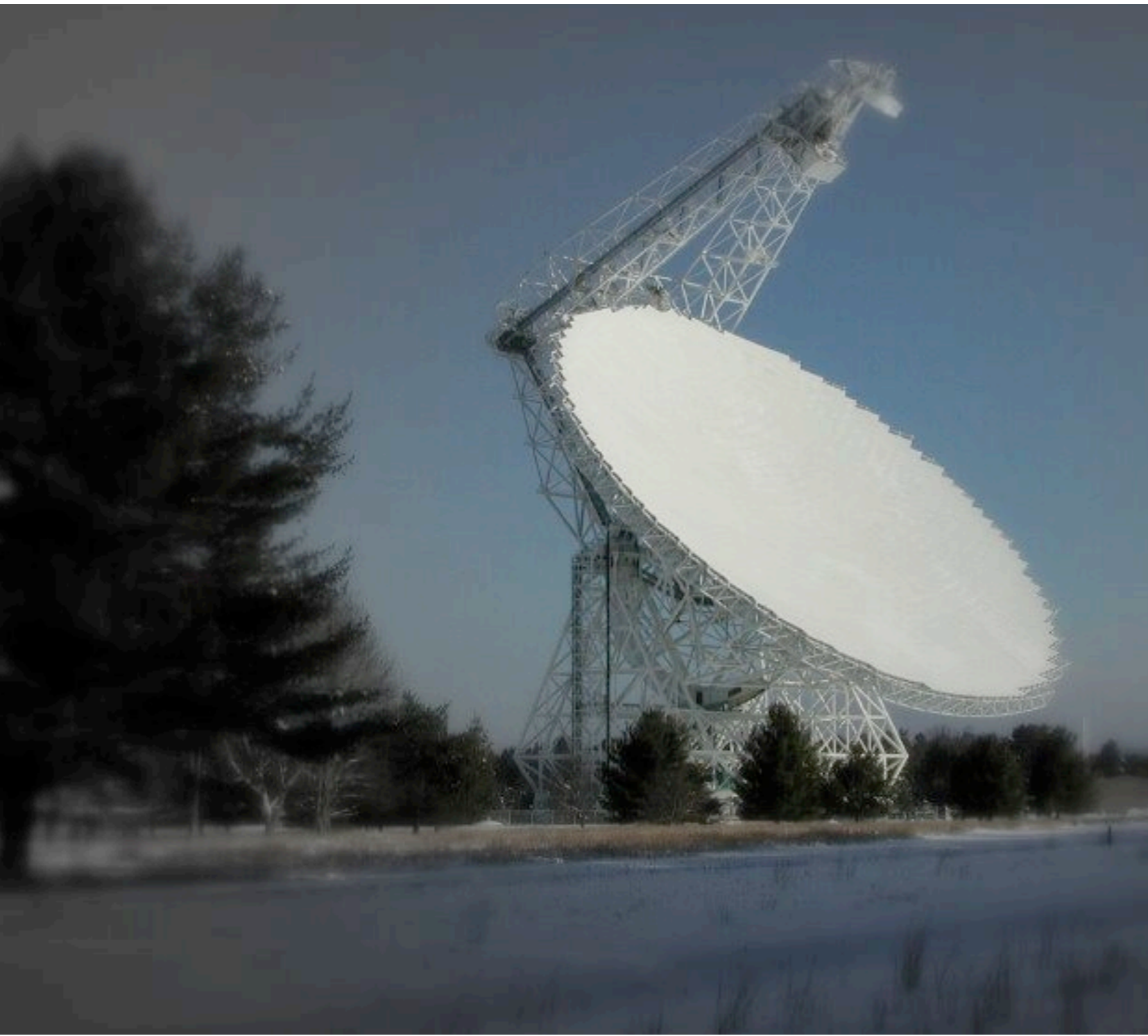




The Green Bank Telescope

a powerful instrument for enhancing ALMA science



Unblocked Aperture

Low sidelobes gives high dynamic range

Resistance to Interference

Excellent spectral Baselines

Excellent sensitivity to low surface brightness

Frequency coverage from 100 MHz–100 GHz

Spectroscopy, Continuum, Pulsar, VLBI

>85% Sky Coverage

Pointing to 1"–2" accuracy

Surface good for 3mm work

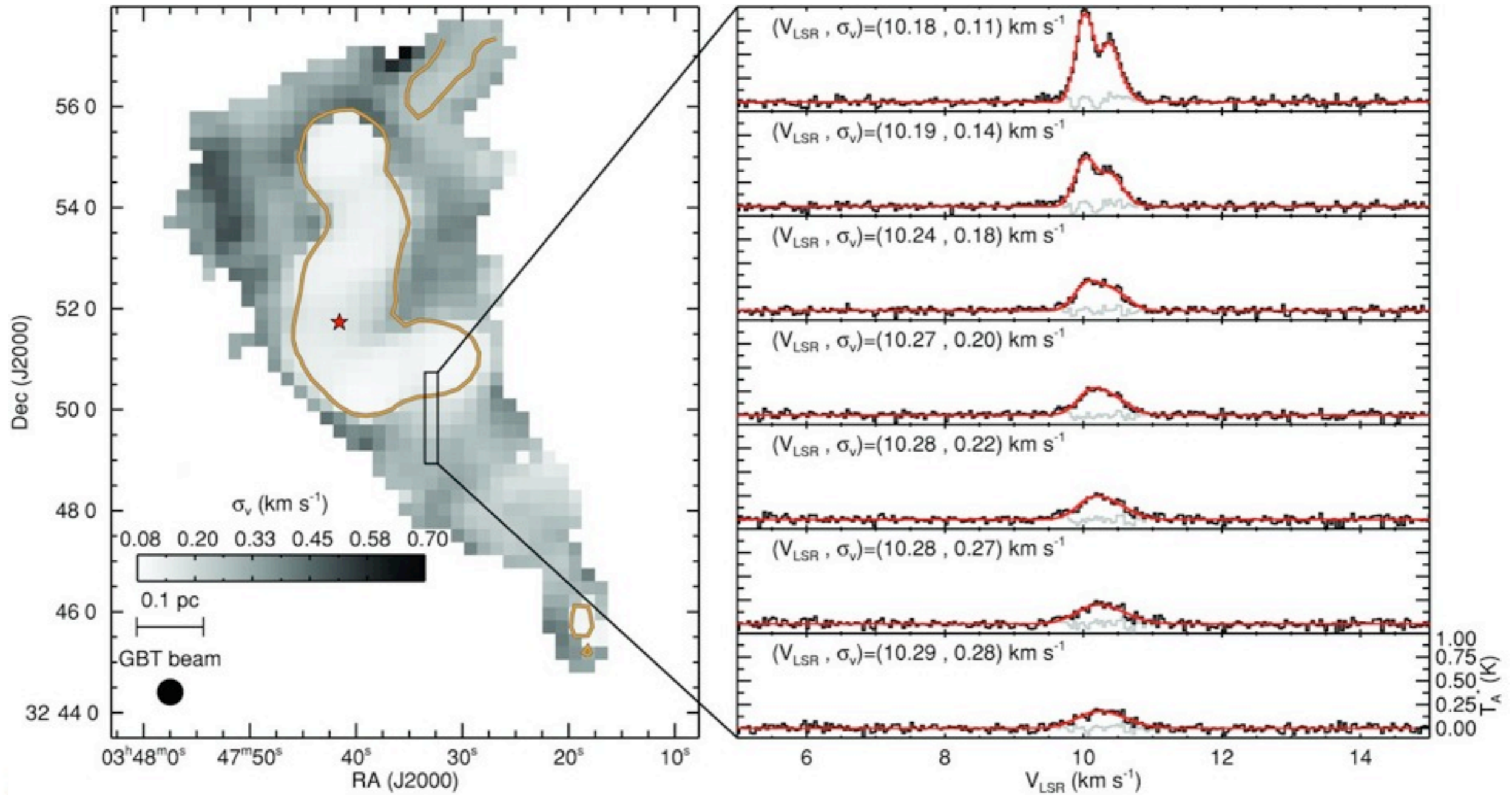
