# Finding Milky Way Black Holes with the NGVLA



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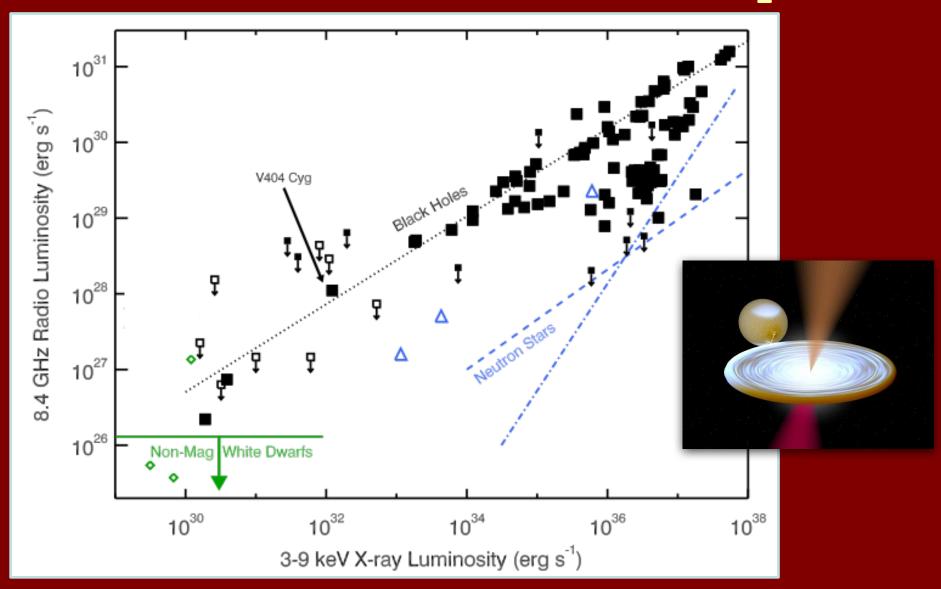
B. de Bivort

### Searches for accreting BH binaries can answer:

- How many BH X-ray binaries are there?

  Current estimates span  $10^2 10^8$  (Tetarenko+ 2016)
- What natal kicks do BHs receive?
- What is the mass spectrum of stellar-mass BHs?
- What is the relative importance of dynamical and binary channels in forming LIGO mergers?

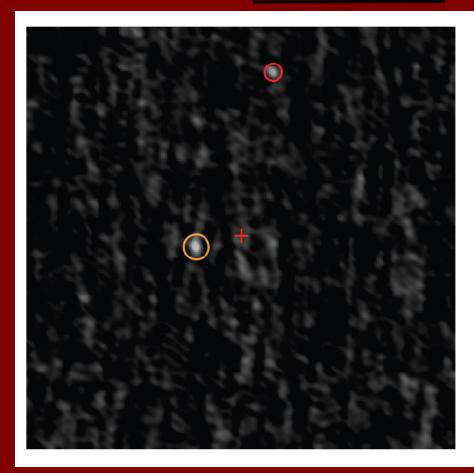
### Finding low-luminosity Black Holes with radio/X-ray

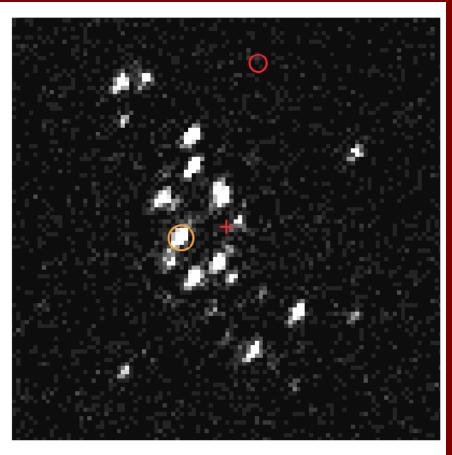


### What BH candidates look like

25"

