

#### **The VLA Calibration Pipeline**

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# **Pipeline Overview**

- With the start of Jansky VLA Full Operations (January 2013), we started a new operational model:
  - Deliver automatically-flagged and calibrated visibility data
  - You will check the calibration, do additional flagging if needed, then selfcalibrate and image visibility data to meet science goals, using resources at home institution or NRAO computing resources
- Automated pipeline should run correctly on all "standard" continuum Stokes I science SBs, may work for other set-ups as well
- Current versions available:
  - CASA integrated pipeline is compatible with ALMA pipeline infrastructure, improved diagnostics in weblog, used as real-time pipeline since Sep 2015
  - "scripted" pipeline is a collection of python scripts that use CASA tasks wherever possible, but also uses toolkit calls; readable and easy to modify



# **Pipeline Operation**

- Real-time pipeline:
  - Minimal human intervention
    - Pipeline is run automatically on *every* science SB as it completes (not just "continuum")
  - Pipeline output undergoes basic quality assurance checks by NRAO staff, and detailed checks upon request; reports generated are archived as pipeline products
- At your home institution:
  - Instructions for installation and operation of the VLA CASA Calibration
     Pipeline are available at <a href="http://go.nrao.edu/vla-pipe">http://go.nrao.edu/vla-pipe</a>
    - Uses CASA 5.1.1, similar to current real-time pipeline
    - See the VLA CASA pipeline guide at <a href="http://go.nrao.edu/vla-casa-tut">http://go.nrao.edu/vla-casa-tut</a>
- Scripted pipelines for CASA versions through 5.0.0 also available
  - Provides more flexibility in how to use the pipeline, options suitable for spectral line datasets, mixed correlator set-ups, multi-band observations, etc.



# **Pipeline Requirements**

- Will the pipeline work for you?
  - The pipeline successfully completes on >90% of all science SBs observed on the VLA; whether the output can be used for science depends on the science goal, and whether the observation was correctly set up
    - Pipeline includes Hanning smoothing, RFI flagging, and weight calculations that may not be appropriate for spectral line projects (but can modify scripted pipeline)
    - No polarization calibration in default pipeline yet (but see later...)
    - Will probably work well for data taken since May 2012, may work for earlier EVLA data, likely that extra flagging may be needed in these cases
- "Standard" continuum Stokes I science SB means:
  - 128 MHz spws, but may work on other set-ups as well
    - Can work for narrower BWs, depends on the strength of the calibrators
    - Heuristics currently make some assumptions about the strength of the calibrators, in particular, the delay calibrator
  - Contains correctly labeled and complete scan intents
    - And also that the observation has been set up correctly!



# **Pipeline Requirements**

- Calibrator strength:
  - Conservative limit on strength of BP and complex gain calibrators can be derived from requirement for initial gain calibration to work at high end of Q-band
  - Heuristic for delay calibration currently requires the SNR=3 limit on initial gain calibration per integration





# **Pipeline Requirements**

- Correct observation set-up
  - Independent of whether you want to run the pipeline!
  - Remember: simple observing set-ups are always easier to calibrate
  - Do not skimp on calibration to spend more time on your target you may end up not being able to calibrate the target data at all
    - Spending 3 minutes pointing could buy you more sensitivity than doubling the time on your target
- Scan intents
  - The pipeline relies entirely on correct scan intents to be defined in each SB
  - In order for the pipeline to run successfully on an SB it must contain, at minimum, scans with the following intents:
    - A flux density calibrator scan that observes one of the primary calibrators (3C48, 3C138, 3C147, or 3C286) – this will also be used as the delay and bandpass calibrator if no bandpass or delay calibrator is defined
    - Complex gain calibrator scans