Active Instrument Development Program

Site Protected by a 13000 km² Radio Quiet Zone

GBT Studies of Star Formation

Direct Observation of a Sharp Transition to Coherence in Dense Cores

Pineda et al 2010, ApJ

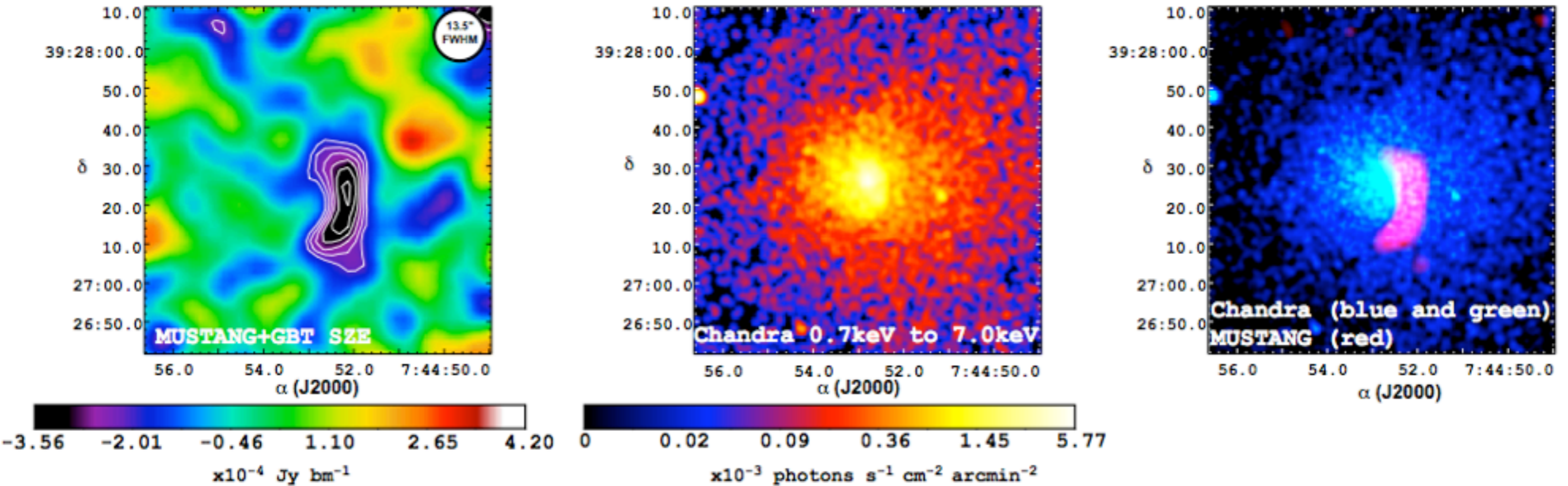


MACS0744+3927

previously unknown/unexpected weak shock near the core of this cluster

Korngut et al. (submitted)

MUSTANG
64-pixel 90 GHz bolometer array
now routinely imaging the SZE
at $\approx 10''$ resolution

