12m Array (includes DDT, Cycle 4-7 Carryover and resubmissions)
 - 3000 hours available on the ACA*
 - 3000 hours available on the Total Power Array*
 - ~2500? Hours of ACA time will be available through the Supplemental Call in mid-Cycle 8.

ALMA Capabilities: ACA

- **Observing Time:**

- We strongly encourage ACA only observations in a wide range of science and large observing times.
- “medium size” proposals of about 10-30 hours are likely to get time.

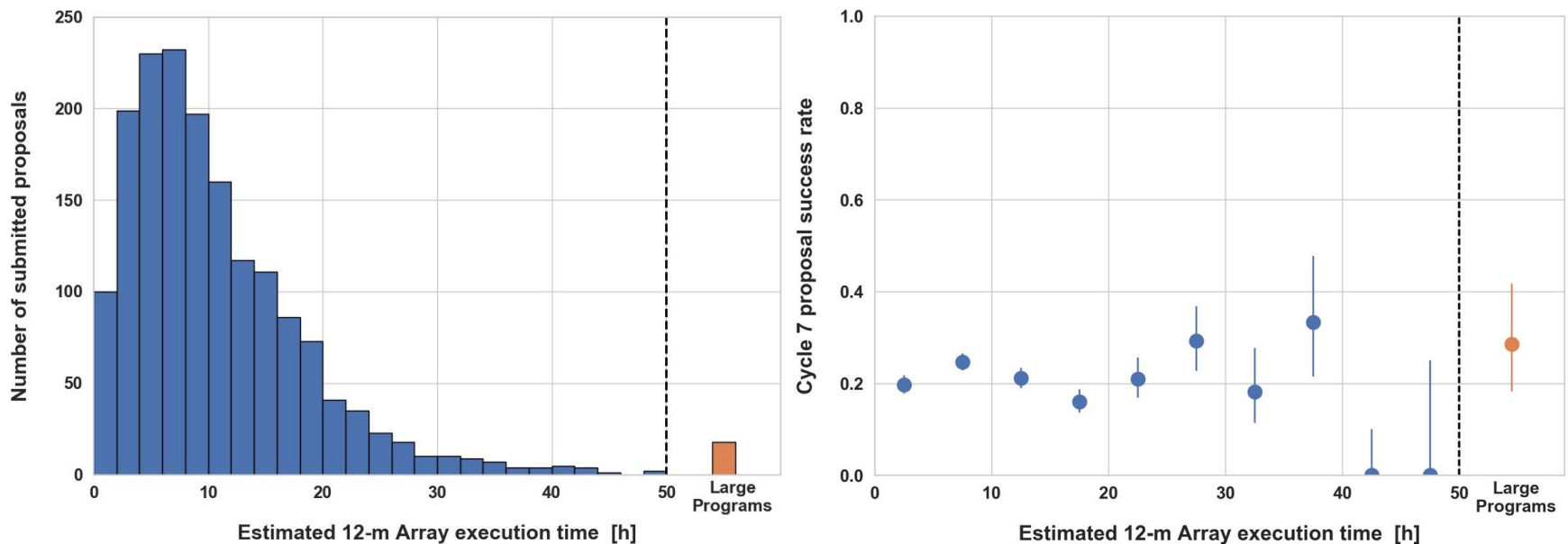


Figure 1: (Left) Number of proposals submitted as a function of the 12-m Array execution time in Cycle 7. (Right) The fraction of proposals (with 1sigma confidence intervals) that are assigned priority Grade A or B as a function of the estimated 12-m Array time.

