ALMA Future Sc. Program Development Workshop Aug. 2016, Charlottesville

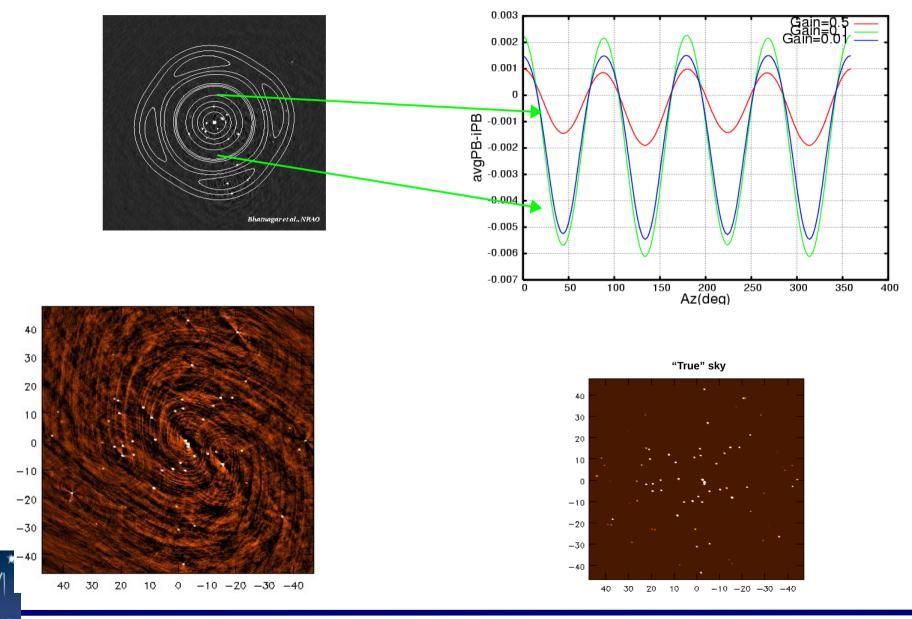


Wide-field Wide-band Full-polarization Imaging: Impact of Direction Dependent PB effects



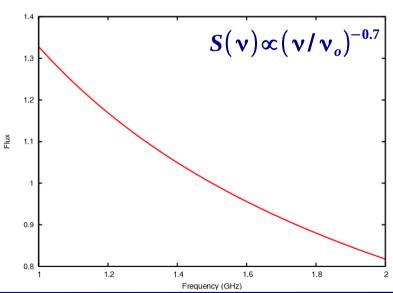
S. Bhatnagar, U. Rau, P. Jagannathan, K. Kundart

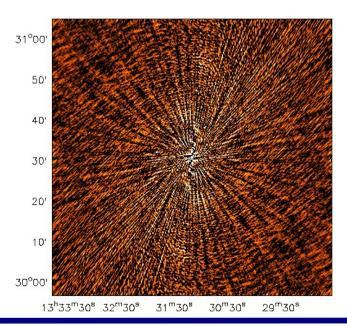
What is wide-field?



What do we call wide-band?

- When fractional signal bandwidth used for imaging > ~20%
 - Plus source spectral index >= -1.0
 - Plus target dynamic range > 1000
- Spectral effects for higher source spectral index will become significant at lower bandwidth ratios
 - Empirical Dynamic range : $\frac{I\alpha}{100}$
 - Spectral line imaging, by definition, does not require wide-band imaging algorithms

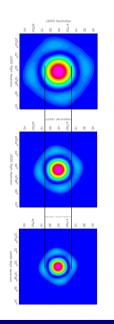


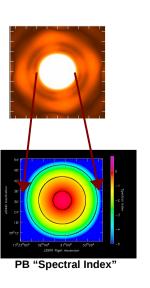


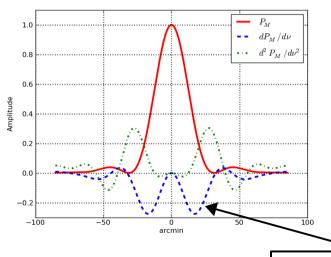


Wide-band Wide-field Imaging

- Wide band data to image beyond the ~50% point of the PB at a reference frequency
 - Bandwidth ratio > ~20%
 - FoV > ~HPBW @ reference frequency
 - Variable PB:
 - Long integration (rotation), Mosaicking (pointings at different PA), in-beam polarization is large (AA)





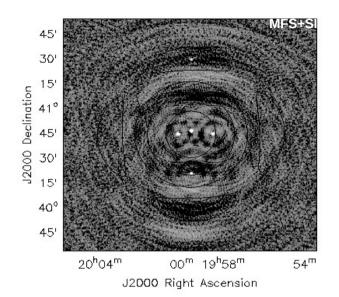


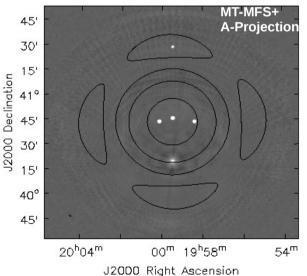
PB Frequency dependence (blue curve)