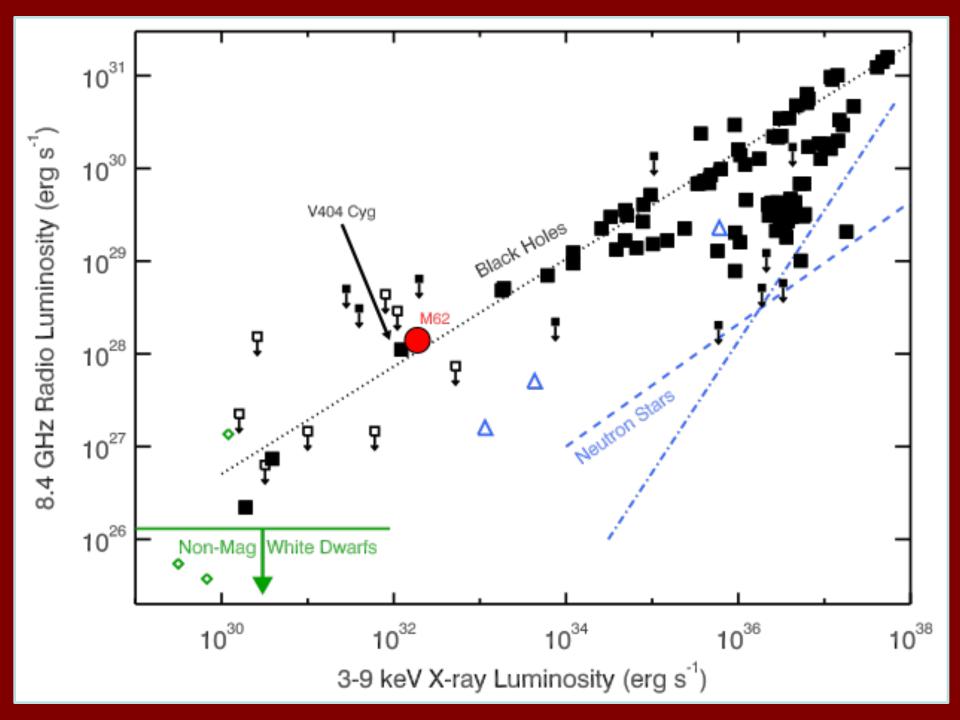
Chomiuk et al. 2013





VLA

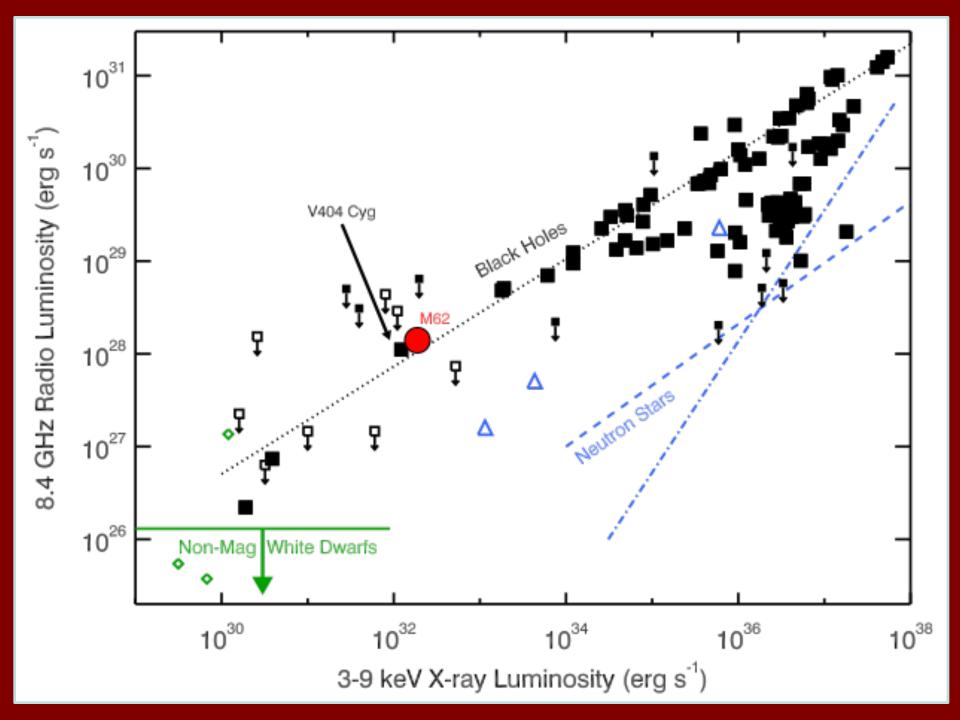
Chandra



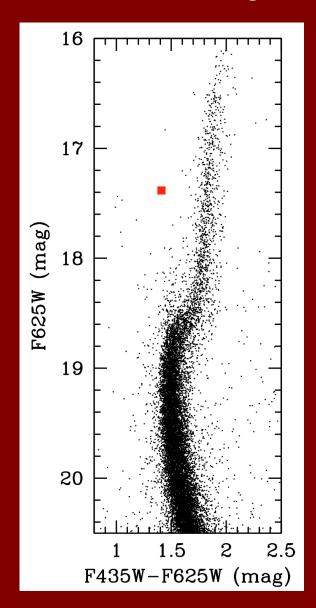
### We can find contaminants.

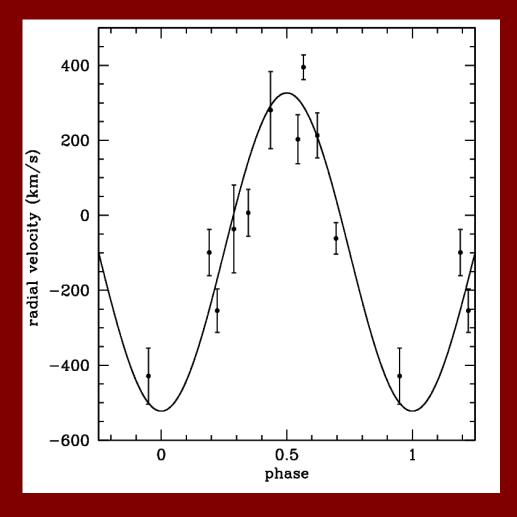