ALMA Capabilities

ACA Supplemental Call:

- In Cycle 8, ALMA will offer a stand-alone ACA Supplemental Call for Proposals.
- The Supplemental Call will open on 15 September 2020 and the proposal deadline will be on 8 October 2020 (dates may change!).
- Observations from the Supplemental Call will be scheduled from January 2021 to September 2021.
- **The anticipated amount of time available will be announced in the Call. While stand-alone ACA proposals accepted from the Main Call may be assigned priority "A", "B", or "C", all accepted proposals from the Supplemental Call will be assigned priority "C".**
- Proposals submitted to the Supplemental Call will be peer reviewed through a distributed system, in which the PI of each submitted proposal, or a designee from among the co-Is, will be responsible for reviewing ten other proposals submitted in the same call.
 - Summary from the Cycle 7 Supplemental Call process can be found at:
<https://almascience.nrao.edu/news/results-of-the-aca-standalone-cycle-7-supplemental-call>
- More information about the supplemental call can be found at:
<https://almascience.nrao.edu/proposing/7m-array-supplemental-call>

ALMA Capabilities – NEW!!!

Dual-Anonymous Proposal Review

- Proposals in Cycle 8 will implement a dual-anonymous process for proposal reviews. While proposers will still enter their names and affiliations in the Observing Tool, their identities will be concealed from the reviewers.
- **It will be the responsibility of the investigators to write their proposals such that anonymity is preserved.**
- Guidelines on how to prepare such proposals is available now in an ALMA Science Portal news item and, later, in the CfP - <https://almascience.nrao.edu/news/items-for-planning-cycle-8-proposals>

ALMA Capabilities

Standard vs Non-Standard modes???

GONE!

- Unlike in previous cycles, there will no longer be a distinction between standard and non-standard modes so... there is no more 20% cap on the time request for non-standard modes!!!
- Proposal types in Cycle 8 will include Regular, Very Long Baseline Interferometry (VLBI), Target of Opportunity, and Large Program. VLBI proposals work in concert with the Global mm-VLBI Array (GMVA) or the Event Horizon Telescope (EHT).
- **GMVA programs must also submit a proposal to the GMVA by its 1 February 2020 deadline.** Additional information about proposing with ALMA using the GMVA was made available in the GMVA Call for Proposals in early January 2020.

However, Large Program Observing Modes will STILL be restricted. They cannot include:

- Polarization observations
- Bandwidth switching projects (having less than 1 GHz aggregate bandwidths over all spectral windows)
- Solar observations
- VLBI observations
- Non-standard calibrations (user-defined calibrations selected in the OT)
- Astrometric Observations

ALMA Capabilities

Full ALMA Operations (All Cycle 8 Capabilities plus):

- **Receiver bands:**
 - Include Bands 1 and 2
 - Band 1 summary report from 2019 June - <https://zenodo.org/record/3240351>. Full ALMA Band 1 Science Case: <http://arxiv.org/abs/1310.1604>
 - Band 2 summary report from 2019 June - <https://zenodo.org/record/3240407>
- **Baselines:**
 - All observing bands out to 16 km. Some may never be considered a standard mode.
- **Observing Time:**
 - Up to 4500 hours+ for successful proposals of PI programs expected on the 12m Array (includes DDT, Cycle 7+ Carryover and resubmissions)
- **Observing Modes:**
 - Full operations include full Stokes plus circular polarization at all observing bands including mosaics and Total Power.

ALMA Timelines and Milestones

The ALMA Cycle 8 Timeline

Date	Milestone
17 March 2020 (15:00 UT)	Release of Cycle 8 Call for Proposals, Observing Tool, and supporting documents and opening of the Archive for proposal submission
NO EARLIER than 1500 UT on 19 May, 2020	Proposal submission deadline for Cycle 8 proposals
End of July 2020	Announcement of the outcome of the proposal review process
9 September 2020	Deadline for submission of Phase 2 material for Cycle 8 accepted proposals
October 2020	Start of ALMA Cycle 8 science observations
September 2021	End of ALMA Cycle 8