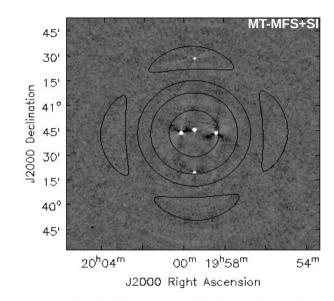


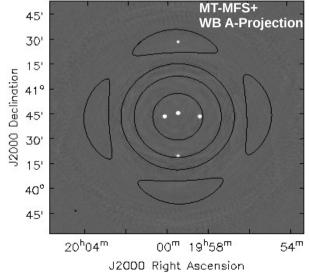
Wide-band Wide-field Imaging

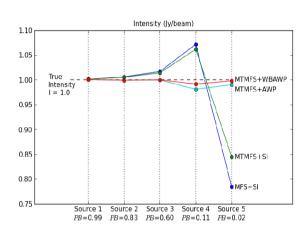
Characterization of the (WB) A-Projection + MT-MFS

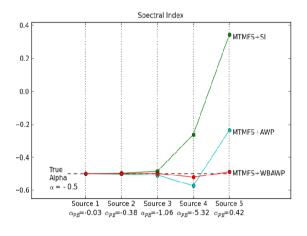










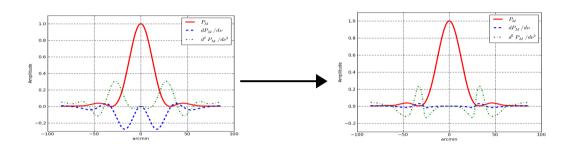


2016, AJ (accepted) [arXiv:1605.07640]

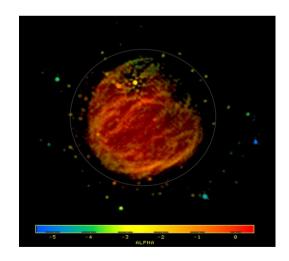


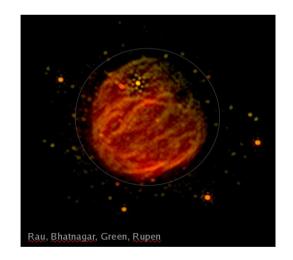
Wide-band Wide-field Imaging

- WB A-Projection + MT-MFS
 - WB A-Projection for PB



MT-MFS for sky







(2011, ApJ, 739, L20 [arXiv:1106.2796v2])

Full-Mueller (Polarization) WF-WB Imaging

Direction Dependent (DD) Measurement Eq. (ME) in the image domain:

$$I^{Obs} = [M] \cdot [I^o]$$

$$\begin{vmatrix} I_I^{Obs} \\ I_Q^{Obs} \\ I_U^{Obs} \\ I_V^{Obs} \end{vmatrix} = \begin{vmatrix} M_{11} & M_{12} & M_{13} & M_{14} \\ M_{21} & M_{22} & M_{23} & M_{24} \\ M_{31} & M_{32} & M_{33} & M_{34} \\ M_{41} & M_{42} & M_{43} & M_{24} \end{vmatrix} \cdot \begin{vmatrix} I_I^o \\ I_U^o \\ I_V^o \end{vmatrix}$$
 • Diagonal: "pure" poln. products • Off-diagonal: Include poln. leakage

DD ME in the visibility domain:

$$V_{ij}^{Obs} = \left[A_i \otimes A_j^T\right] * \left[V_{ij}^o\right] = \left[A_{ij}\right] * \left[V_{ij}^o\right]$$

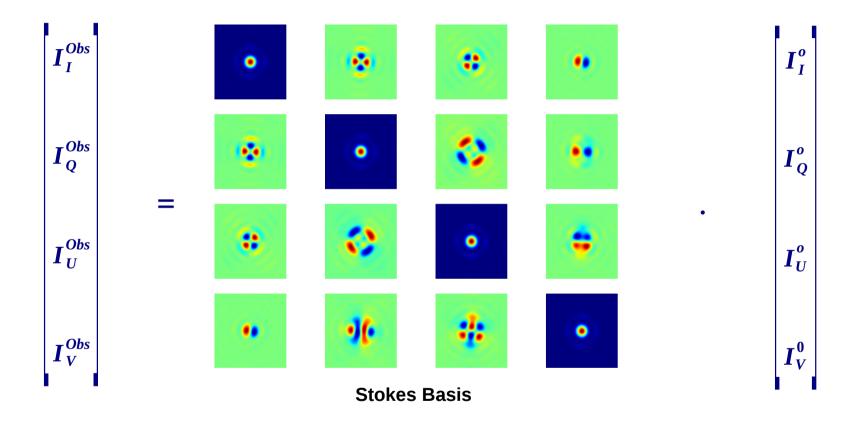
where
$$A_i = \begin{bmatrix} A_p & A_{p \rightarrow q} \\ A_{q \rightarrow p} & A_q \end{bmatrix}$$

