### Performance and Properties of the GBT



#### Richard Prestage for the PTCS Team

GBT High Frequency Science Workshop – 21 September 2015



Atacama Large Millimeter/submillimeter Array
Karl G. Jansky Very Large Array
Robert C. Byrd Green Bank Telescope
Very Long Baseline Array



#### **Talk Outline**

- Current offset pointing / tracking performance
- Current surface accuracy
- Recent, current and potential future servo upgrades
- Potential new surface measurement system
- Summary
- Call for collaborators!



# Offset Pointing and Tracking Accuracy



#### Offset Pointing / Tracking at 3 mm

Point on bright quasar



2) Track source of interest

$$g(\rho) = \exp\left[-4\ln 2\left(\frac{\rho}{\theta}\right)^{2}\right]$$

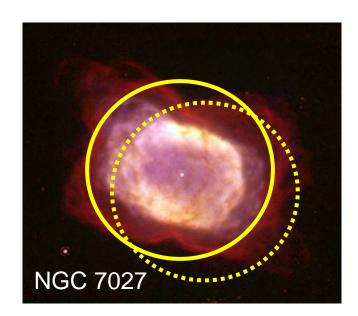
$$\left(\frac{\theta}{740 \text{ arcsec}}\right) \approx \left(\frac{\text{GHZ}}{\nu}\right)$$

$$\left\langle \rho^{2} \right\rangle \equiv \sigma_{2}^{2} = \sigma_{\text{Az}}^{2} + \sigma_{\text{El}}^{2}$$

$$f \equiv \left(\frac{\sigma_{2}}{\theta}\right)$$

$$\text{Good} \quad (\sigma_{s} = 5\%) \to f \approx 0.14$$

$$\text{Usable} (\sigma_{s} = 10\%) \to f \approx 0.20$$

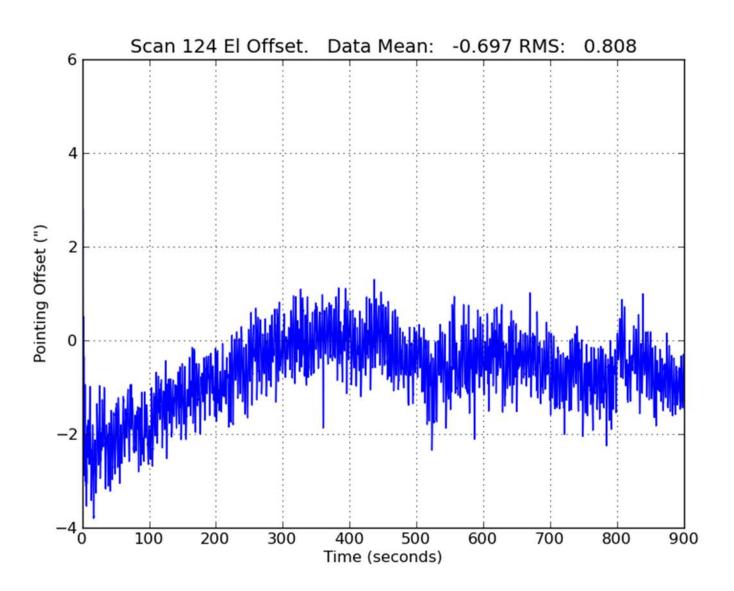




At 115 GHz, beam size = 6.5", so the tracking requirement for 10% error = 1.3" rms

Slide after Todd Hunter

### Offset Pointing / Tracking at 3 mm



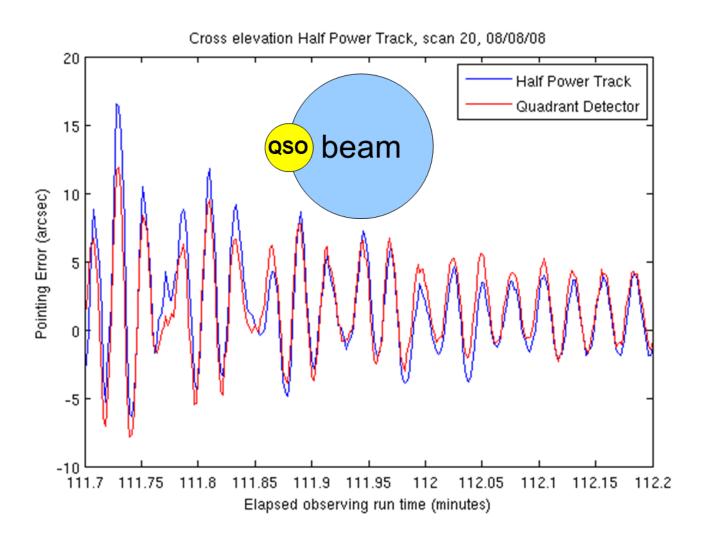


#### Offset Pointing / Tracking at 3 mm

- Recently replaced 22-bit BEI encoders with 26-bit Heidenhains
- Best (benign time-time) performance before encoder upgrade:
  - 1.7" (just acceptable for 90 GHz)
- Now, potentially (not yet measured):
  - $\leq$  I" offset pointing, plus  $\leq$  I" tracking over  $\sim$  30 min
  - Extremely close to 115 GHz specification!
- Under non-ideal conditions, performance degrades due to thermal and wind effects
- Thermal deformations compensated for by pointing/focus
- Wind effects are (primarily) "feed arm sway"

