

VLA and VLBA Proposal Preparation Emmanuel Momjian







Creating NRAO/LBO/GBO proposals

- Proposal Submission Tool (PST)
 - VLA, VLBA, GBT, but ...
 - for ALMA, use ALMA OT
- Many elements in the PST are common to the VLA, VLBA, and the GBT. However, differences will be noted as needed.
- Accessing the PST
 - You must be registered at my.nrao.edu





Creating an NRAO/LBO/GBO proposal

Register and login at my.nrao.edu





Register for an Account
Accounts are used by astronomers to create and submit proposals, prepare for observations, and gain access to proprietary data from the archive.
REGISTER
<u>Lost your Username or Password?</u> You can reset it <u>online.</u>
If you need help, please <u>email us</u> .

Staff | Policies | Diversity





The National Radio Astronomy Observatory is a facility of the National Science Foundation operated under cooperative agreement by Associated Universities, Inc.





my.nrao.edu

- Gives access to various services:
 - Proposal preparation
 - Allows creating and submitting new proposals
 - Gives access to all proposals you are associated with regardless of your role (PI, co-I, contact author, reviewer)
 - Data Processing
 - VLA Observation Preparation Tool (OPT)
 - The VLA, VLBA, GBT data archive





my.nrao.edu



Dashboard

Proposals

Data Processing

Obs Prep

Helpdesk

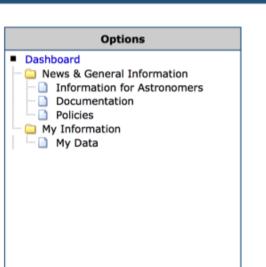
Profile

Hi, Galactico

Monday 03 July 2017

Help

Sign Out



DASHBOARD

Web Browser

We recommend using the Firefox web browser for the Proposal Submission Tool (PST).

Warning: there are problems uploading files to the PST with Safari.

User Accounts

Please remember to update your user profile, especially if you have moved to a new institution. Do not create a new account.

Telescope News

Next Proposal Deadline: August 01, 2017 5 PM EDT (21 hours UT) in 29 days

Proposal Finder Tool - Search cover sheets of approved NRAO telescope proposals.

Important All proposal authors must be registered users

Important Information for VLA/GBT/VLBA/HSA/VLBI Proposers

VLA Configuration Plans and Proposal Deadlines

Proposal Help Desk





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Timeline: Proposal Deadlines

- Two per year: Typically February I and August I
- Next deadline: August 1st for 2019A (i.e., first half of 2019).
- The VLA configurations in 2019A are
 - A (most extended, B_{max} =36.4 km), and
 - B (2nd most extended, B_{max} =11.1 km).
- Note: Director's Discretionary Time proposals are not subject to these deadlines (more on this later).





Timeline: After deadline

- Proposals are evaluated by the Science Review Panels (SRPs) and the Time Allocation Committee (TAC)
- Observers are informed of allocated time (if any) and scientific priority (A, B or C) in a "disposition letter" about a month before next call.
- Proposal scheduling priorities:
 - A: Highest priority, most likely to be observed;
 - B: Next highest priority, scheduled on best effort basis;
 - C: Filler time
- For VLA: Schedules can be submitted about a month before configuration.
 - If needed, schedules can be made in advance as a Test project in the OPT to seek advise/validation.
- For VLBA: Schedules can be submitted once disposition letter goes out.