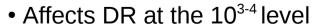


Antenna Aperture Illumination

DD Effects in Full-pol. Imaging

DD "Mueller" matrix:



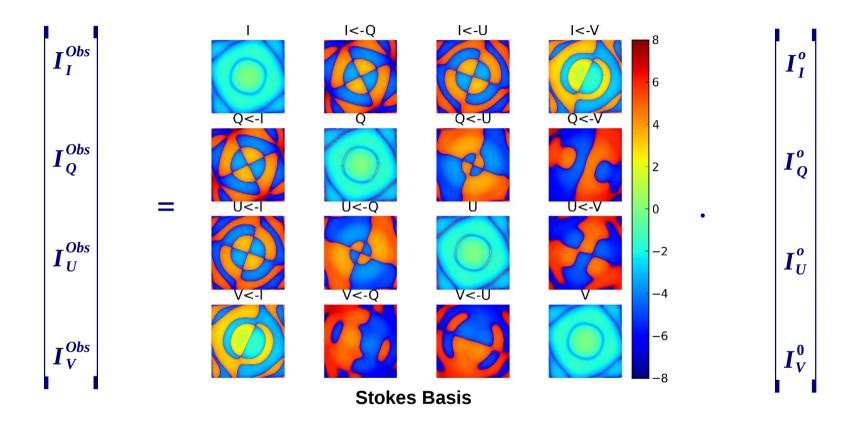


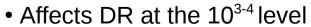
[•] PB Stokes-Q, -U is few% of Stokes-I



DD Effects in Full-pol. Imaging

DD "Mueller" matrix:





[•] PB Stokes-Q, -U is few% of Stokes-I



Issues in Wide-field Wide-band Full-Pol. Imaging

PB Effects

- In-beam effects : DD Leakage
- Parametric Aperture Illumination model (Holographic measurements not sufficient)
- Pointing Errors
- Mosaic patterns
- Variations with frequency
 - Frequency dependence of intrinsic Q and U
 - Frequency dependence due to PB

Computing load

- Fundamentally more expensive : Larger CF support
- Larger memory footprint: Fundamentally required, any which way you cut it



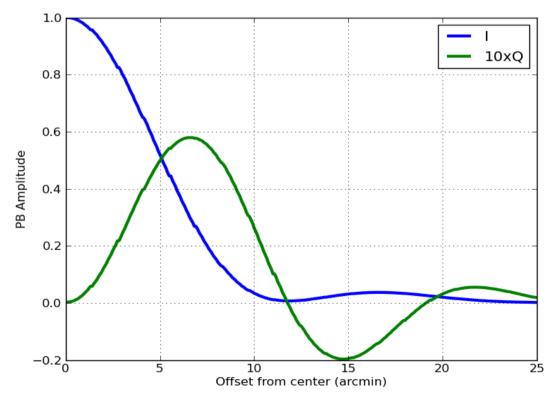
Full-pol. Imaging: In-beam Leakages

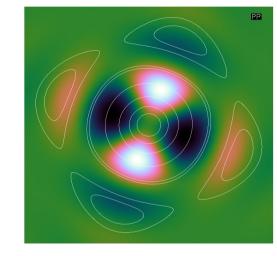
Leakage (Off-diagonal elements of the Mueller matrix)

Vary with direction (position in the beam), Parallactic Angle (time) and

frequency







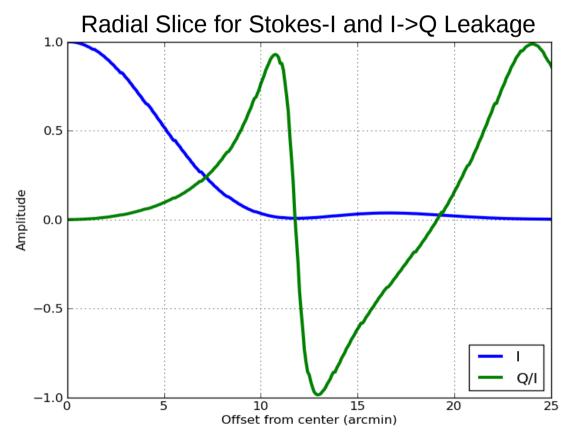


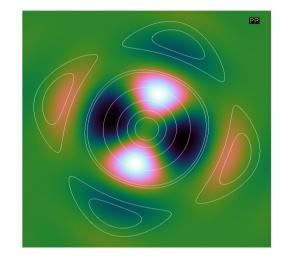
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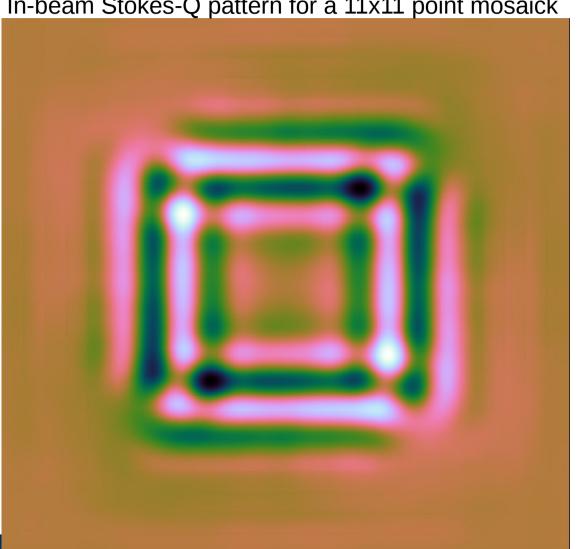