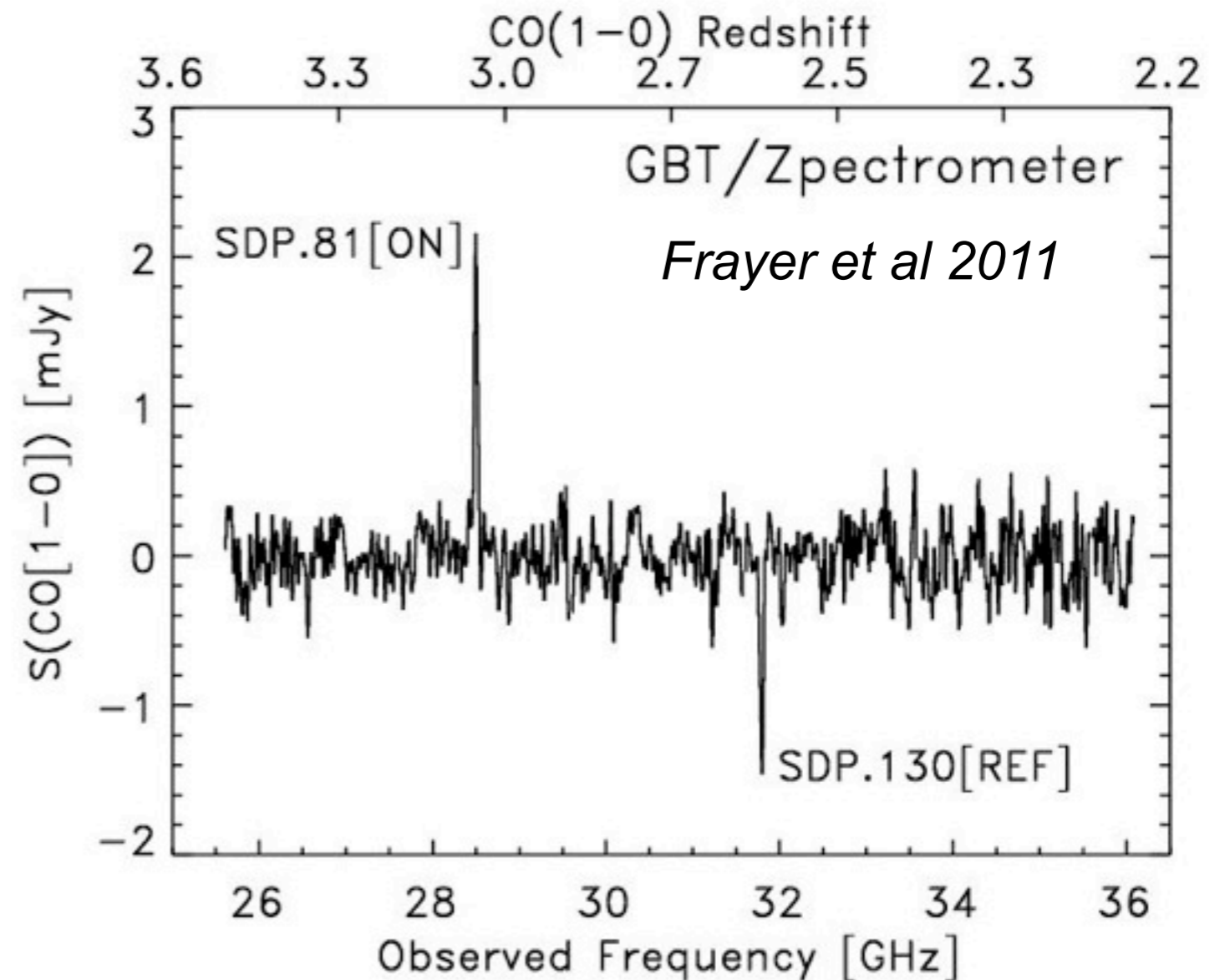


9'' (FWHM) on GBT; 40''x40'' FOV
Sensitivity: rms ~ 3 mK $\text{sec}^{1/2}$
– 3'x3' region to 0.4 mJy (RMS) in one hour
– integrates down for at least 12 hours

Blind High- z CO detections