Types of proposals

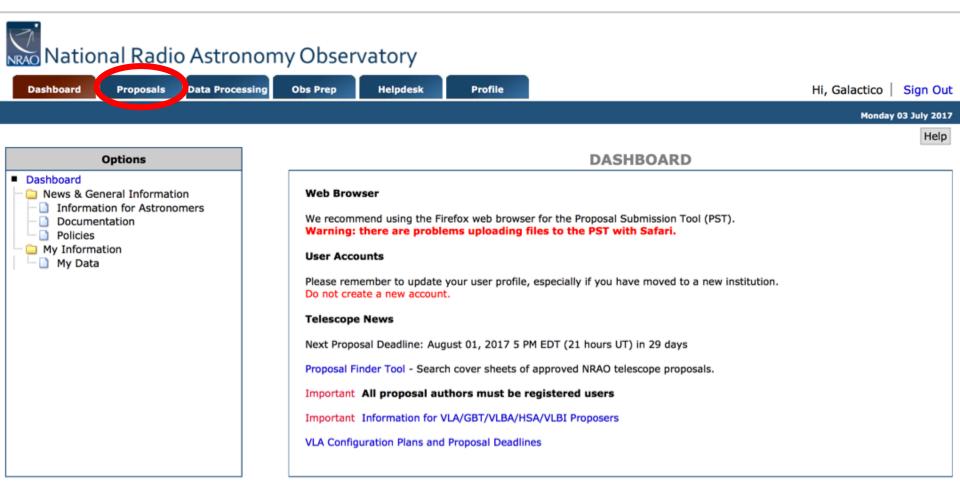
- Proposals submitted at deadlines:
 - Regular (< 200h) ≤ 4 pages science justification
 - Large (≥ 200h) ≤ 10 pages science justification; requires data reduction and release plan
 - Triggered ≤ 4 pages science justification
 - pre-planned observations of transients whose event times are unknown a priori; well-defined triggering criteria are required
- Director's Discretionary Time
 - Not tied to proposal deadline, limited time request
 - For a Target of Opportunity (unexpected, unpredicted, e.g. supernova in nearby galaxy) or
 - Exploratory Time for high risk/high yield or last minute projects
 - Must have a good reason for why this was not proposed at a regular deadline.







Log into my.nrao.edu and go to "Proposals"









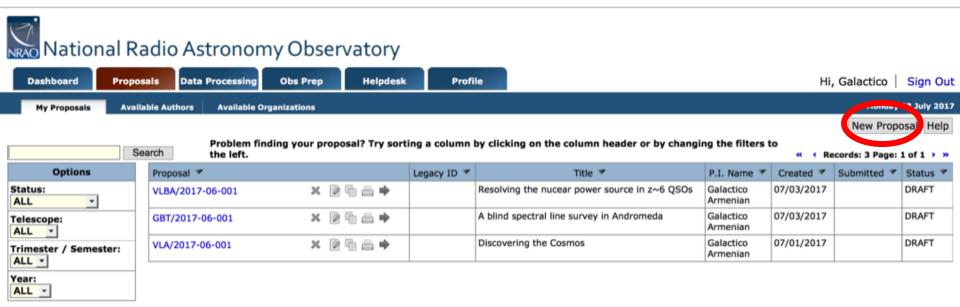
The National Radio Astronomy Observatory is a facility of the National Science Foundation operated under cooperative agreement by Associated Universities, Inc.







Click on "New Proposal"







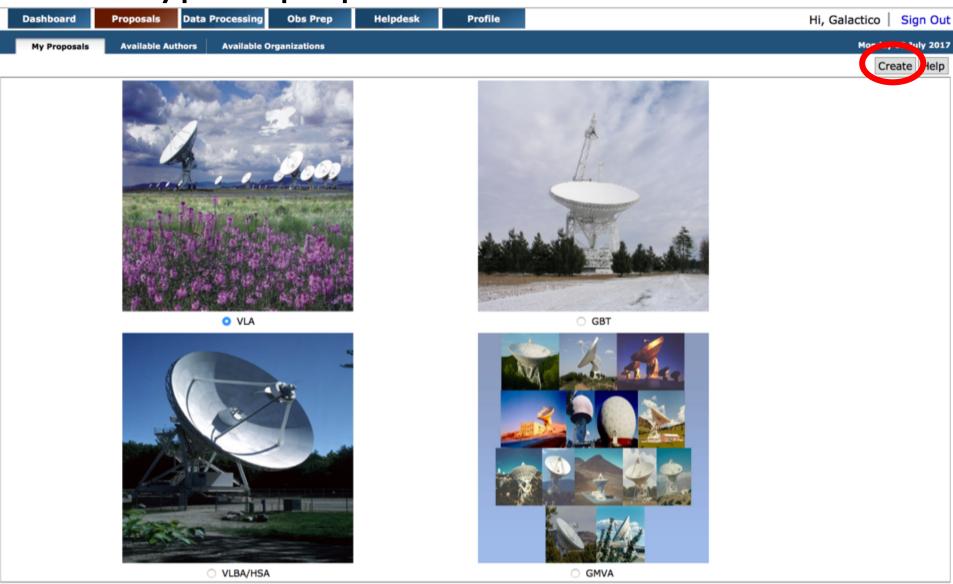
The National Radio Astronomy Observatory is a facility of the National Science Foundation operated under cooperative agreement by Associated Universities, Inc.