ALMA Timelines and Milestones

The ALMA Cycle 8 Supplemental Call Timeline

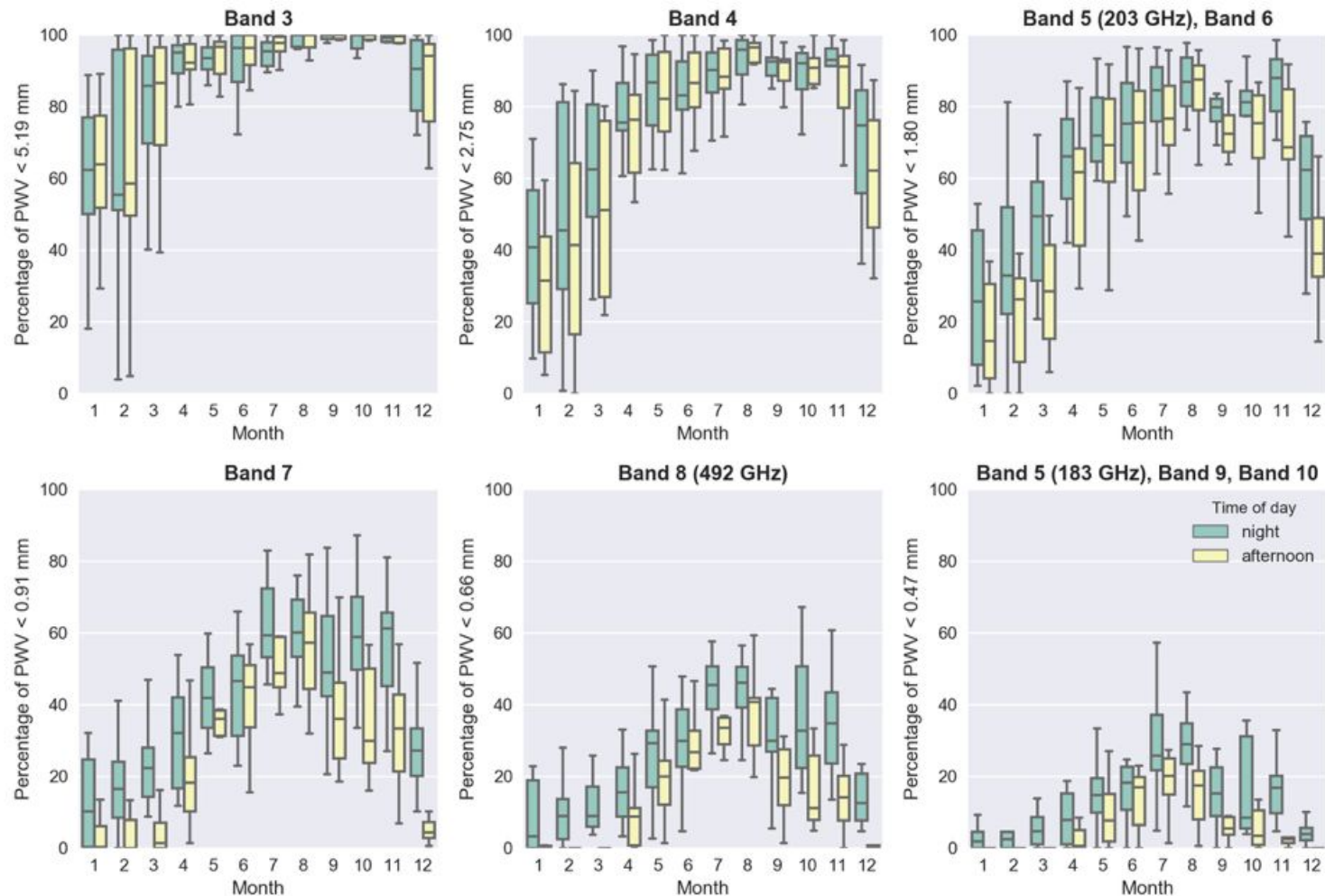
Date	Milestone
15 September 2020 (15:00 UT)	Call for Proposals and Supplemental Call submission server opened
08 October 2020 (15:00 UT)	Deadline to submit Supplemental Call proposals
22 October 2020	Proposals released to reviewers
29 October 2020 (15:00 UT)	Deadline for reviewer to report conflicts of interest on proposal review assignments
19 November 2020 (15:00 UT)	Deadline to submit reviews and ranks
Early December 2020	Notification emails sent to PIs
January 2021	Successful Supplemental Call proposals enter the observing queue

ALMA Array Configuration Schedule (Cycle 8)