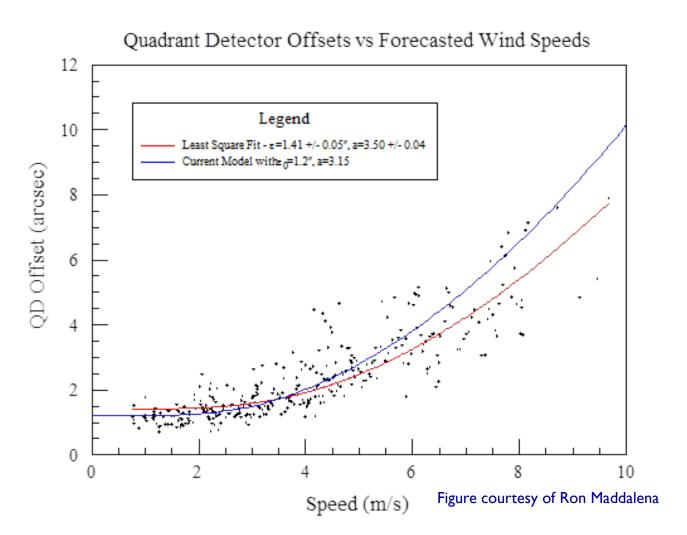


### Monitoring feed-arm sway with the QD



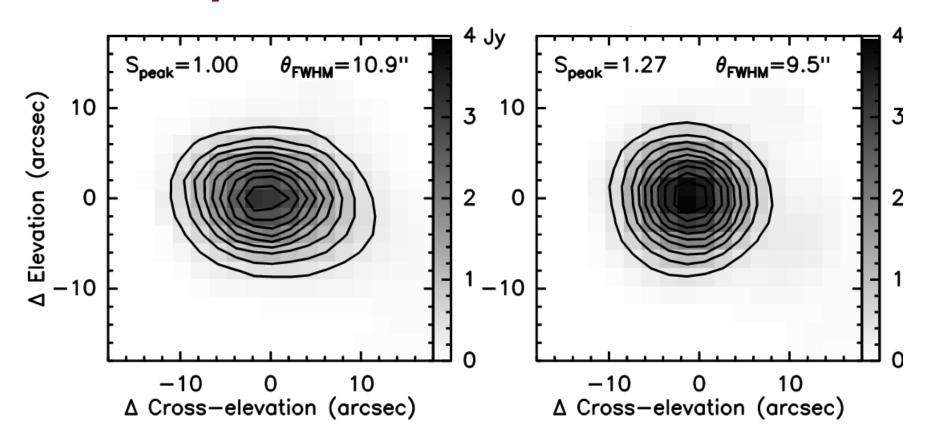


### Feed-arm sway as a function of wind speed





### For array detectors, correct off-line



MUSTANG images of a bright quasar obtained on a windy ( $\approx 6 \text{ m s}^{-1}$ ) night combining several scans with a total integration time of about 20 minutes. Beam size under ideal conditions is 9.0" ((Ries et al., 2011)



# Summary offset pointing / tracking performance

- Under benign (no-wind) night-time conditions:
  - $\le 1$ " offset pointing, plus  $\le 1$ " tracking ( $\sim 1.5$ " total)
- Current DSS equation:

$$\sigma_{tr}^2 = \sigma_0^2 + \left(\frac{s}{a}\right)^4$$

$$\sigma_0 = 1.3$$
"  $a = 3.5 \, ms^{-1} \, arcsec^{-2}$ 

$$= 2.5$$
" @ 5 m s<sup>-1</sup> (11 mph)



### **Current Surface Accuracy**



### Surface Requirements for 3mm

• Aperture efficiency given by:

$$-\eta_{aperture} = \eta_{surface} \times \eta_{illumination} \times \eta_{spillover} \times \eta_{blockage}$$

For the GBT:

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- off-axis design: \eta_{blockage} = I
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- I4dB taper: 
$$\eta_{\text{spillover}} = 0.9895$$
;  $\eta_{\text{illumination}} = 0.707$ 

- Ruze formula: 
$$\eta_{\text{surface}} = \exp(-(4\pi\sigma/\lambda)^2)$$

• GBT specification on surface rms,  $\sigma \le (\lambda / 4\pi)$ :