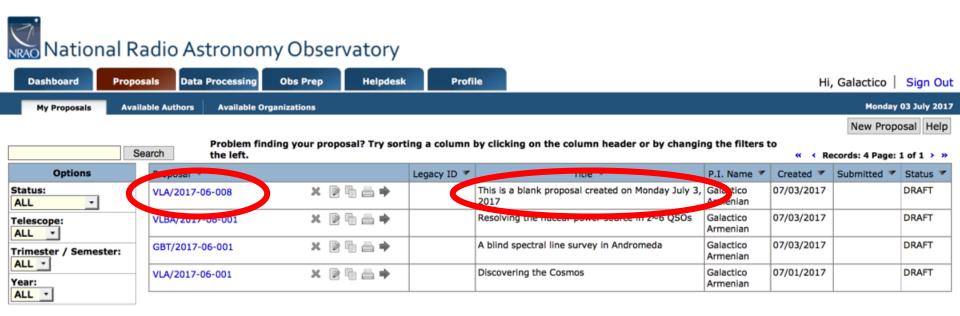




Select type of proposal and then "Create"



"Blank" proposal now appears in "My Proposals"



Click on the name of the proposal to see its subsections and edit them.



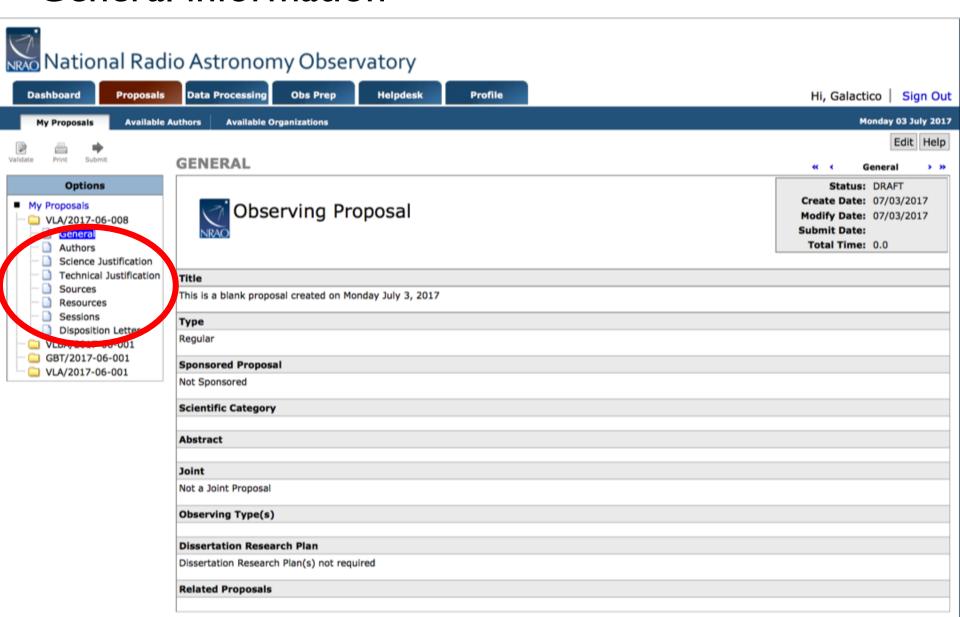


The National Radio Astronomy Observatory is a facility of the National Science Foundation operated under cooperative agreement by Associated Universities, Inc.





