

The ALMA Proposal Preparation Process

How to get started and what to expect



[Your Name as Presenter here]



Atacama Large Millimeter/submillimeter Array
Expanded Very Large Array



This talk is for you if...

- You are new to ALMA and have not yet had experience with the relevant documentation...
- You have not downloaded the ALMA Observing Tool (OT) or even know where to get it.
- You have a fabulous science case that will be essential to follow-up with ALMA facilities...
- You would like examples of science use cases for ALMA
- You were familiar with past Cycles and wonder what Cycle 5 capabilities are now available and what changes will be made before the Call for Proposals.

This talk will be available online for reference after this workshop.

Proposal Checklist

- Read relevant documentation (CfP Guide, Primer, etc.)
- Create an ALMA account by registering at the Science Portal (almascience.org)
- Download the Observing Tool (OT) & related guides
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 - New capabilities for Cycle 5!
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Cycle 5 Documentation & Timeline

- Call for Proposals
- Proposer's Guide
- ALMA Primer
- OT Guide
- ALMA Tech Handbook
- Timeline for Cycle 5
 - Mar. 21 – Call for Proposals
 - Apr. 20 – Proposal Deadline
 - Aug 2017 – Results to Pis
 - Sept 2017 – PIs submit SBs
 - Oct. 2017 – Start of Cycle 5
 - Duration – 11-12 months




Observing with *ALMA*
A Primer for Early Science



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ALMA



Atacama Large Millimeter/submillimeter Array

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More news...

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Jun 09, 2017

More...

Status

ALMA Cycle 5 Pre-Announcement

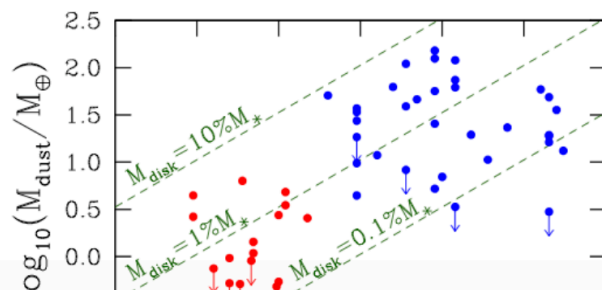
Refereed publications:

Last observed source:

Current configuration: C40-2

More...

Science Highlights - Possible Disk Truncation in Ophiuchus Brown Dwarfs



The sensitivity, resolution and the wavelength coverage of ALMA makes it an ideal tool for studying the properties of the cold outer disks of young stars and low mass objects. Such observations can aid us in understanding the formation of their central objects and their likelihood of ultimately hosting planets. In a recent Astronomy & Astrophysics [paper](#), Dr. Testi and his collaborators made use of ALMA Band 7 to observe an unbiased sample of spectroscopically confirmed Ophiuchus brown dwarfs with infrared excesses.



www.almascience.org

ALMA Science Portal @ NRAO

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Downloading the ALMA OT



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Observing Tool

The ALMA Observing Tool (OT) is a Java application used for the preparation and submission of ALMA Phase 1 (observing proposal) and Phase 2 (telescope runfiles for accepted proposals) materials. It is also used for preparing and submitting Director's Discretionary Time (DDT) proposals. The current [Cycle 4](#) release of the OT is configured for the present capabilities of ALMA as described in the [Cycle 4 Call For Proposals](#). Note that in order to submit proposals you will have to register with the ALMA Science Portal beforehand.

Note that preparation of Cycle 3 Phase 1 and DDT proposals needs to be done using the Cycle 3 version of the Observing Tool. This version of the OT can be found in the [DDT page](#), or the Phase 2 menu.

Download & Installation

The OT will run on most common operating systems, as long as you have **Java 8** installed ([see the troubleshooting page if you](#) are experiencing Java problems). The ALMA OT is available in two flavours: Web Start and tarball.

The **Web Start** application is the recommended way of using the OT. It has the advantage that the OT is automatically downloaded and installed on your computer and it will also automatically detect and install updates. There are some issues with Web Start, particularly that it does not work with the Open JDK versions of Java such as the "Iced Tea" flavour common on many modern Linux installations. The Sun/Oracle variant of Java should therefore be installed instead. If this is not possible, then the tarball installation of the OT is available.

The **tarball** version must be installed manually and will not automatically update itself, however there should be no installation issues. For Linux users, we also provide a download complete with a recommended version of the Java Runtime Environment. Please use this if you have any problems running the OT tarball install with your default Java.