Full-pol. Imaging: Mosaic Sensitivity Pattern

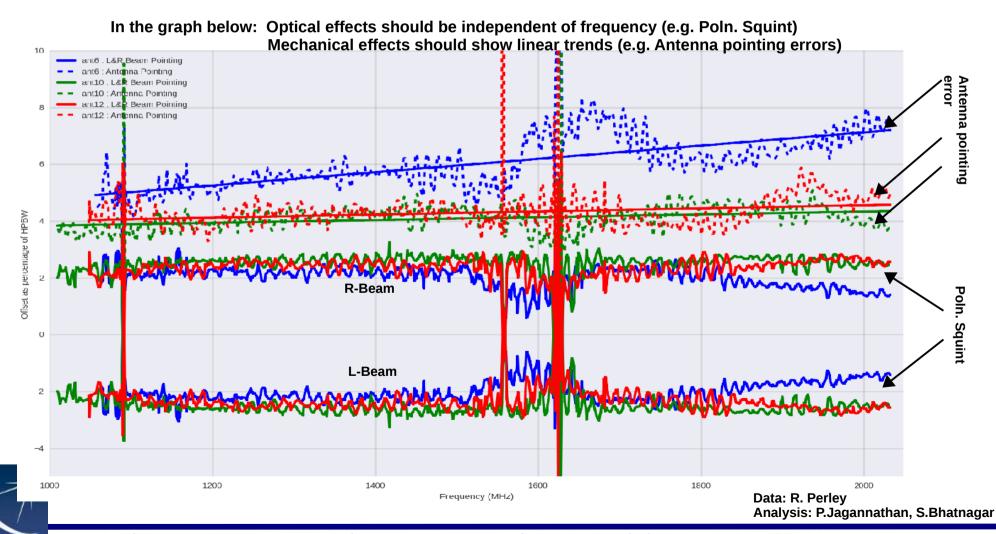
In-beam Stokes-Q pattern for a 11x11 point mosaick



- Heterogeneous case; rotation due to PA change also ignored
- The resulting pattern is combination of overlapping Clover-leaf pattern of each pointing
- · In-beam DD leakage spreads all across the mosaicked region.

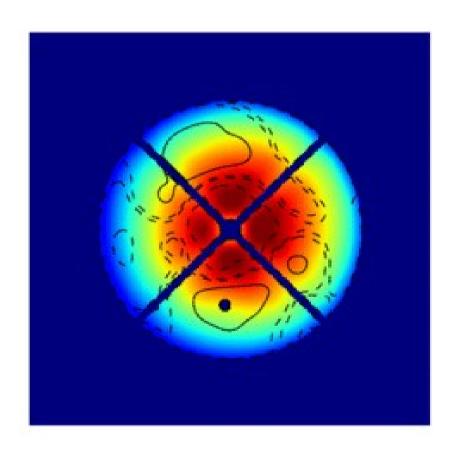
[Jagannathan, Bhatnagar, Rau & Taylor]

- Parametric model of antenna Aperture Illumination
 - Difference between Ant6 and Ant10 in "homogeneous array"



Parametric model of antenna Aperture Illumination

Aperture Illumination

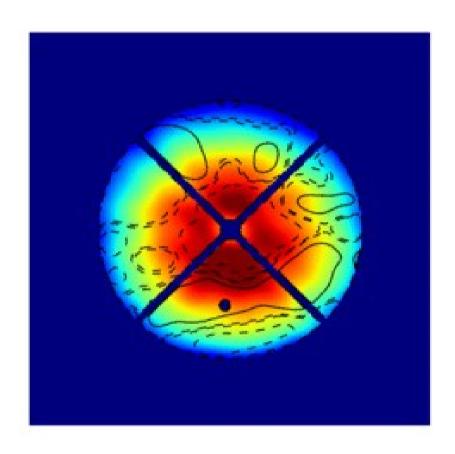


- Strongest effects
 - Antenna size, Pointing errors
 - Quadrapods
 - Antenna-to-antenna variations