- Antenna configurations for the main 12-m array will use a new nomenclature in Cycle 8.
 - Configurations will be called C-1, C-2, and so on up to C-10, with C-1 having similar characteristics to the C43-1 configuration of Cycle 7, and likewise for the others.
 - Cycle 8 will NOT include the two longest baseline 12-m array configurations, C-9 and C-10.
- Maximum baselines in Cycle 8 will therefore be 8.5 km in configuration C-8.
- Configurations C-9 and C-10 with maximum baselines of 13.9 km and 16.2 km, respectively, will again be available in Cycle 9.
- **NOTE: No PI observing takes place in Feb!**
- The forward-looking configuration schedule (through Cycle 9) can be found at:
<https://almascience.nrao.edu/observing/observing-configuration-schedule/long-term-configuration-schedule>

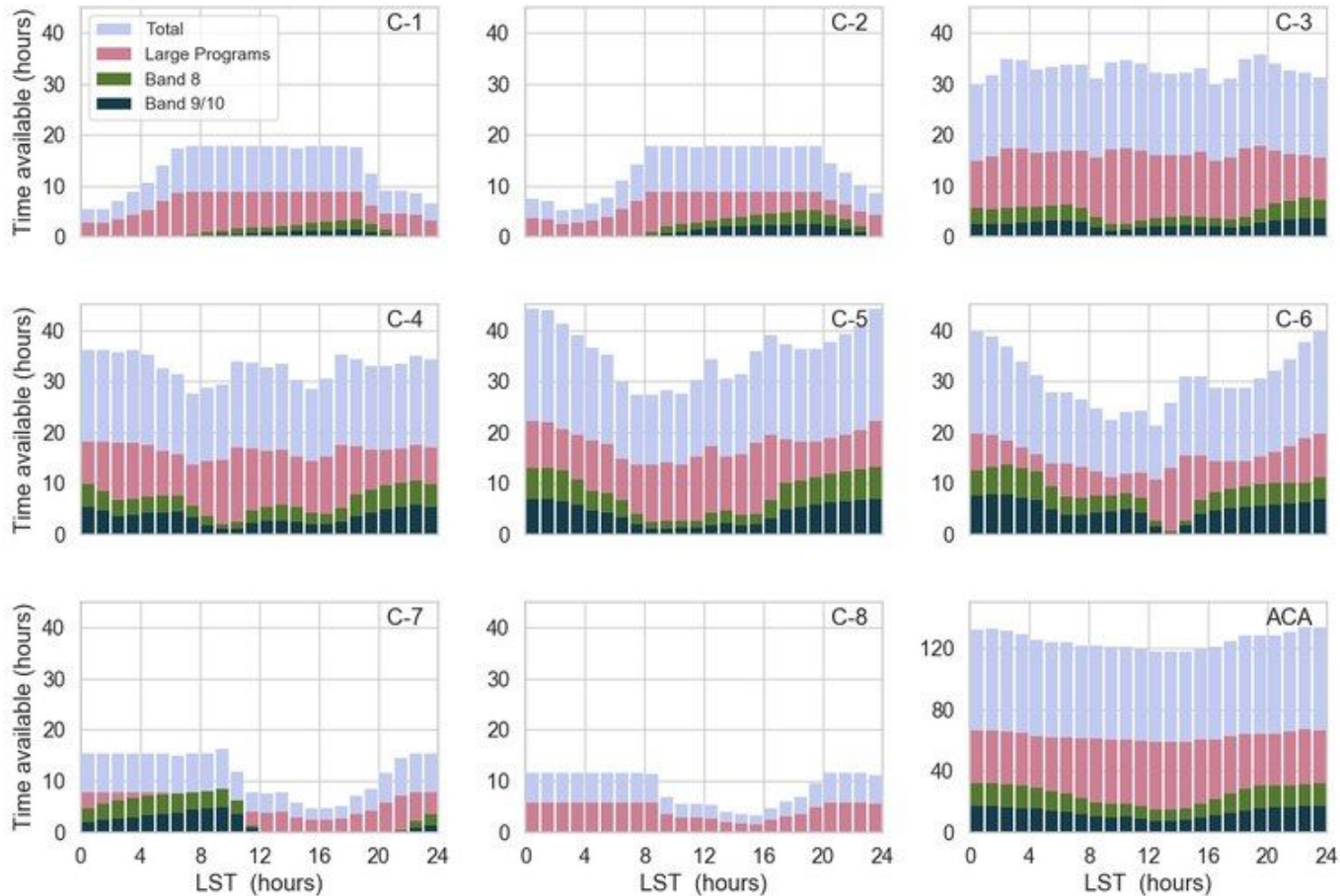
Start date	Configuration	Longest baseline	LST for best observing conditions
2020 October 1	C-8	8.5 km	~ 22h – 10h
2020 October 20	C-7	3.6 km	~ 23h – 11h
2020 November 10	C-6	2.5 km	~ 1h – 13h
2020 December 01	C-5	1.4 km	~ 2h – 14h
2020 December 20	C-4	0.78 km	~ 4h – 15h
2021 January 10	C-3	0.50 km	~ 5h – 17h
2021 February 1-28	No observations due to February Maintenance		
2021 March 1	C-1	0.16 km	~ 8h – 21h
2021 March 26	C-2	0.31 km	~ 9h – 23h
2021 April 20	C-3	0.50 km	~ 11h – 1h
2021 May 10	C-4	0.78 km	~ 13h – 3h
2021 May 31	C-5	1.4 km	~ 15h – 5h
2021 June 23	C-6	2.5 km	~ 16h – 6h
2021 July 28	C-5	1.4 km	~ 17h – 7h
2021 August 18	C-4	0.78 km	~ 19h – 8h
2021 September 10	C-3	0.5 km	~ 20h – 9h