- i.e. 
$$\eta_{surface} = 1/e = 0.367$$

- For 115 GHz we need:
  - $-\sigma \leq 210 \,\mu m \,rms$

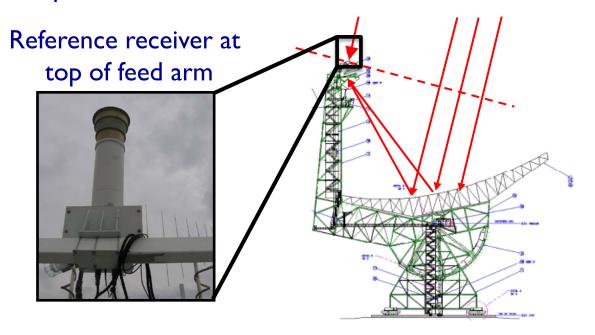
# Surface Measurement: High-resolution with-phase holography

- Technique is > 30 years old (Bennett et al. 1976)
  - Measure complex beam pattern (phase and amplitude)
  - Fourier transform to get phase and amplitude of E field on aperture
  - Convert phase to surface error, and apply mechanical corrections
- 2 Receivers: room-temp. LNAs, 10kHz filters, Hilbert transform correlator

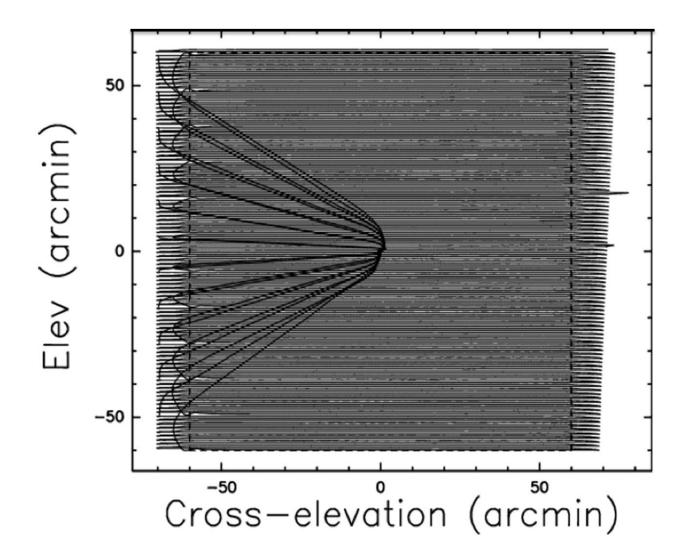
Main receiver in Gregorian turret





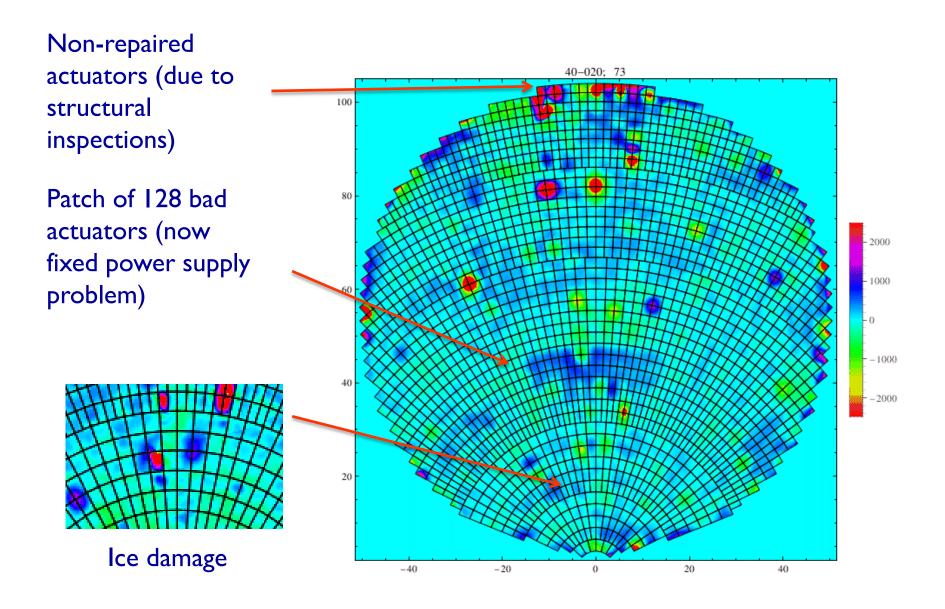


### Scan Pattern (2° x 2° Map)





### Recent (not ideal) Results

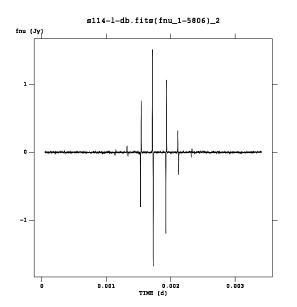


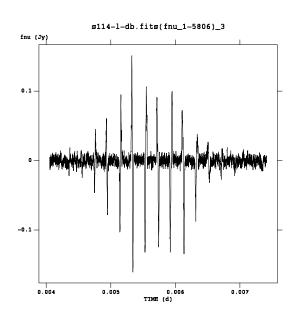
# Surface Measurement: Low-resolution "OOF" holography