# (Real-Time) Heuristics (I)

- Assuming requirements are met, the pipeline:
  - Loads the data
  - Hanning smooths\*
  - Retrieves information about the observing set-up from the data
  - Applies deterministic flags (online flags, shadowed data, end channels of subbands, etc.)
  - Identifies primary calibrators and loads models
  - Derives all prior calibrations (antenna position corrections, gain curves, atmospheric opacity, requantizer gains)
  - Iteratively determines initial delay and bandpass solutions, including running RFLAG (RFI flagging algorithm), and identifying other system (deformatter) problems

\*May want to modify inputs and/or omit entirely for spectral line reductions, depending on level of Gibbs ringing due to RFI or strong emission lines



# (Real-Time) Heuristics (II)

- Heuristics (cont.): the pipeline:
  - Derives initial gain solutions, does flux density bootstrapping and derives spectral index of all calibrators
  - Derives final delay, bandpass, and gain calibrations
  - Applies all calibrations to the MS
  - Runs RFLAG algorithm on all fields, including target\*\*
  - Runs statwt to derive proper relative weights per antenna/spw<sup>\*\*</sup>

\*\*May want to modify inputs and/or omit entirely for spectral line reductions



# Flow chart (CASA pipeline)

Each major step is executed as a separate pipeline task:



#### Flow chart

flag bad deformatters/spws (hifv\_flagbaddef)

run rflag on calibrated delay & BP cals (hifv\_checkflag)

split out calibrators with spectral averaging into calibrators.ms, do a test gaincal to determine short and long solints (hifv\_solint)

make amp gain table for flux density bootstrapping, and do flux density bootstrapping; includes deriving sp. index of calibrators and re-inserting into main ms (hifv\_fluxboot)

make final cal tables (delay, BP, amp, phase) – includes redoing split, fluxboot, etc. (hifv\_finalcals)



#### **Flow chart**





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#### **Observation Overview**

Project	uid://evla/pdb/30673845	Pipeline Version	40896 (Pipeline-CASA51-P2-B)
Principal Investigator	Dr. Tim Bastian	CASA Version	5.1.1-5 r40000
Observation Start	2015-08-02 17:16:40 UTC	Pipeline Start	2017-10-18 16:20:16 UTC
Observation End	2015-08-02 17:52:28 UTC	Execution Duration	1:52:33

#### **Observation Summary**

		Num Antennas	Time (UTC)		Baseline Le				
Measurement Set	Receivers		Start	End	On Source	Min	Max	RMS	Size
Scheduling Block ID: uid://evla/pdbsb/31019491									
Session: session_1									
15A-397.sb31019491.eb31020561.57236.7198700463.ms	20cm (L)	26	2015-08-02 17:16:40	2015-08-02 17:52:27	0:14:56	793.4 m	34.4 km	14.8 km	6.2 GB



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	End Time	2015-08-02 17:52:27		Para 1
	Total Time on Source	0:39:36		1772 <sup>200</sup> 1772 <sup>200</sup> 1772 <sup>200</sup>
	Total Time on Science Target	0:14:56	Intent vs Time	Tield ve Tiere
	LISTOBS OUTPUT		Track scan intent vs time	Track observed field vs time
	Spatial Setup		Spectral Setup	
	Science Targets	'J0842+1835'	All Bands	'20cm (L)'
	Calibrators	'J0542+4951' and 'J0738+1742'	Science Bands	'20cm (L)'
			VLA Bands: Basebands: Freq range: [spws]	L: A0C0: 1.493 GHz to 2.005 GHz: [0,1,2,3] L: B0D0: 993.000 MHz to 1.505 GHz: [4,5,6,7]
	Antenna Setup		Sky Setup	
	Min Baseline	793.4 m	Min Elevation	64.31 degrees