Parametric model of antenna Aperture Illumination

Aperture Illumination

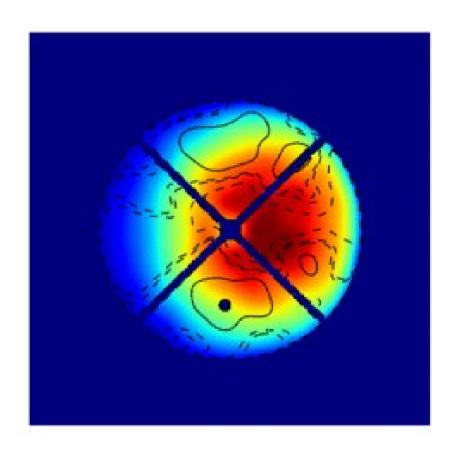


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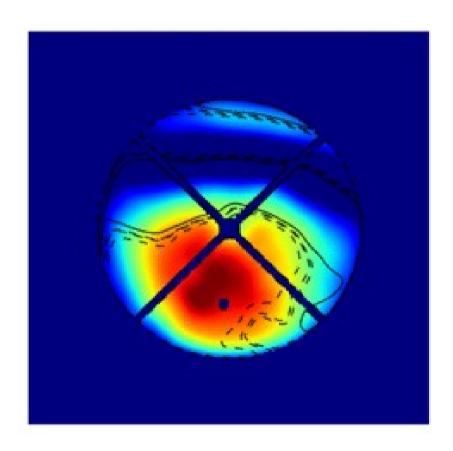


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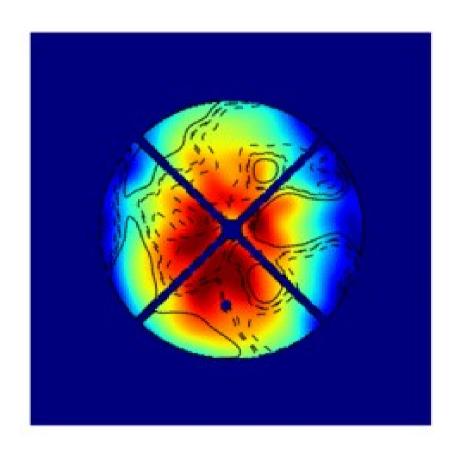


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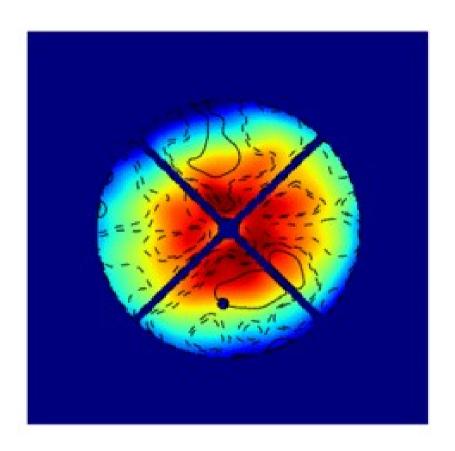


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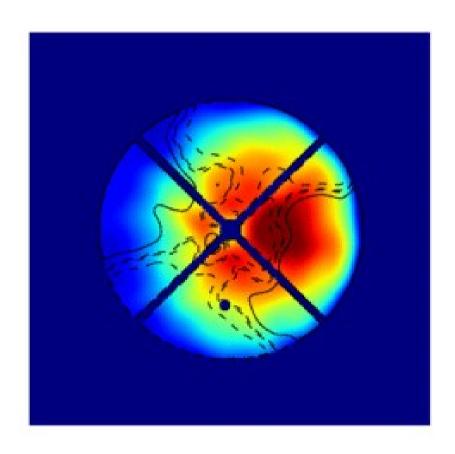


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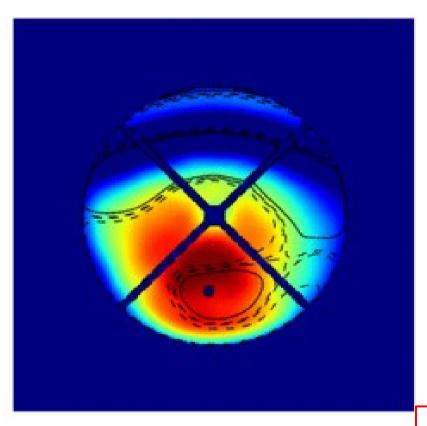


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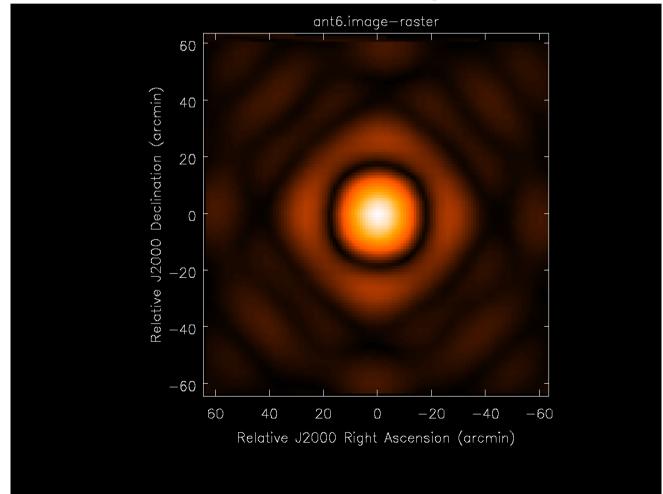


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- Cost equations
 - Cheaper to fix the hardware?
 - Cheaper to handle in post processing?