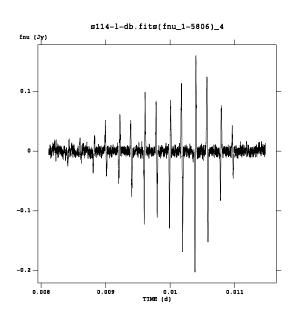
- Make three Nyquist-sampled beam maps, one in focus, one each ~ five wavelengths radial defocus
- Model surface errors (phase errors) as combinations of low-order Zernike polynomials. Perform forward transform to predict observed beam maps (correctly accounting for phase effects of defocus)
- Sample model map at locations of actual maps (no need for regridding)
- Adjust coefficients to minimize difference between model and actual beam maps.



## Typical data



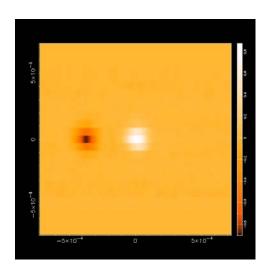


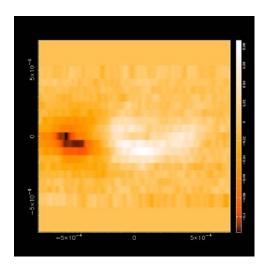


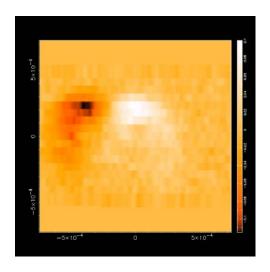
Q-band (43 GHz)

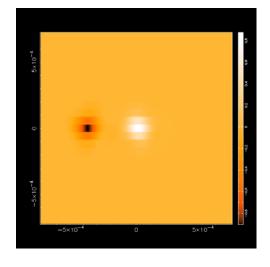


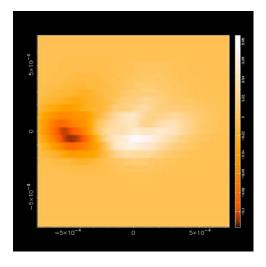
## Typical data: before (rms = $370 \mu m$ )

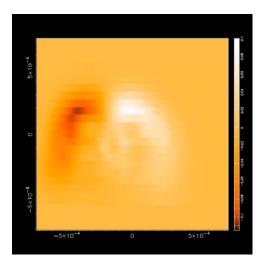




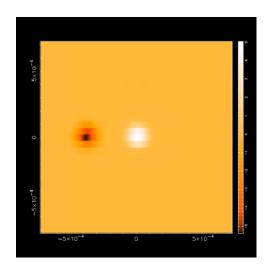


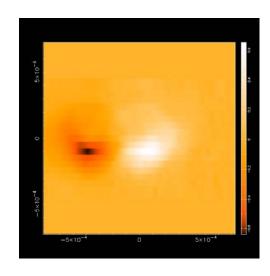


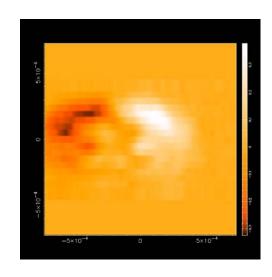


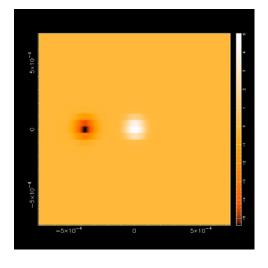


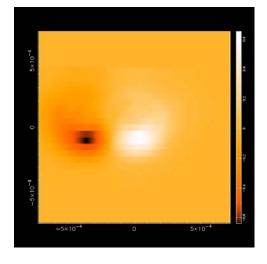
## Typical data: after (rms = $80 \mu m$ )