Background AGN are visible as galaxies in optical images.



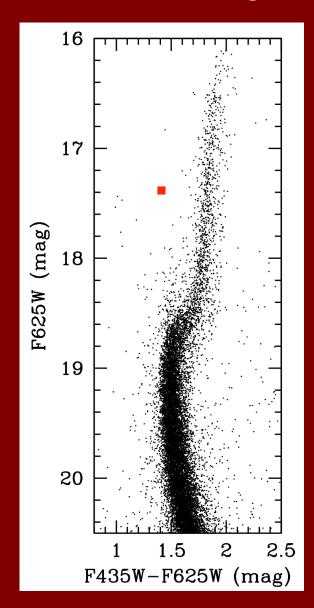
### M62 BH Candidate

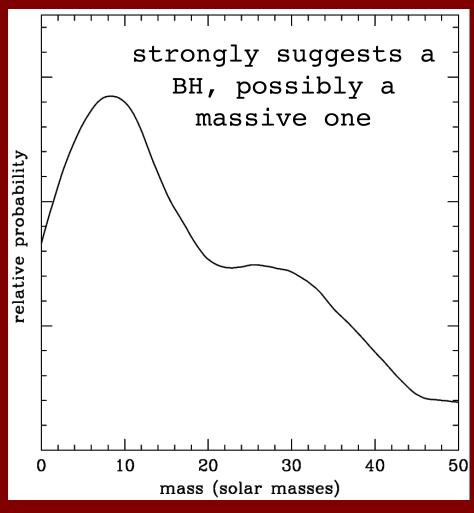




Radial velocities of the red giant companion constrain the mass of the invisible compact object