ALMA Observing Strategies (Cycle 8)



- Box and whisker plots of the percentage of time that the precipitable water vapor (PWV) is less than the thresholds adopted for the various ALMA bands versus the month of the year. Results are shown for both night time (green) and mid-afternoon (yellow), and assume a source elevation of 60 degrees. The horizontal line within a box indicates the median, the boundaries of a box indicate the 25th- and 75th-percentile of the distribution, and the whiskers indicate the highest and lowest values of the distribution. The PWV measurements were obtained by the APEX weather stations between 2007 and 2017.

ALMA Observing Strategies (Cycle 8)



- Effective observing time available per configuration for executing PI projects. As an example, up to 36 hours are expected to be available in C-4 at LST=00 h for all observations and up to 18 h may be allocated to Large Programs. The total number of hours excludes time spent on observatory calibration, maintenance, reconfigurations, and other activities. The time available for Large Programs is shown in pink and time for high-frequency observations in green and dark blue. The configuration schedule and, consequently, the total number of hours available per configuration may change in response to proposal pressure. The data files containing these histograms are available [here](#).

NAASC Sources of Support

- **ALMA Helpdesk:** User support is a priority so questions are usually answered within 48 hours (with around the clock staffing in the week leading up to the proposal deadline) - <https://help.almascience.org>
- **Student Observing Support:** Successful ALMA proposals will be invited to apply for up to \$35k to support undergraduate or graduate student involvement - <https://science.nrao.edu/opportunities/student-programs/sos>
- **NAASC Financial Support for Workshop/Conferences:** The NAASC invites scientists to apply for funding in support of upcoming conferences and workshops. - <https://science.nrao.edu/facilities/alma/community/NAASC-Conference-and-Workshop-Support>
- **Page Charges:** Upon request NRAO covers page charges for authors at US institutions when reporting results from ALMA/VLA - <https://library.nrao.edu/pubsup.shtml>
- **Face-to-face Visitor Support:** Upon request NRAO will cover the travel expenses of up to 2 people from 2 teams per week to come to the NAASC to get support for data reduction, proposal preparation, etc... We also have long term visitor support as well - <https://science.nrao.edu/facilities/alma/visitors-shortterm>
- **ALMA Ambassadors:** You too can become an ALMA Ambassador. For program eligibility visit - <https://science.nrao.edu/facilities/alma/ambassadors-program>



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