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Observer: Dr. Tim Bastian Project: uid://evla/pdb/30673845	
Observation: EVLA	
Data records: 3088800 Total elapsed time = 2386 seconds	
Observed from 02-Aug-2015/17:16:40.0 to 02-Aug-2015/17:56:26.0 (UTC)	
ObservationID = 0 ArrayID = 0	
Date Timerange (UTC) Scan FldId FieldName nRows SpwIds Average Interval(s) ScanIntent	
02-Aug-2015/17:16:40.0 - 17:25:30.0 1 0 J0542+4951 689000 [0,1,2,3,4,5,6,7] [2, 2, 2, 2, 2, 2, 2, 2] [SYSTEM_CONFIGU	RATION#UNSPECIFIED]
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0 NONE J0542+4951 05:42:36.137916 +49.51.07.23355 J2000 0 995800	
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2	79.22	79.40	79.22	79.23	79.23	79.22	79.22	79.46	79.23	79.24	79.24	79.26	79.40	79.23	79.89	79.28	79.21	79.23	79.23	79.22	79.23	79.23	79.22	79.22	79.72	79.23			
3	79.25	79.94	79.24	79.30	79.40	79.23	79.86	79.92	79.26	79.40	79.52	79.70	79.78	79.25	79.68	79.67	79.22	79.25	79.34	79.24	79.32	79.28	79.24	79.27	80.20	79.26			
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#### **Task Summaries**

Task	QA Score	
1. hifv_importdata: Register VLA measurement sets with the pipeline	No QA	N/A
2. hifv_hanning: VLA Hanning Smoothing	No QA	N/A
3. hifv_flagdata: VLA Deterministic flagging		0.98
4. hifv_vlasetjy: Set calibrator model visibilities	No QA	N/A
5. hifv_priorcals: Priorcals (gaincurves, opacities, antenna positions corrections, rq gains, and switched power)	No QA	N/A
6. hifv_testBPdcals: Initial test calibrations	No QA	N/A
7. hifv_flagbaddef: Flag bad deformatters	No QA	N/A
8. hifv_checkflag: Flag possible RFI using rflag and tfcrop	No QA	N/A
9. hifv_semiFinalBPdcals: Semi-final delay and bandpass calibrations	No QA	N/A
10. hifv_checkflag: Flag possible RFI using rflag and tfcrop	No QA	N/A
11. hifv_semiFinalBPdcals: Semi-final delay and bandpass calibrations	No QA	N/A
12. hifv_solint: Determine solint and Test gain calibrations	No QA	N/A
13. hifv_fluxboot: Gain table for flux density bootstrapping	No QA	N/A
14. htfv. finalcals: Final Calibration Tables. e:///Users/cchandle/talks/2017/DRWkshp/pipeline-20171018T162015/html/t1-4.html	No QA	N/A



- The following pipeline steps provide key checks for calibration quality:
  - hifv\_flagdata deterministic flagged data fraction
  - hifv\_priorcals
     gain compression due to RFI (or the Sun)
  - hifv\_solint solution intervals used for short/long phase cals
  - hifv\_fluxboot
  - hifv\_finalcals
- fitted calibrator flux densities and spectral indices final calibration tables to be applied to the data
- hifv\_plotsummary some useful diagnostic plots of calibrated data
- If something funny shows up in these steps you can look at the intermediate steps to see what might have gone wrong



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20. hif_makeimages	Files used for template flagging steps.											
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Most Visited -