Webstart

Tarball



OT Video Tutorials



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OT Video Tutorials

The OT video tutorials provide an audio-visual demonstration of different aspects of proposal preparation in the OT. Novice users should start with the first video and work their way down, while more experienced users may want to jump straight to one of the specialised videos.

OT Video Tutorial 1: Useful to Know

This video will help you get started with the OT and introduce you to some handy tips and tricks. Topics covered include navigating the OT, using the help function, the template library, time estimation, validation, opening & submitting projects including re-submissions, and the concept of non-standard modes.



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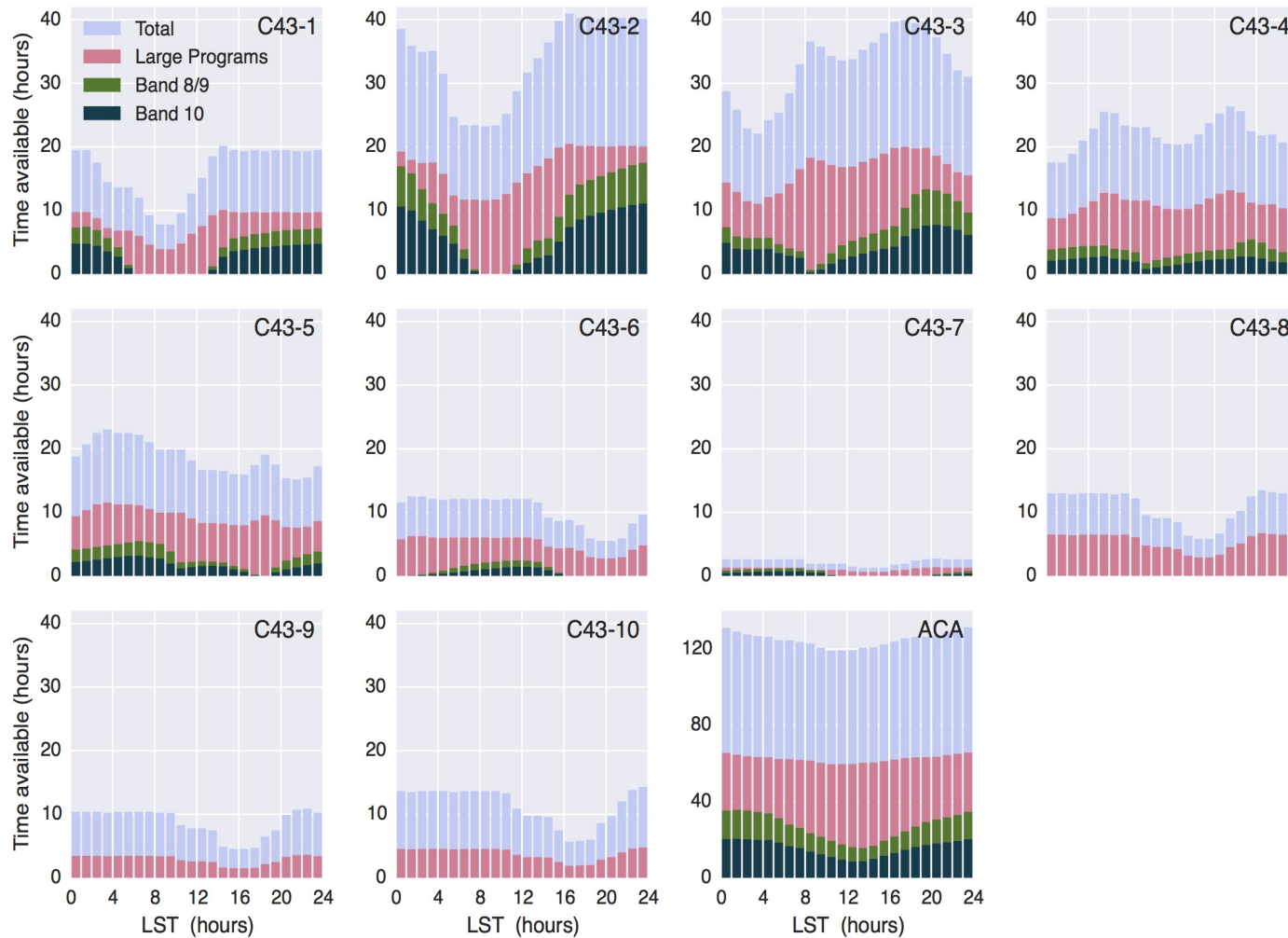
Science Case