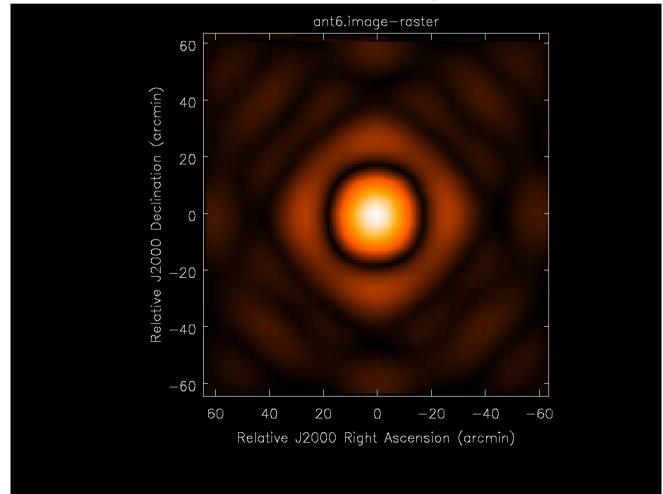


- EVLA Squint
 - Expected: Lateral translation between RR- and LL-beams
 - Measure: Translation + Rotation (~1 deg!)





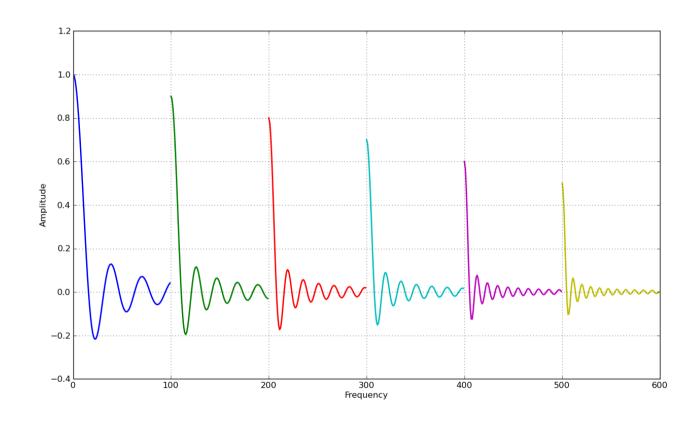
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Full-pol. Imaging: Q/U vs. Frequency

- MFS approach to gain from WB sensitivity
 - PSF structure scales with frequency
 - Effective PSF amplitude also changes with frequency (Sp. Ndx., PB scaling)

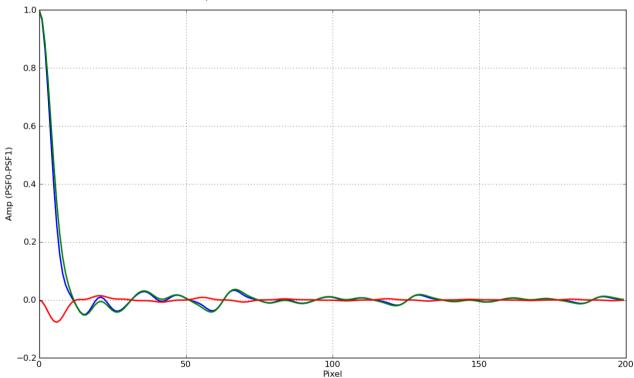




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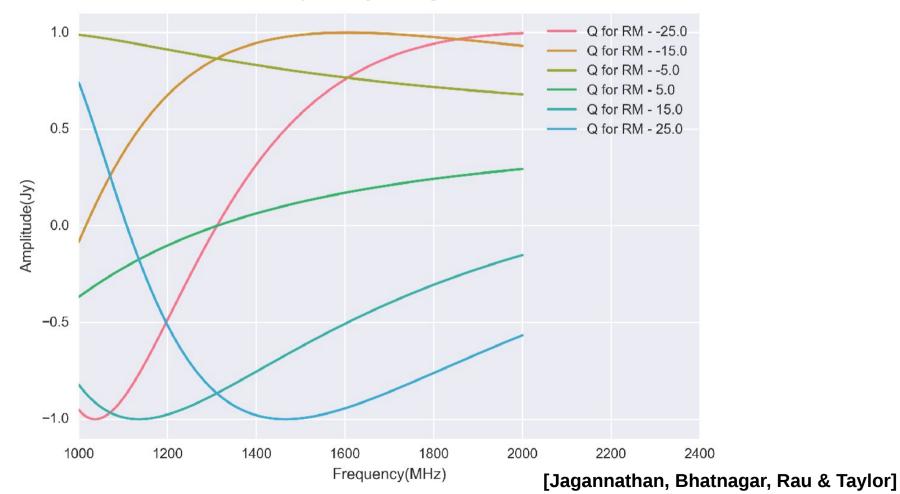
$$PSF(x_o)_{Continuum} = \sum_{v} I(x_o, v) PSF(x - x_o, v)$$