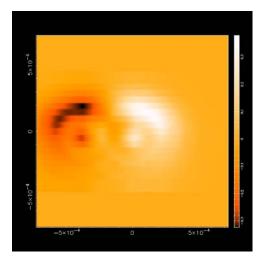












#### **Estimated error budget**

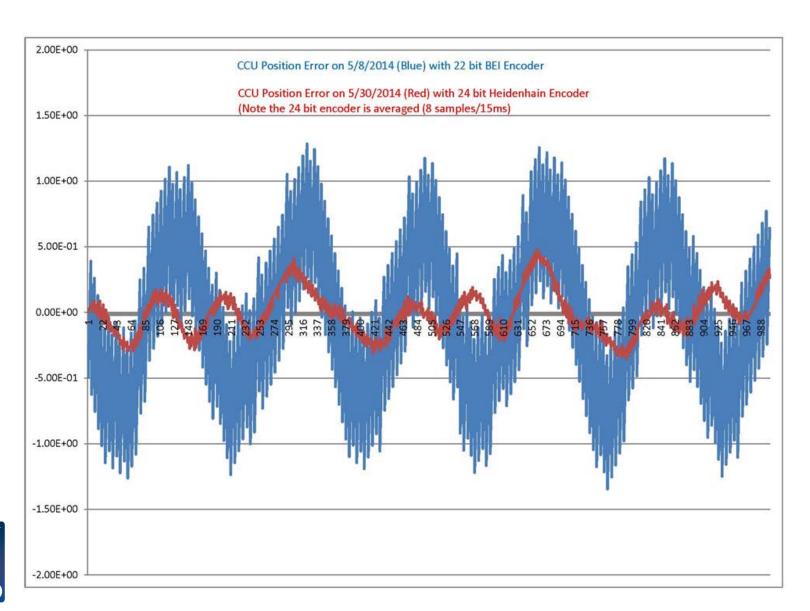
- Small-scale errors:
  - panel manufacturing plus gravity error: I 27 μm
  - residual actuator error: 80 µm
  - panel corner setting error: 80 µm
  - panel-scale thermal errors: 100 μm
  - subreflector: 75 μm
  - total small-scale error:  $80+80+127+100+75 = 211 \mu m$
- Large-scale thermal error: 100 μm
- Total error:  $211+100 = 235 \mu m$



# **Current and Potential Servo Improvements**



## **Encoder Replacement (Complete)**



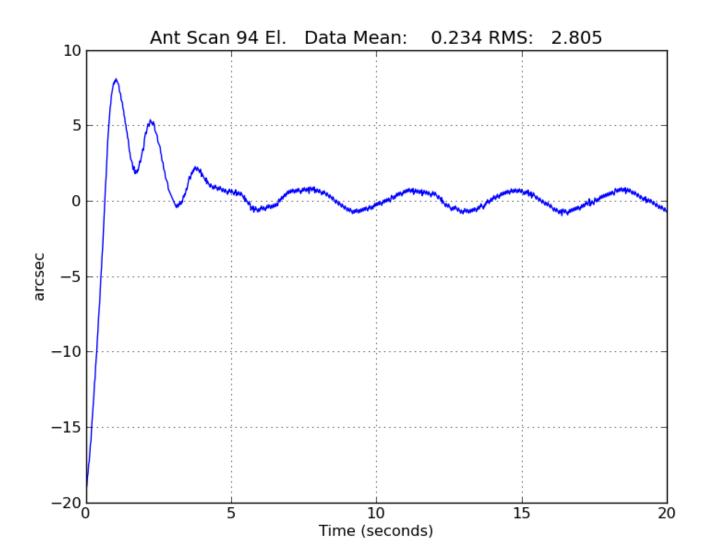


# Servo / Motor Control System Upgrade (Ongoing)

- Primarily a required maintenance upgrade
  - Existing system is > 20 years old, and obsolete
- Replacing existing custom, wire-wrap "rate loop" cards with modern, PCB, plug-compatible version
  - Controls the (analog) velocity loop
  - Controls the individual motors (torque bias; faulted motors...)
- Replace the current VxWorks / VME, proprietary "Central Control Unit", with an open-source, Linux / PC version
  - Initial installation; replicate current functionality
  - Replace "lead-lag" servo algorithm with PID algorithm
  - Replace analog rate loop with digital rate loop



# Servo / Motor Control System Upgrade (Ongoing)



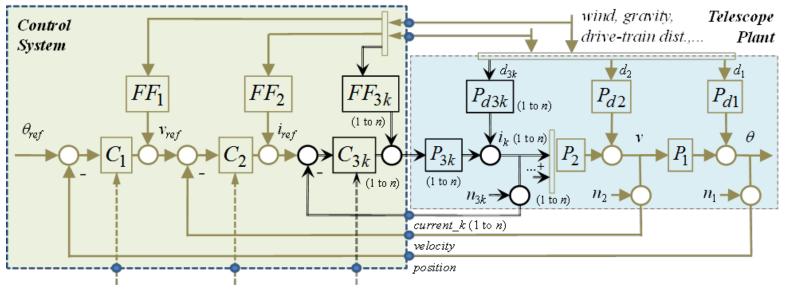


# Servo / Motor Control System Upgrade (Proposed)

- Develop a modern control system replacement for main drive and subreflector servos
- Collaboration with Case Western Reserve University (Prof.Mario Garcia-Sanz, Trupti Ranka)
- High-fidelity plant model based on improved FE model (Art Symmes)
- Develop "Observers" using additional sensor data (quadrant detectors, accelerometers)
- Improve wind rejection, damp feed-arm vibration
- Compensate for "feed arm sway" (timescales of ~ seconds) using the subreflector