- Must include:
 - Astronomical Importance
 - Estimated intensity, S/N
- May include:
 - Figures
 - Tables
 - References
- Free-form PDF document
 - 12+ font, English only
 - 20 MB file size
 - 4 pages (6 for Large Projects)

Table 1: Cycle 5 Configuration Schedule

Start date	Configuration	Longest baseline	LST for best observing conditions
2017 October 1	C43-7	3.7 km	~ 21h - 10h
2017 October 5	C43-8	6.8 km	~ 22h - 11h
2017 October 25	C43-9	12.8 km	~ 23h - 12h
2017 November 10	C43-10	16.5 km	~ 1h - 13h
2017 December 1-18	No observations due to large antenna reconfiguration		
2017 December 19	C43-6	1.8 km	~ 4h - 15h
2018 January 10	C43-5	1.1 km	~ 5h - 17h
2018 February 1-28	No observations due to February shutdown		
2018 March 1	C43-4	0.7 km	~ 8h - 21h
2018 March 30	C43-3	0.46 km	~ 10h - 0h
2018 May 15	C43-2	0.27 km	~ 12h - 3h
2018 June 15	C43-1	0.15 km	~ 14h - 5h
2018 July 15	C43-2	0.27 km	~ 17h - 7h
2018 August 15	C43-3	0.46 km	~ 18h - 8h
2018 August 30	C43-4	0.7 km	~ 19h - 9h
2018 September 15	C43-5	1.1 km	~ 20h - 10h

ALMA Observing Strategies (Cycle 5)



- Histograms of the anticipated amount of observing time available versus LST for the antenna configurations in Cycle 5. Also shown are histograms of the time available for Large Programs, as well as high frequency observations (Bands 8, 9, and 10) based on historical PWV data

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New Capabilities

- Band 5
- Improved spectral scan mode
- 90 Degree Walsh switching at Band 9
- Solar Observations (Bands 3 + 6)
- VLBI full polarization continuum (Bands 3, 6 + 7)

Large Projects (started in Cycle 4)

- Any project >50 hours
- Standard observing modes
- Automatic 'A' grade
- +2 pages for Science Case
 - Data/Project Mgmt. Plans
 - Enhanced Data Products

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Project Structure

Proposal Program

Unsubmitted Proposal

- Project
 - Proposal
 - Planned Observing
 - ScienceGoal (Copy of B6 12CO (2-1): NGC3256)
 - General
 - Field Setup
 - Spectral Setup
 - Calibration Setup
 - Control and Performance
 - Technical Justification

Template library. Turn the keys on the JTree below & r...

- Template library. Turn the keys on the JTree below & r...
 - Proposal
 - Planned Observing
 - ScienceGoal (B3 spectral sweep CO)
 - ScienceGoal (B7 continuum: COSMO)
 - ScienceGoal (B7 CO(9-8): Cosmic Ey
 - ScienceGoal (B9 continuum: Cosmic
 - ScienceGoal (B3 spectral sweep: PK)
 - ScienceGoal (B3 continuum: GRB To
 - ScienceGoal (B6 continuum: GRB To
 - ScienceGoal (B7 continuum: GRB To
 - ScienceGoal (B6 continuum: GRB To
 - ScienceGoal (B3 continuum: GRB To
 - ScienceGoal (B6 12CO (2-1): NGC32
 - ScienceGoal (B6 13CO (2-1): NGC32
 - ScienceGoal (B6 spectral line: Massi
 - ScienceGoal (B9 spectral line: Massi
 - ScienceGoal (B3 continuum: Protoste
 - ScienceGoal (B6 continuum: Protoste

Editors

Spectral Spatial Field Setup

Source Name NGC3256 Resol

Choose a Solar System Object? ☐

Name of object Unspecified

System J2000

Sexagesimal

display? ☒

Parallax 0.00000

mas

Source Coordinates

RA 10:27:51.6000

PM RA 0.00000

mas/yr

Dec -43:54:18.000

PM DEC 0.00000

mas/yr

Source Radial Velocity

0.000

km/s

hel

z

0.000000000

Doppler Type

RELATIVISTIC

Target Type

☐ Individual Pointing(s) ☒ 1 Rectangular Field

Expected Source Properties

Peak Continuum Flux Density per Beam 0.17400

Jy

Continuum Polarization Percentage

0.0

%

Peak Line Flux Density per Beam

0.00000

Jy

Line Width

0.00000

km/s

Line Polarization Percentage

0.0

%

Rectangle

Coords Type ☐ ABSOLUTE ☒ RELATIVEField Center
Coordinates

System

J2000

Offset(Longitude)