General information



Major Elements of a Proposal

- General (Title, Type, Abstract, etc....)
- Authors
- Science Justification
- Technical Justification
- Sources what do you want to observe
- Resources instrumental setup
- Sessions which of your sources do you want to observe with which of your resources
- Disposition Letter (will contain a copy of the disposition letter after the proposal review process)

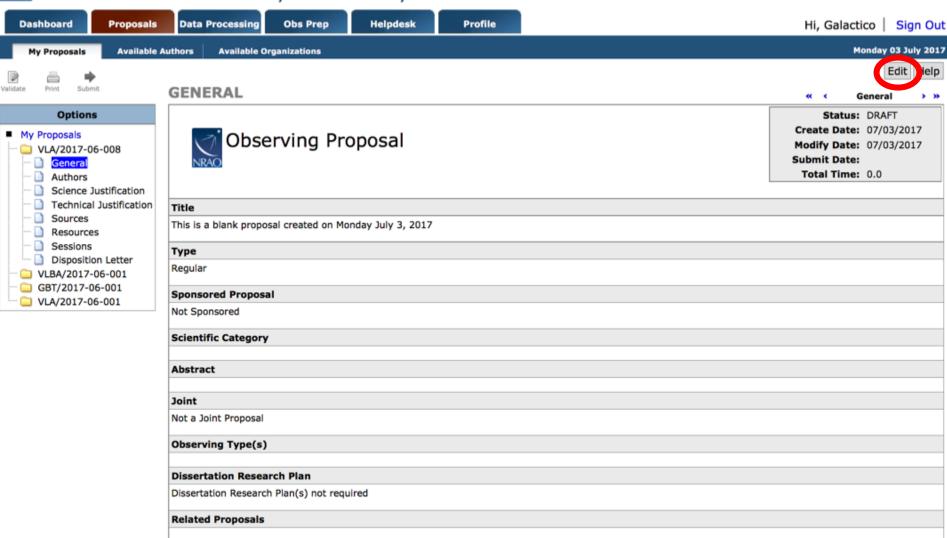




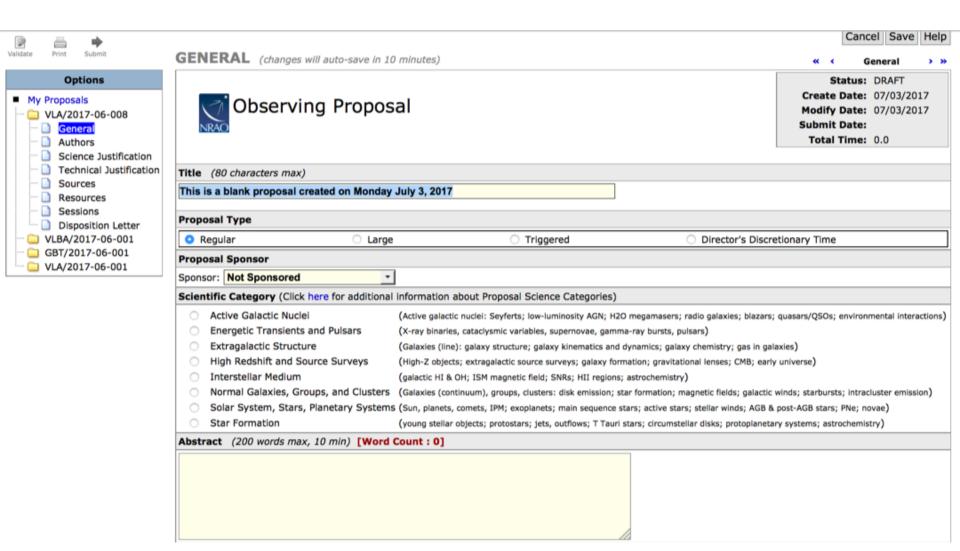


Click "Edit"