### M62 BH Candidate



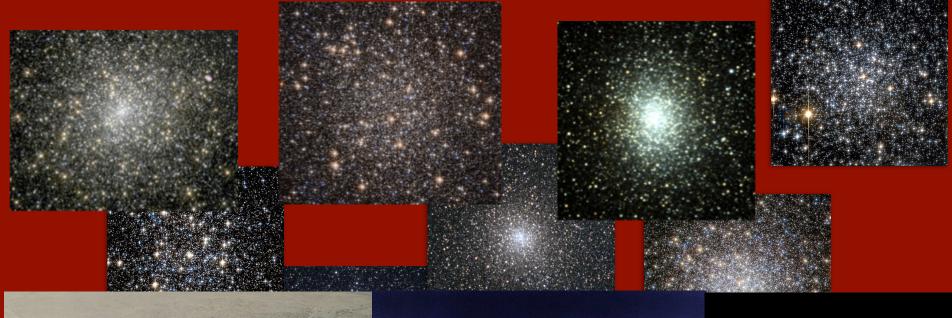


Radial velocities of the red giant companion constrain the mass of the invisible compact object

## Searching for more black holes in globular clusters



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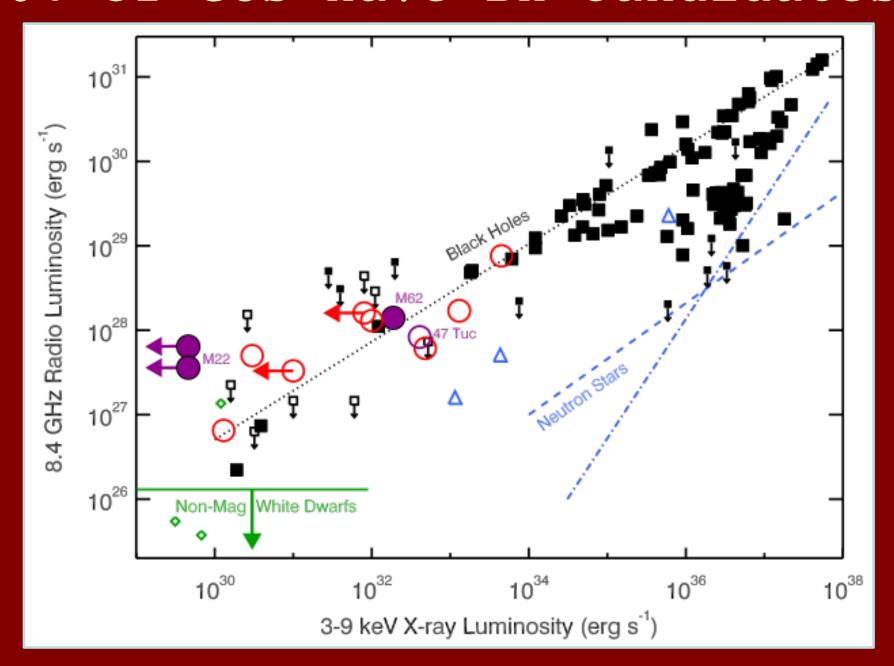