GBT discovery of highly-redshifted CO(1-0) emission from optically obscured, lensed, Herschel sub-millimeter galaxies.

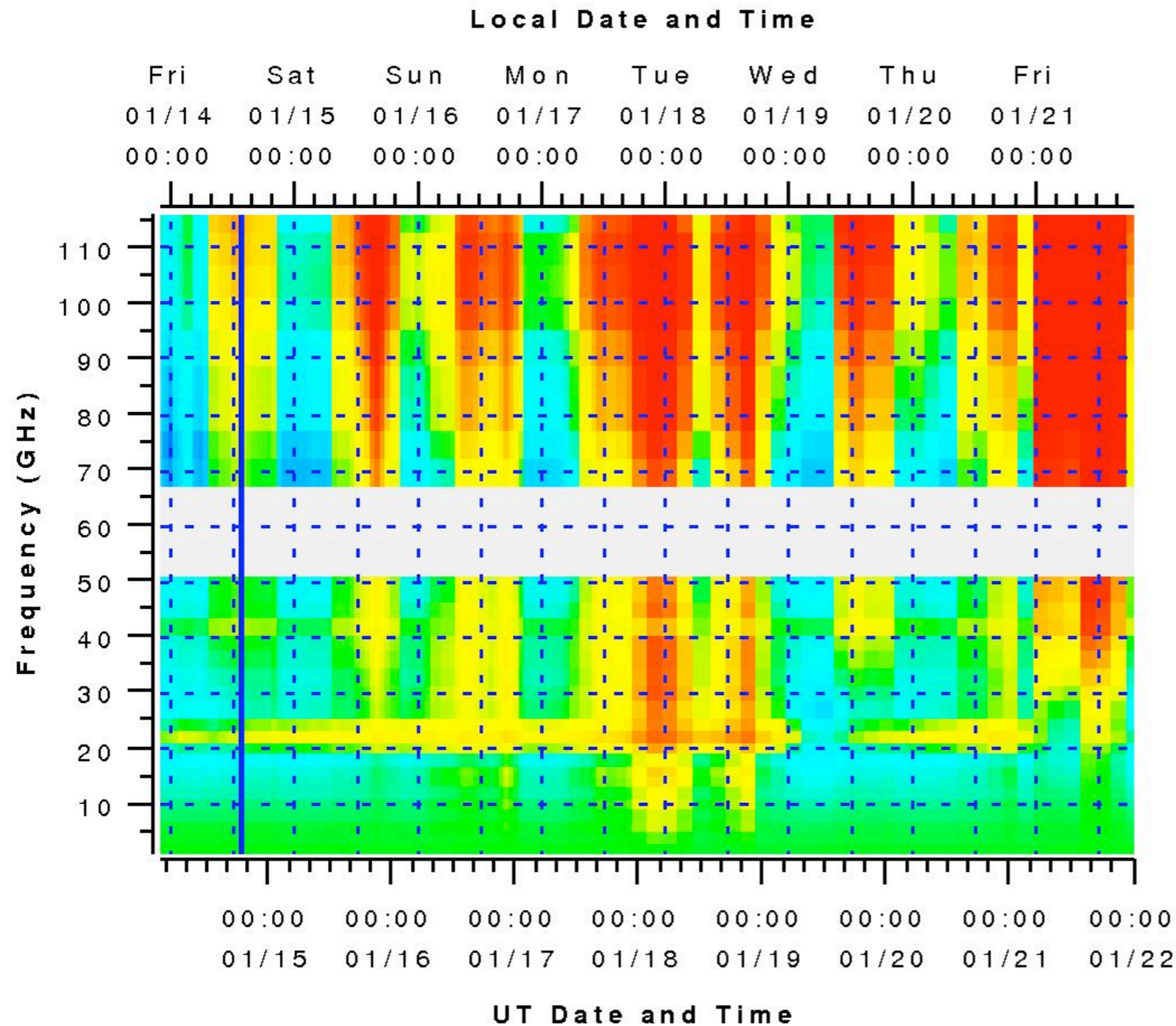
The line strength reveals the presence of large reservoirs of cold molecular gas in young systems.