#### **New Servo Architecture**



Accelerometers and Quadrant detector

#### This project is not funded!



## **Potential Surface Improvements**



#### Leica ScanStation P40

System Accuracy	
Accuracy of single measurement ^ Range accuracy Angular accuracy	1.2 mm + 10 ppm over full range 8" horizontal; & vertical
3D position accuracy	3 mm at 50 m 6 mm at 100 m
Target acquisition	2 mm standard deviation at 50 m
Dual-axis compensator	Liquid sensor with real-time onboard compensation, selectable on/off, resolution 1", dynamic range ±5', accuracy 1.5"

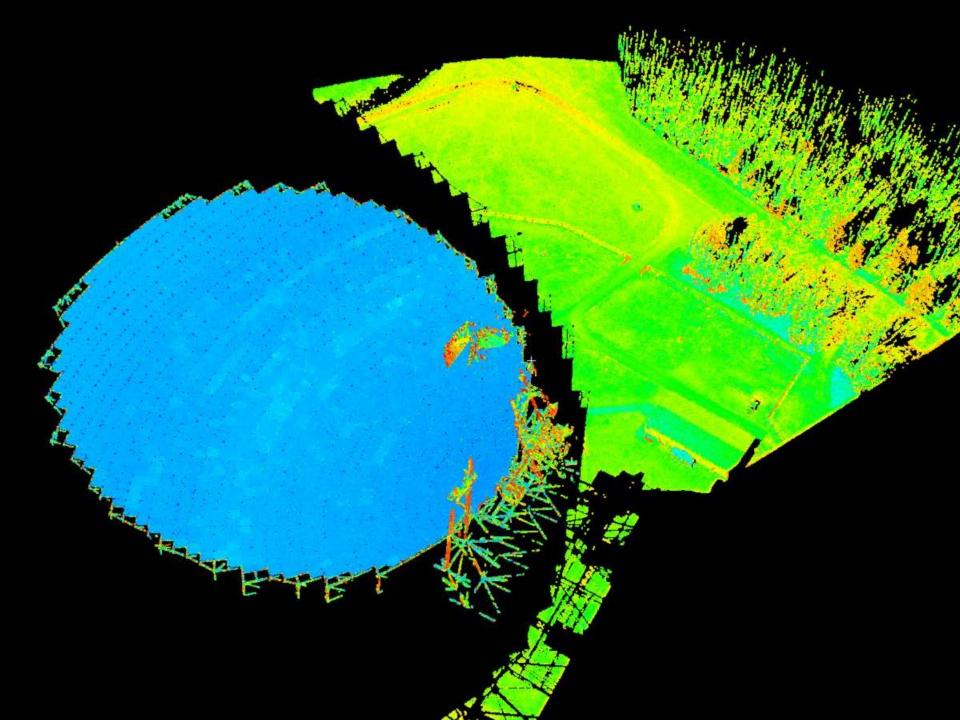
Distance Measurement S	ystem			
Туре		eed time-of-flight FD) technology	enhanced	by Waveform
Wavelength	1550 nm (invis	1550 nm (invisible) / 658 nm (visible)		
Laser class	1 (in accordan	1 (in accordance with IEC 60825:2014)		
Beam divergence	< 0.23 <u>mrad</u> (	< 0.23 mrad (FWHM, full angle)		
Beam diameter at front window	≤3.5 mm (FW	≤ 3.5 mm (FWHM)		
Range and reflectivity	Minimum rang	je 0.4 m		
	Maximum range at reflectivity			eflectivity
		120 m	180 m	270 m
	P30	18 %	-	-
	P40	8 %	18 %	34 %
Scan rate	Up to 1000'00	Up to 1000'000 points per second		
Range noise *	0.4 mm rms at 10 m 0.5 mm rms at 50 m			
Field-of-View Horizontal Vertical	360° 270°			
Data storage capacity		256 GB internal solid-state drive (SSD) or external USB device		
Communications/ Data_transfer	•	Gigabit Ethernet, integrated Wireless LAN or USB 2.0 device		
Onboard display		Touchscreen control with stylus, full colour VGA graphic display (640 × 480 pixels)		
Laser plummet	Centring accu	(IEC 60825:2014) Iracy: 1.5 mm at 1 meter: 2.5 mm at I/OFF		