1. hifv\_importdata

2. hifv\_hanning

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4. hifv\_vlasetjy

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16. hifv\_targetflag

19. hif\_makeimlist

20. hif\_makeimages

17. hifv\_statwt 18. hifv\_plotsummary

Tasks in execution order

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A Home By Topic By Task

#### Other Before Shadowed Online Edge Total Unwanted Flagging Data Selection (by intent) Task ANOS **Antennas** Intents Flags Template Autocorr Channels Clipping Quack Baseband Science All Data 11.00% 3.12% 19.83% 0.00% 18.02% 4.86% 0.00% 0.00% 6.14% 0.00% 0.00% 0.00% Science Spectral Windows 3.12% 19.83% 0.00% 18.02% 4.86% 0.00% 0.00% 6.14% 0.00% 0.00% 0.00% 11.00% Bandpass 3.12% 1.09% 0.00% 0.00% 0.17% 0.00% 0.00% 6.44% 0.00% 0.00% 0.00% 6.62% Flux 3.12% 1.09% 0.00% 0.00% 0.17% 0.00% 0.00% 6.44% 0.00% 0.00% 0.00% 6.62% Phase 3.12% 41.29% 0.00% 0.00% 0.00% 0.00% 0.00% 6.45% 0.00% 0.00% 0.00% 6.45% Target 3.12% 6.12% 0.00% 0.00% 5.65% 0.00% 0.00% 6.09% 0.00% 0.00% 0.00% 11.73% 11. hifv\_semiFinalBPdcals 15A-397.sb31019491.eb31020561.57236.7198700463.ms 3.12% 19.83% 0.00% 18.02% 4.86% 0.00% 0.00% 6.14% 0.00% 0.00% 0.00% 11.00% Summary of flagged data. Each cell states the amount of data flagged as a fraction of the specified data selection, with the Flagging Agent columns giving this information per flagging agent.

#### Flagging reason vs time

Plots of flagging reason vs time. The reasons for flagging the data are defined in the plot legend

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. hifv_testBPdcals	15A-397.sb31019491.eb31020561.57236.7198	700463.ms										
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10. hifv_checkflag	ea17&&*	
11. hifv_semiFinalBPdcals	ea16&&*	
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Summary of gencal opacities

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#### Antenna positions

No antenna position corrections to apply.

#### Requantizer gains

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#### Switched Power plots

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Switched Pow r Plots: SwPower SPgain plots | SwPower Tsys plots





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#### Flux Density Bootstrapping (hifv\_fluxboot)





#### Flux Density Bootstrapping (hifv\_fluxboot)

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### Final Calibration Tables (hifv\_finalcals)

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## Final Calibration Tables (hifv\_finalcals)





### Final Calibration Tables (hifv\_finalcals)





















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# Calibrator Images (hif\_makeimages)

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# Calibrator Images (hif\_makeimages)





# **Pipeline Products and Outputs**

- Real-time pipeline products and output
  - Flag and calibration tables
  - Calibrated and flagged MS (available for 15 days, not archived)
  - Logs, including weblog used by quality assurance (QA) staff and QA report if requested
  - Products can be requested through the helpdesk (<u>https://help.nrao.edu</u>, VLA Pipeline Department), for download via FTP or shipping on hard disk(s)
- You may request a detailed QA report from the data analysts
- If you are happy with the pipeline calibration, then:
  - Do further flagging if necessary
  - Split out your target and image



# **Pipeline Products and Outputs**

- If you have the SDM or uncalibrated MS and the calibration and flag tables, instructions for applying flags and calibration tables may be found at <u>http://go.nrao.edu/vla-pipe</u>; see Drew's talk!
- In some cases the pipeline and/or the MS may need to be modified
  - Download the SDM from the archive plus pipeline scripts
  - Follow the directions at above link
- In some cases the pipeline heuristics may not be appropriate for your data (e.g., some L-band set-ups do not work well with the pipeline yet)
  - Reduce data by hand



# **Imaging comparison**



Left: L-band image of G55.7+3.4 produced from data flagged and calibrated by hand; the rms noise is 11.5  $\mu$ Jy/beam. Right: an image made from data flagged and calibrated by the VLA calibration pipeline; the rms noise is 12.2  $\mu$ Jy/beam. Differences in the source structure and/or source flux density are dominated by the uncertainty in the deconvolution process, not the calibration and flagging (images provided by Urvashi Rao).