Ronald J Maddalena
National Radio Astronomy Observatory
Green Bank, WV

**GBT Dynamic Scheduling
matches the project
to the weather**

Overview: DSS Relative Efficiencies without Limits (η/η_{mi})



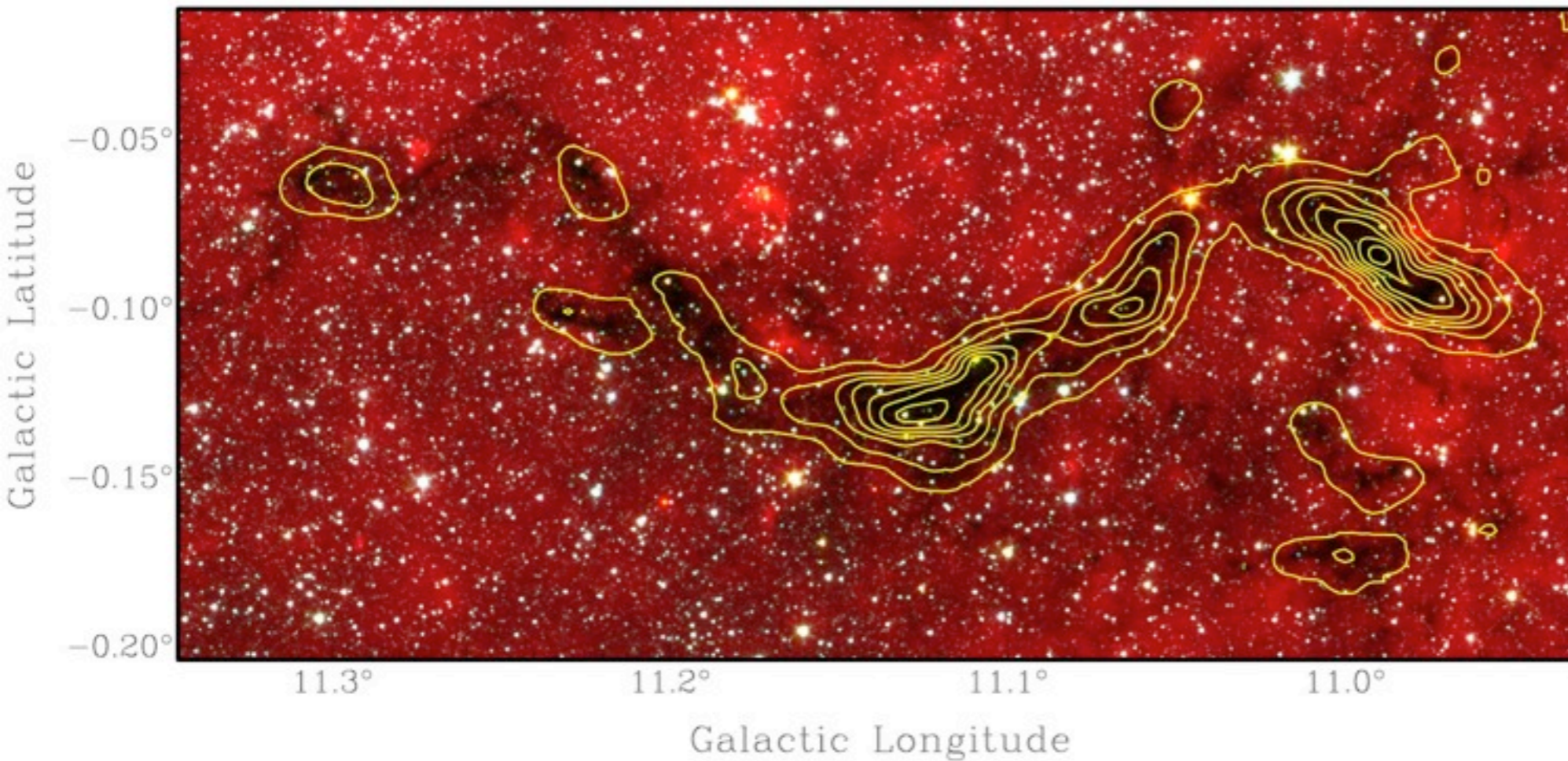
1776 hours of observing at >18 GHz dynamically scheduled in 2010 -- this amount should rise in the coming years