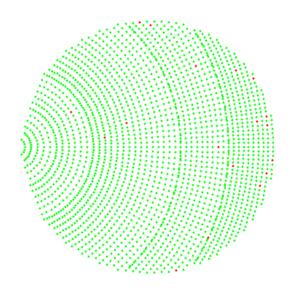
3D accuracy: 6mm @ 100m

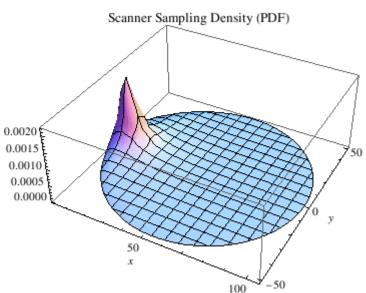
Range noise: 0.5mm @ 50m

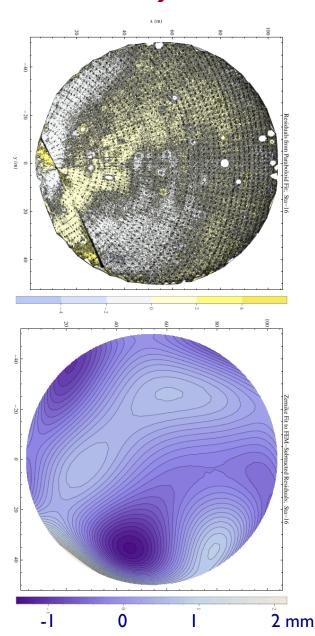




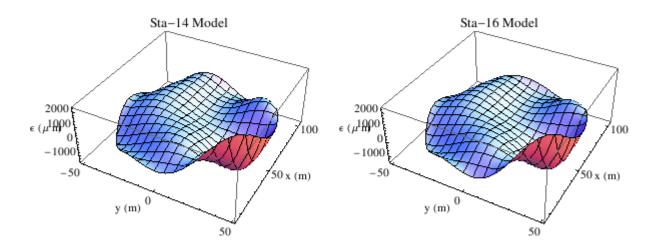
# **Data Processing (Fred Schwab)**



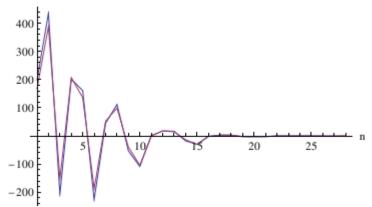


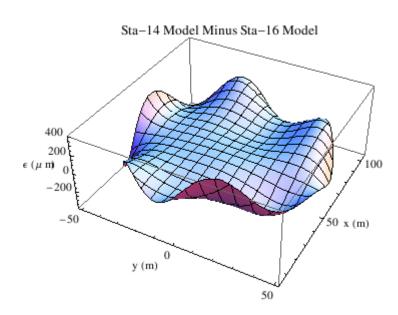


## Data Processing (Fred Schwab)

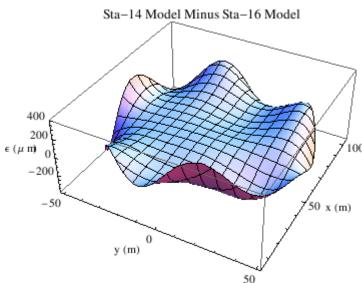


Sta-14 (blue) and Sta-16 (red) best-fit Zernike coefficients Zn (microns)

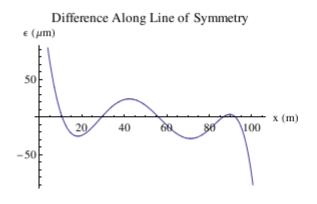


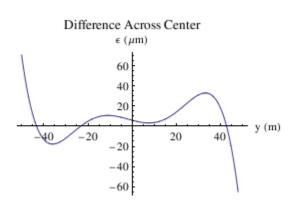


## Data Processing (Fred Schwab)



Difference of two successive runs (best case!)