#### Known failure modes and issues

- In general the pipeline does very well, but there are possible failure modes:
  - No flux density or gain calibrator intents defined, or flux density calibrator not one for which we have models
    - work around in scripted pipeline
  - Wrong scan intents
    - work around in scripted pipeline
  - Does not always identify deformatter problems (but does NOT usually have false positives – L-band may be an exception)
    - flag remaining bad spws
  - Calibrators are too weak for given spw bandwidth
    - heuristics have been developed for the case of a weak bandpass calibrator, you are welcome to try them on your data!



### **Spectral line data**

- Several steps in the real-time pipeline may not be appropriate for spectral line data:
  - Hanning smoothing (increases effective channel width)
  - Last run of RFLAG on target (may eliminate your line as interference!)
  - Statwt calculates rms based on scatter of channels per spw, per visibility; may want to run manually with channel selection turned on to eliminate use of channels containing line emission in calculating the rms
- With the above modifications, the pipeline will work with spectral line data as long as the calibrators are strong enough



#### **Mixed correlator set-ups**

- With the new WIDAR capabilities it is common to observe both wide and narrow spws to obtain both continuum and spectral line data simultaneously, or multiple receiver bands
  - A single heuristic (e.g., gain calibration solution interval) for entire dataset may not be appropriate
- Solution:
  - Run pipeline through application of deterministic flags, including Hanning smoothing if you are going to use it
  - Split the MS by spw and/or scans
  - Run pipeline on split MSs WITHOUT Hanning smoothing (you have already applied it, if you are going to use it)
  - Warning: output flagging statistics may not be correct



# **Special cases**

- Incorrect scan intents
  - Best to use the scripted pipeline (otherwise have to edit SDM)
  - Can run through msinfo.py, then re-set the following string variables to refer to the correct scan and field IDs:

flux\_field\_select\_string='2'
bandpass\_scan\_select\_string='8'
bandpass\_field\_select\_string='4'
delay\_scan\_select\_string='8'
delay\_field\_select\_string='4'
calibrator\_scan\_select\_string='4,5,7,8,10,11,12'
calibrator\_field\_select\_string='1,2,3,4,5,6,7'
phase\_scan\_list=[1,3,5,7,9,11,13,15]

 If a standard flux density calibrator was not observed, you may still be able to use the pipeline IF you know the flux density and spectral index of one of your other calibrators, with a bit more work – contact the NRAO <u>h</u>elpdesk



# **Special cases**

- Accurate flux density bootstrapping
  - hify fluxboot uses medians to bootstrap flux densities: fairly robust, but in some cases (e.g., high frequencies with pointing, elevation dependent gains) you can do better by flagging the gain table used for the bootstrapping



- Run pipeline through hifv\_fluxboot
- In a separate instance of CASA, flag input gain table using "plotcal" or "plotms"
- Re-run hifv\_fluxboot specifying name of input gain table, then continue with the rest of the pipeline
- With care (match elevation of flux cal, flag bad data), can reproduce flux



density scale to a few % at Q-band

# Summary and VLASS Developments

- Heuristics for Stokes I continuum now well-tested, stable, minor modifications allow the pipeline to be used for certain spectroscopy projects as well
- Heuristics available for user-testing:
  - Weak bandpass calibrators
- Development for the VLA Sky Survey available for user-testing:
  - Improved RFI flagging
    - Uses cross-hands to help identify and flag RFI
  - More robust flux density bootstrapping (single-band only)
    - Flags input data to remove bad solutions before fitting
  - Polarization calibration
    - Makes some assumptions about calibrator strengths
    - Supports unpolarized and unknown polarization leakage calibrators
- [VLASS imaging pipeline, specific to VLASS observing mode]



















### Additional polcal plotsummary plots





#### **Future Developments**

- Heuristics being developed for:
  - Multi-band delays
  - Gain compression
  - Low-frequencies (<IGHz)</li>
  - Self-calibration and imaging for PI science
- Heuristics developed in consultation with expert users and staff; testing, feedback, suggestions welcome!

For how to run the pipeline see Drew's talk next! Or go to <u>http://go.nrao.edu/vla-pipe</u>