GBT Instrument Development Program

- K-band Focal Plane Array 7 pixels for 18–27 GHz
 - Completed and in regular use
- 4mm two-pixel receiver for 68–92 GHz spectroscopy
 - Under construction, commissioning late 2011, see <http://www.gb.nrao.edu/4mm>
- FPGA based spectrometer with up to 128 spectral windows
 - Under construction with CASPER group UCB, testing late 2011
- MUSTANG-2 bolometer array >100x faster than MUSTANG
 - Proposed but not yet funded
- W-band 100 pixel Focal Plane array 84–116 GHz (ALMA Band 3)
 - Under study but not yet funded

The GBT K Band Focal Plane Array is up and running



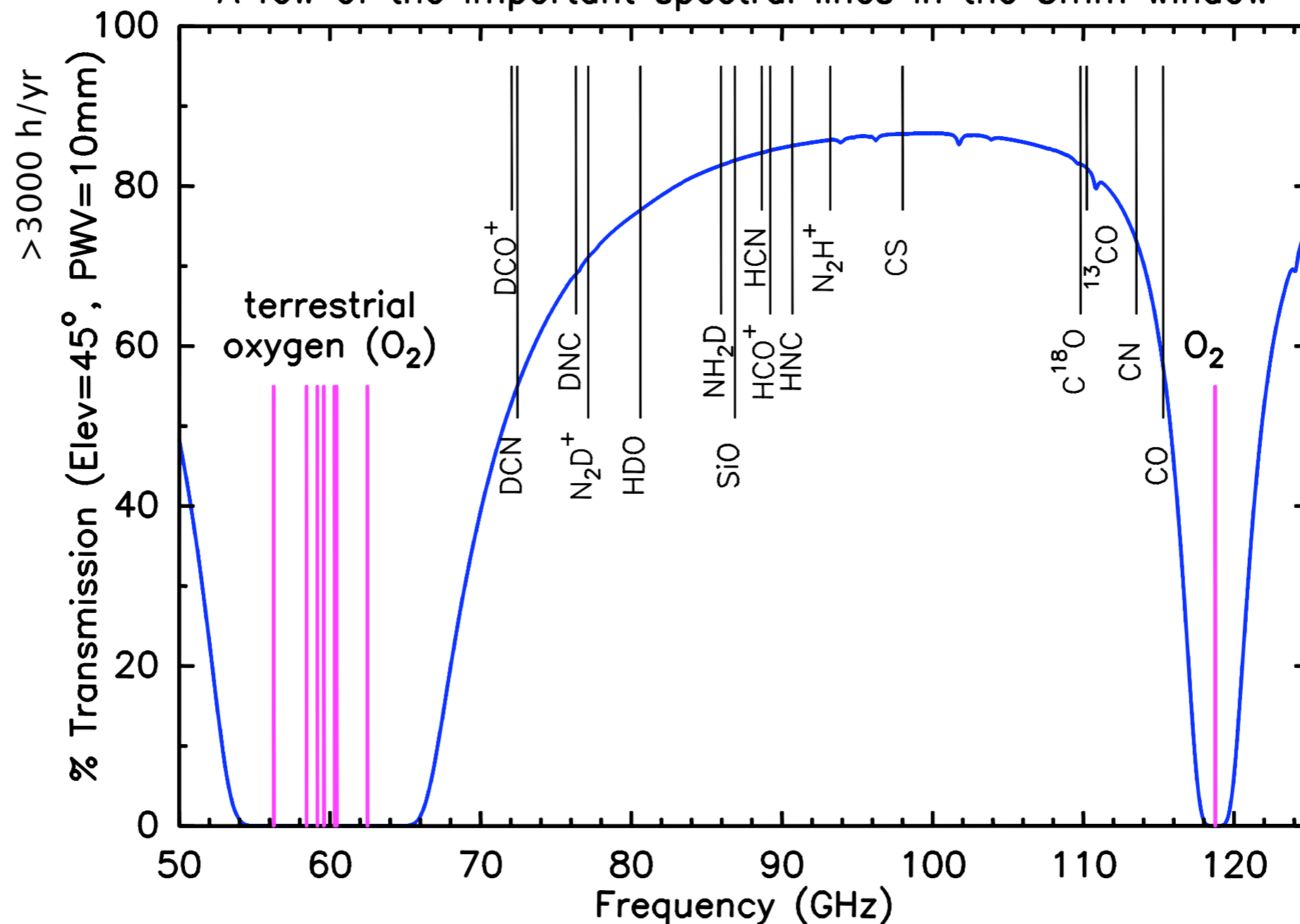
**KFPA
NH₃ (1,1)**

**Ammonia mapping of dark clouds
Finn & Jackson**

4mm Receiver Spectral Coverage

Available on the GBT in late 2011

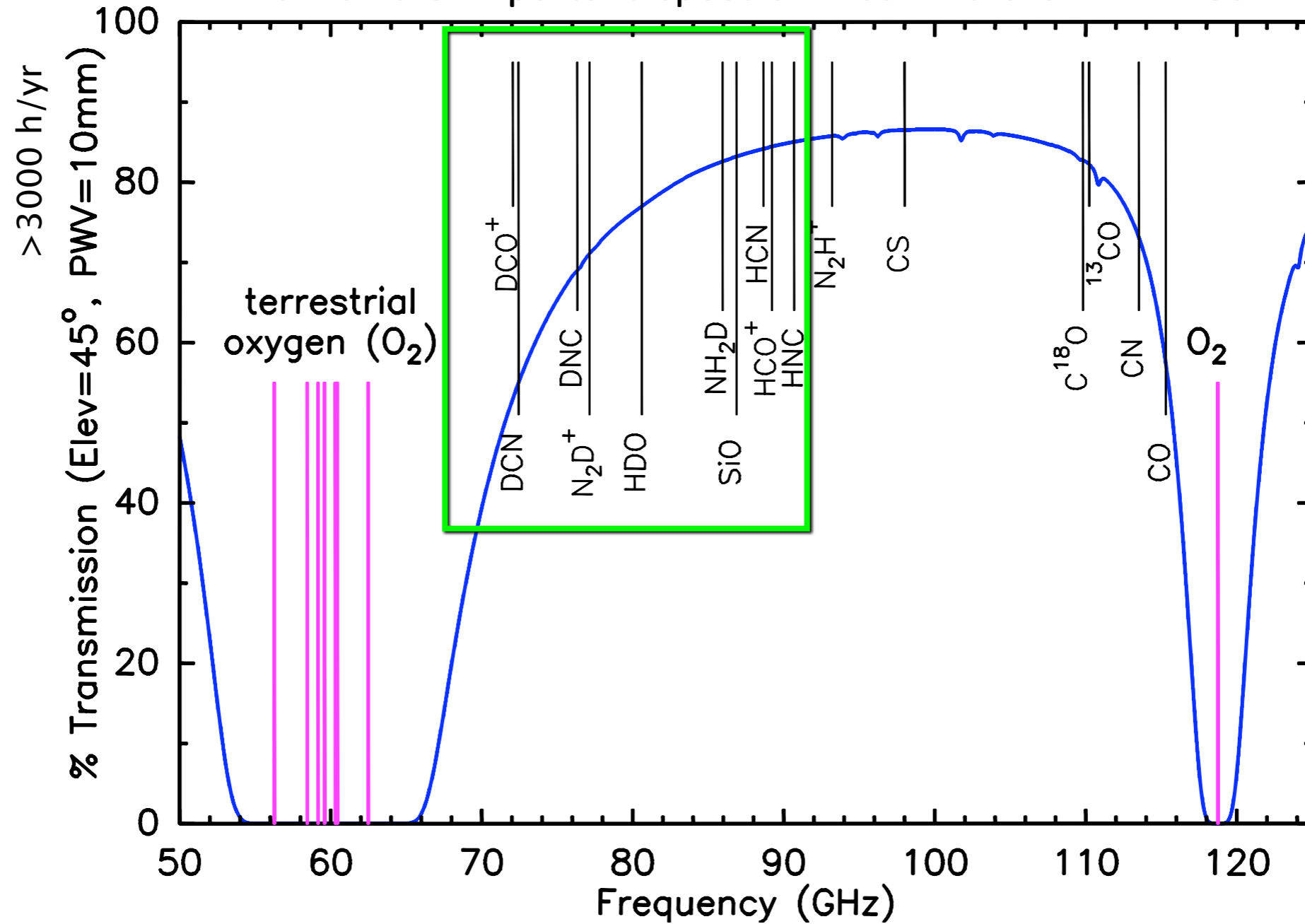
A few of the important spectral lines in the 3mm window