#### **RMS Differences:**

- $|r| < 50m = 68 \mu m$
- $|r| < 45m = 48 \mu m$
- $|r| < 40m = 38 \mu m$

RMS of single a measurement: ~ 35µm

RMS of two subsets of the same run  $\sim 18 \mu m$ 

best case OOF holography: ~ 40µm

#### **Estimated error budget**

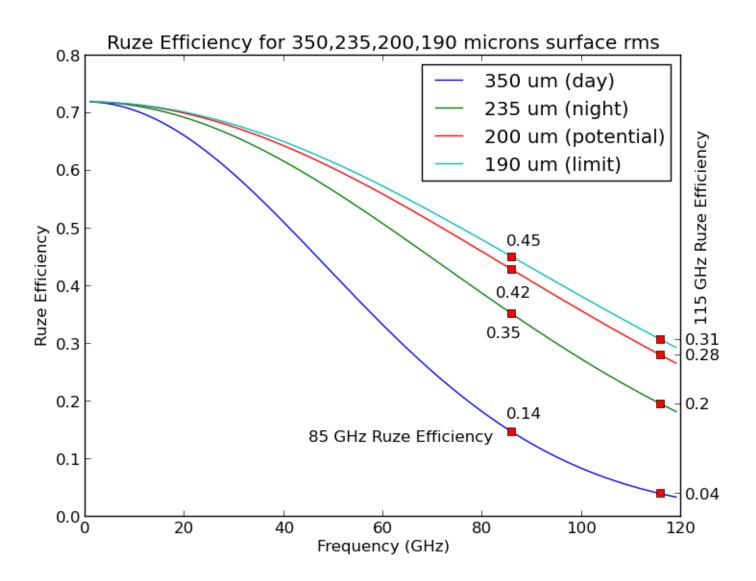
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  - panel manufacturing plus gravity error: I 27 μm
  - residual actuator error: 80 μm
  - panel corner setting error: 80 μm
  - panel-scale thermal errors: 100 μm
  - subreflector: 75 μm
  - total small-scale error: 80+80+127+100+75 = 211 μm
- > Large-scale thermal error: 100 μm 30 μm
- > Total error:  $211+100 = 235 \, \mu \text{m}$  (~ 200  $\mu \text{m}$ !)

201 μm

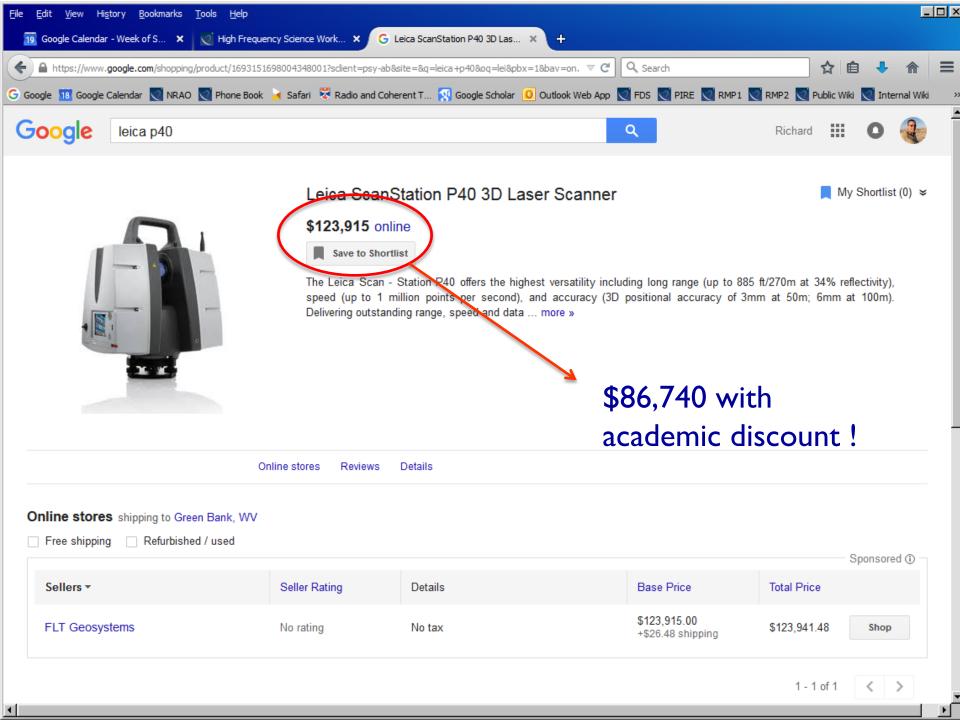
- Current OOF overhead ~ 20 m in 2 hr ~ 17%
- Potential Scan overhead ~ 10 m in 2 hr ~ 8%
- Independent of atmospheric opacity
- Daytime observing becomes possible!



### **Ruze Efficiency**







## **Performance Summary**

Wind	$\sigma_{2}$	
Benign Night-time	< 1.5"	
5 ms <sup>-1</sup>	~ 2.5"	
10 ms <sup>-1</sup>	~ 8"	

Tracking Accuracy

Surface RMS	η <sub>surface</sub> 85 GHz	η <sub>surface</sub> 115 GHz
350μm	0.14	0.04
235μm	0.35	0.20
200μm	0.42	0.28

**Surface Efficiency** 



#### **Collaborators Needed!**

- We are currently about to commission three new ~ 3-4mm instruments!
  - Argus
  - MUSTANG 2
  - U. Mass. / BYU Phased Array Feed + Beamformer
- All are being funded by current or recent NSF AST Grants!
- We need to propose two further Proposals:
  - GBT Servo Improvements (PI Mario Garcia-Sanz, CWRU)
  - New Surface Measurement System (PI me)



Please come by and talk to me in the next few days!