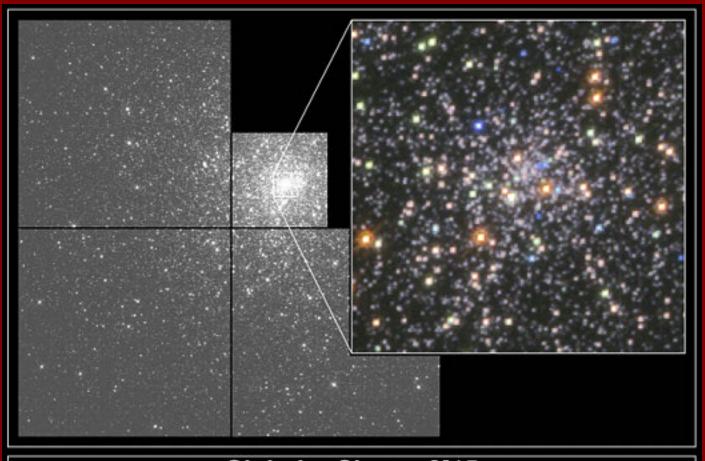


Search 54 clusters for radio emission

### 26% of GCs have BH candidates

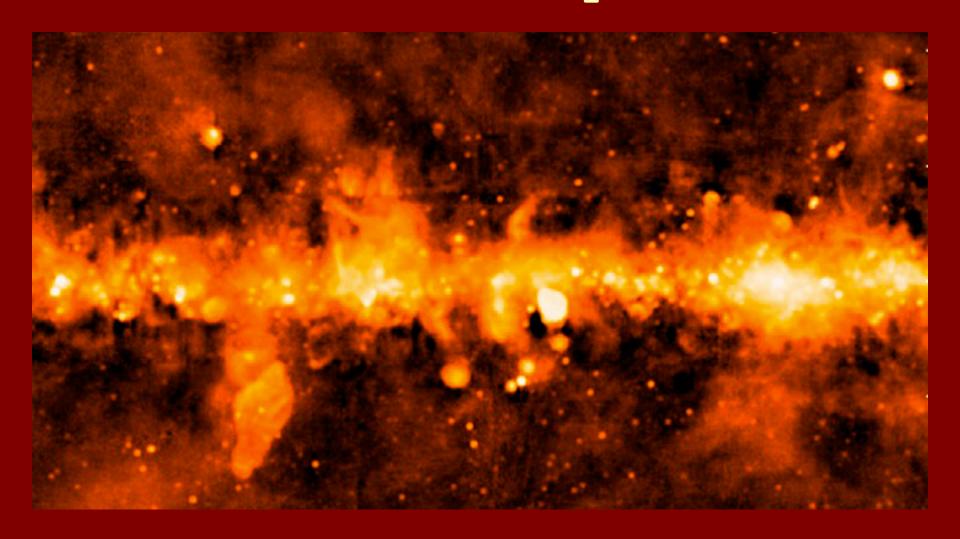


This technique works well in globular clusters, where we focus on small core regions.



Globular Cluster M15
Hubble Space Telescope • Wide Field Planetary Camera 2

### But it can be extended to the Galactic plane.



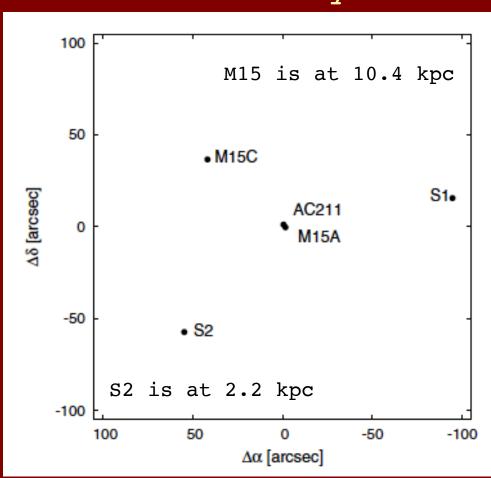
#### BH Searches with NGVLA

- Survey the Galactic plane/bulge---We can detect V404 Cyg out to 10 kpc in just 4 min!
- Wide bands enable focus on flat-spectrum objects
- LSST/WFIRST will enable star/galaxy separation.
- X-ray survey facilities like eROSITA or Lobster will be needed to constrain LR/LX.

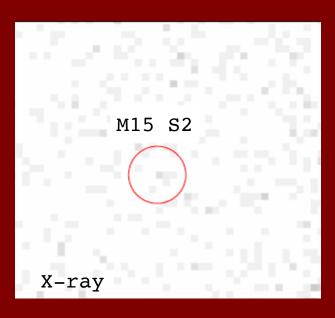
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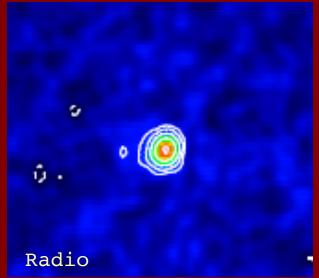
### The curious case of M15 S2

#### Parallax says:



Kirsten et al. 2014





#### Radio bright Galactic objects (BHs?)