4mm Receiver Spectral Coverage

Available on the GBT in late 2011

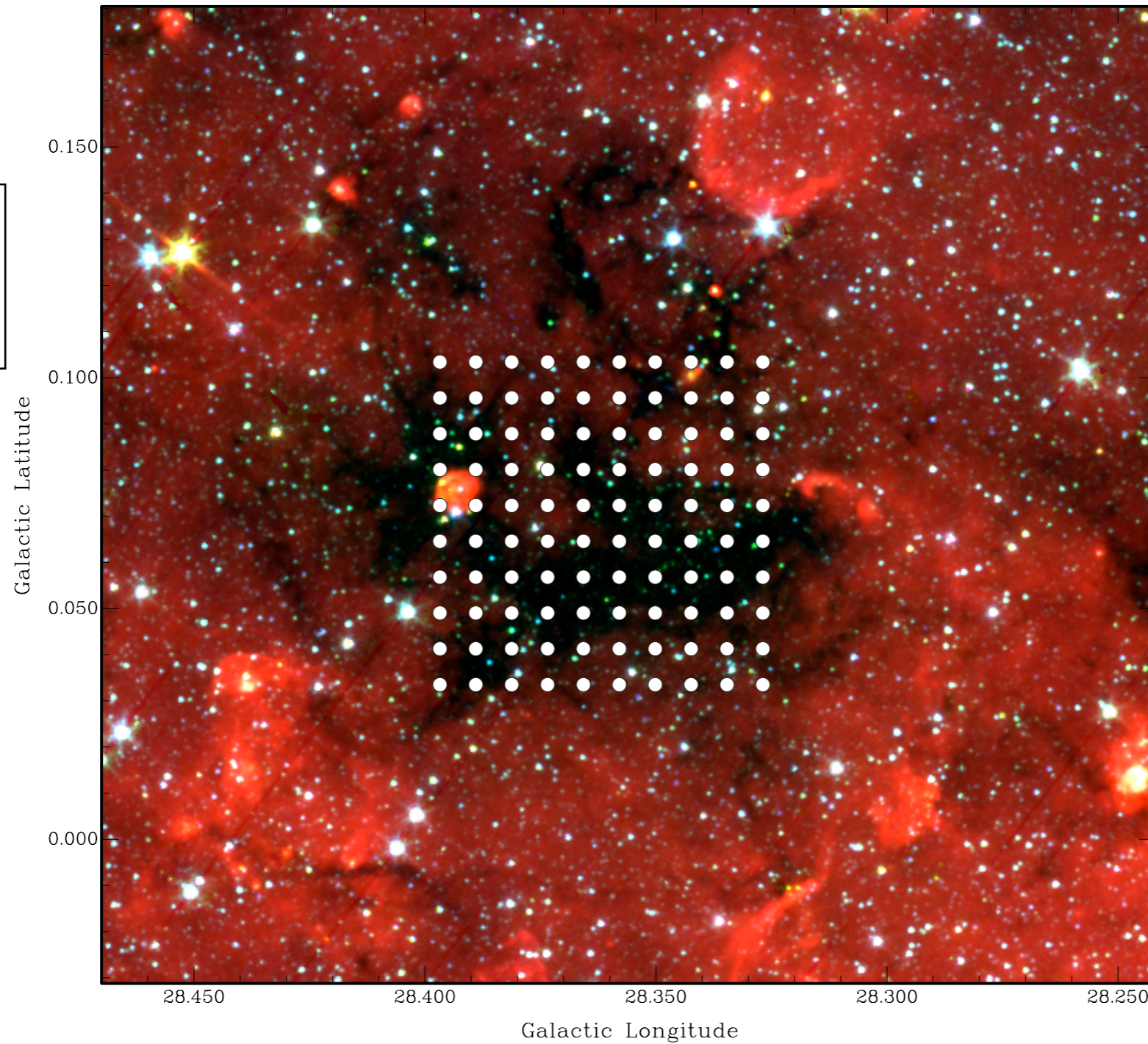
A few of the important spectral lines in the 3mm window



Planned 3mm Focal Plane Array

a wide field mapping complement to ALMA Band 3

GBT 3mm FPA
footprint on
an Infrared
Dark Cloud



ALMA primary
beam at 3mm

Field of View
Sensitivity
Resolution