0.00000

arcsec

Offset(Latitude)

0.00000

arcsec

Feedback

Validation

Validation History

Log

Description

Suggestion

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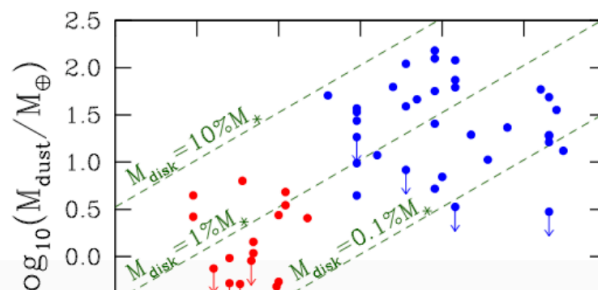
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www.almascience.org

ALMA Science Portal @ NRAO

I could use a hand...

Have no fear, the ALMA Helpdesk is here...



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General ALMA Queries (13)

Early Science - Cycle 1 (31)

Resources & Observer Support (12)

Project Planning (14)

ALMA Observing Tool (OT) (29)

Proposal Handling (5)

Archive & Data Retrieval (4)

Offline Data Reduction and/or CASA (14)

Development Program (1)

Please type your search query here

SEARCH

Knowledgebase

General ALMA Queries (13)

Can I submit a ticket in Japanese?

How close can ALMA observe to the Sun?

Early Science - Cycle 1 (31)

Can I use "breakpoints" in ALMA cycle 1?

The Cycle 1 Technical Handbook has some gaps in its discussion of ALMA receivers (SSB, 2SB, DSB). What else can you tell me about them?

Resources & Observer Support (12)

How do I arrange a visit to one of the ARCs?

Where can I find ALMA documentation and manuals?

Project Planning (14)

What should I include for the content of the Technical Justification and in what format should I submit it?

Where can I find the online ALMA observing simulator developed by the University of Manchester?

ALMA Observing Tool (OT) (29)

What do I do if I can't get the OT to work?

How do I deal with targets with unspecified coordinates in the OT?

Proposal Handling (5)

May I submit an identical proposal to more than one category, e.g. submitting a proposal on distant galaxies both to cosmology and to galaxy categories?

Which category should I submit a proposal on distant galaxies: "cosmology/high-z" or "Galaxies/Nuclei"?