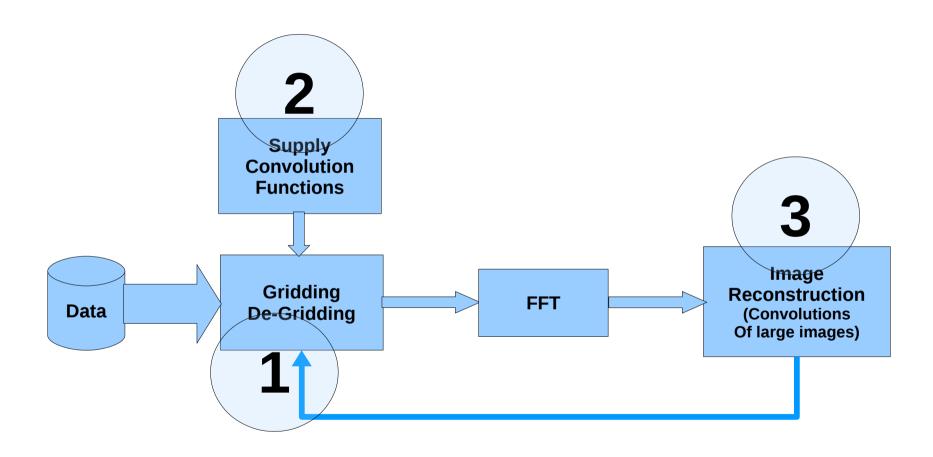


Full-pol. Imaging: Q/U vs. Frequency

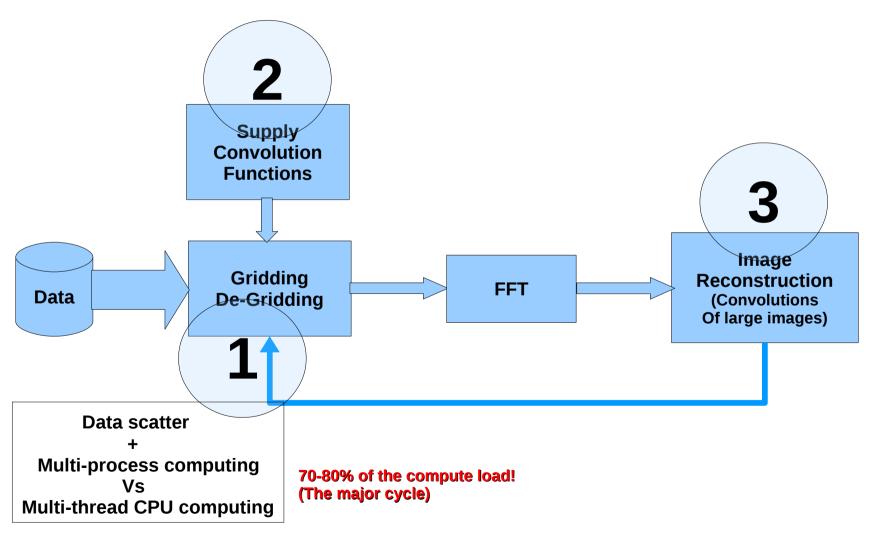
- Possible to use MT approach on Q and U for low RM cases
 - Polynomial model (a la MT-MFS) tested for RM = -25 to +25
 - Even easier for ALMA frequency range