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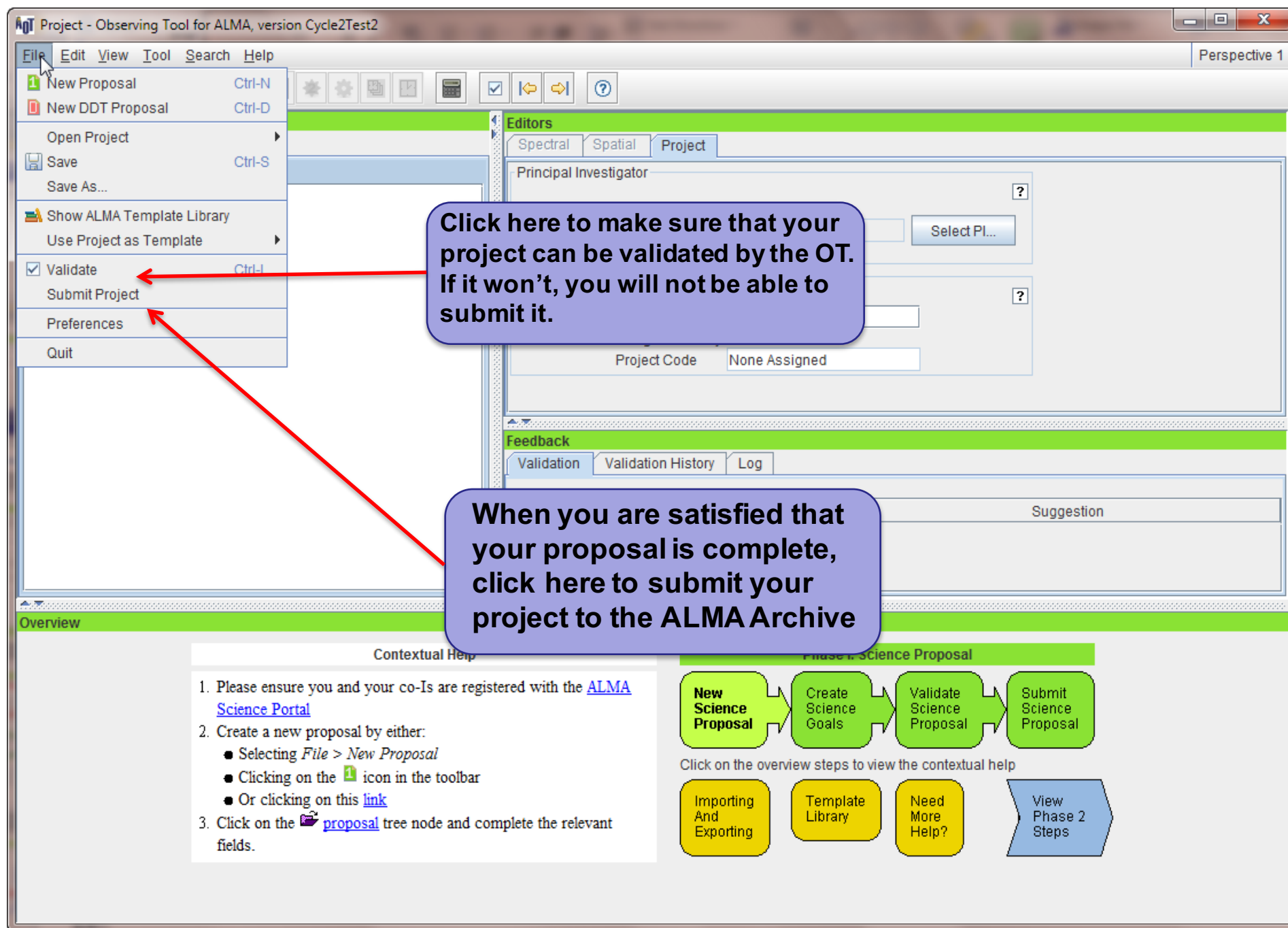
[General ALMA Queries \(14\)](#)[Early Science - Cycle 2](#)[Early Science - Cycle 1 \(31\)](#)[Resources & Observer Support \(12\)](#)[Project Planning \(14\)](#)[ALMA Observing Tool \(OT\) \(29\)](#)[Proposal Handling \(5\)](#)[Archive & Data Retrieval \(4\)](#)[Offline Data Reduction and/or CASA \(15\)](#)[Development Program \(1\)](#)[View Tickets](#)[Submit a Ticket](#)[Knowledgebase](#)[News](#)

Latest Updates

No information available in this view

help.almascience.org

ALMA Helpdesk @ NRAO (logged in view)



After submission

- Remember, you can resubmit as often as needed, but keep in mind that the server is quite busy right before the deadline
- Standard and ToO proposals will be reviewed by the ALMA Proposal Review Committee (APRC) and the ALMA Review Panels (ARP).
- All proposals will be subject to Technical Assessment by a selected group of JAO and ARC experts.
- Proposals will be assessed on the basis of the overall scientific merit of the proposed investigation and its potential contribution to the advancement of scientific knowledge.
- Following approval by the Directors Council, the outcome of the Proposal Review Process will be communicated to the PIs of all valid submitted proposals, expected in August 2017.

After submission

- Phase II (Creating and Queuing Scheduling Blocks)
 - PIs create their own scheduling blocks (with guidance from NAASC staff)
 - Being prompt helps ensure your project can be observed!
- Then wait – dynamic scheduling means your Contact Scientist doesn't know when your project will run. As observations are made, updates are shown in the SnooPI tool on the Science Portal:
<https://almascience.nrao.edu/observing/snoopi>



For more info:

<https://almascience.nrao.edu/>

The Atacama Large Millimeter/submillimeter Array (ALMA), an international astronomy facility, is a partnership of Europe, North America and East Asia in cooperation with the Republic of Chile. ALMA is funded in Europe by the European Organization for Astronomical Research in the Southern Hemisphere (ESO), in North America by the U.S. National Science Foundation (NSF) in cooperation with the National Research Council of Canada (NRC) and the National Science Council of Taiwan (NSC), and in East Asia by the National Institutes of Natural Sciences (NINS) of Japan in cooperation with the Academia Sinica (AS) in Taiwan. ALMA construction and operations are led on behalf of Europe by ESO, on behalf of North America by the National Radio Astronomy Observatory (NRAO), which is managed by Associated Universities, Inc. (AUI), and on behalf of East Asia by the National Astronomical Observatory of Japan (NAOJ). The Joint ALMA Observatory (JAO) provides the unified leadership and management of the construction and operation of ALMA.