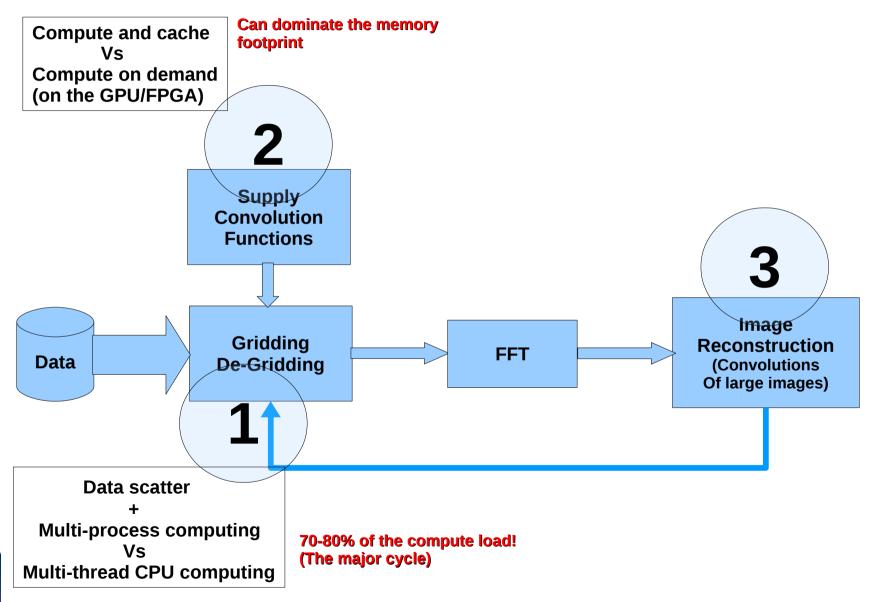


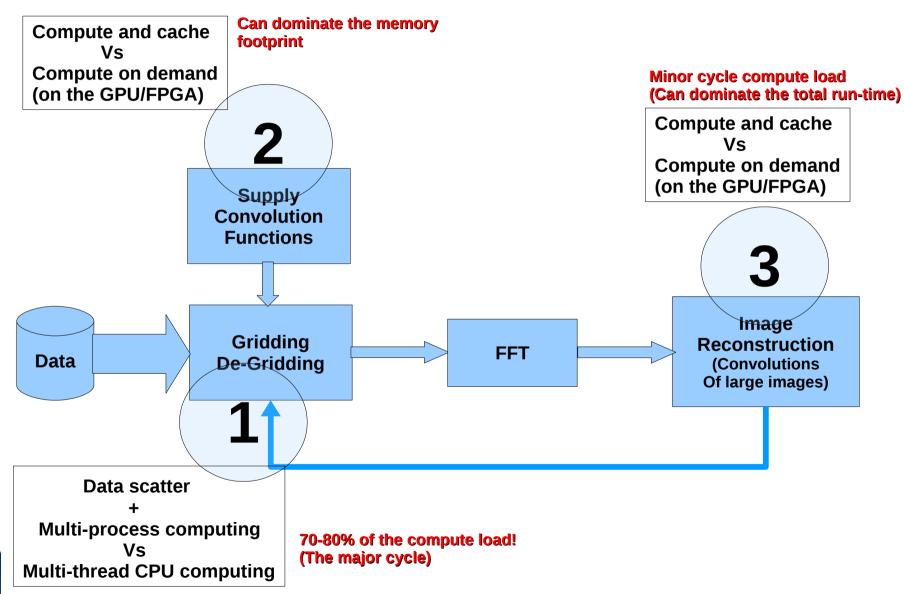






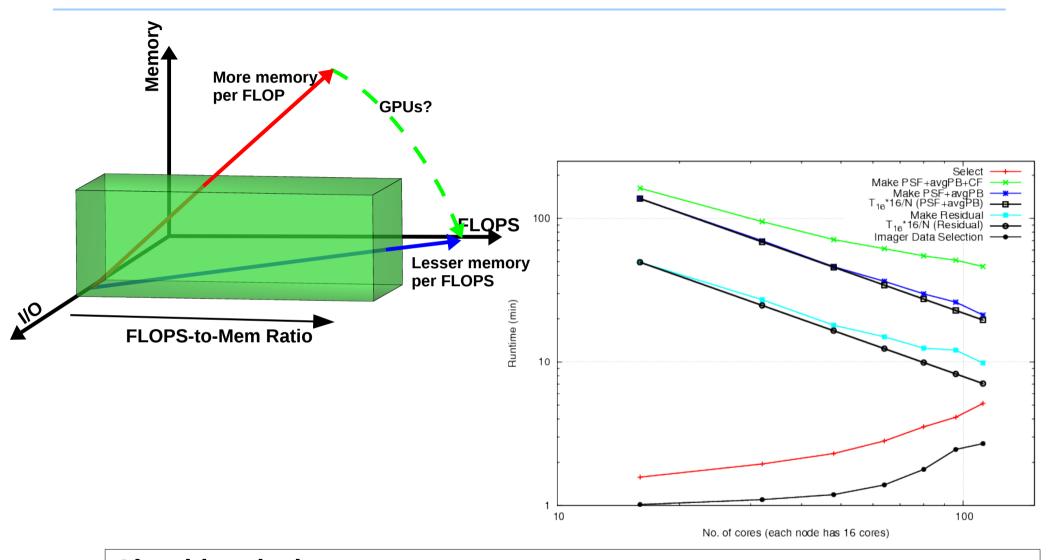








Algorithm architecture





Algorithm design

- Move towards algorithms with higher compute-to-I/O ratio
- Use cheap massively parallel h/w (low memory footprint at the cost of higher computing)

Issues in Wide-field Wide-band Full-Pol. Imaging

PB Effects

- Develop parametric models for antenna aperture illumination
- Extend full Mueller imaging tools for heterogeneous array case
 - Full-pol imaging for EVLA under test as part of P. Jagannathan's thesis

Variations with frequency

- Assess extension of MT-MFS for full pol.
 - Improved multi-scale algorithms in general
- Assess the range of freq. dependencies of the PB to be included

Computing load

- Computing: Parallel processing on a routine basis
- Memory footprint: Use heterogeneous hardware (CPUs + GPUs, etc.)
- Use Cloud Computing as an effective development platform.