General section









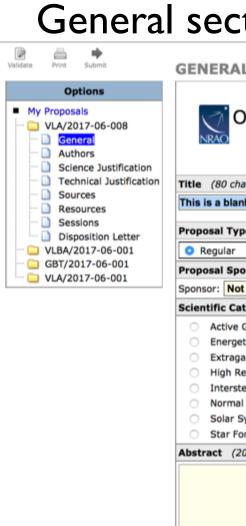
General section

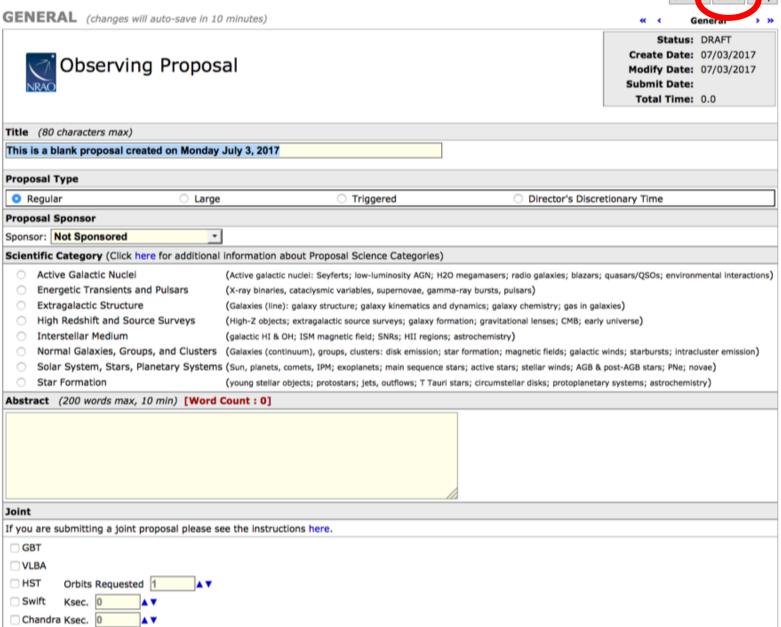
Joint				
If you are submitting a joint proposal please see the instructions here.				
□ GBT				
□VLBA				
☐ HST Orbits Requested 1	▲ ▼			
☐ Swift Ksec. 0	,			
☐ Chandra Ksec. 0	•			
Observing Type(s)				
☐ Continuum ☐ S	Spectroscopy	□ Polarimetry		
☐ Single Pointing(s) ☐ (Grid Mapping/Mosaicing	☐ OTF Mapping		
☐ Sun ☐ N	Monitoring	☐ Solar System		
☐ High Time Resolution ☐ F	Pulsar	Radar		
☐ Geodesy ☐ A	Astrometry	□ VLA Subarrays		
□ Other				
Dissertation Research Plan				
Dissertation Plan is now associated with author(s) and must be set on the Author's page. Dissertation Plan section will appear when a student author is marked "Observing For Thesis".				
Related Proposals				





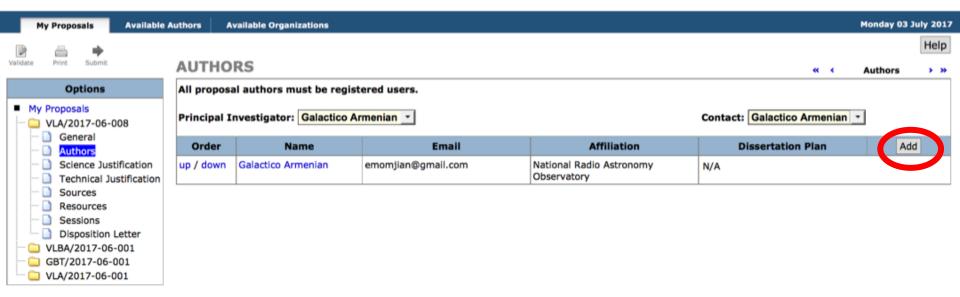
General section





Cano

Authors Section



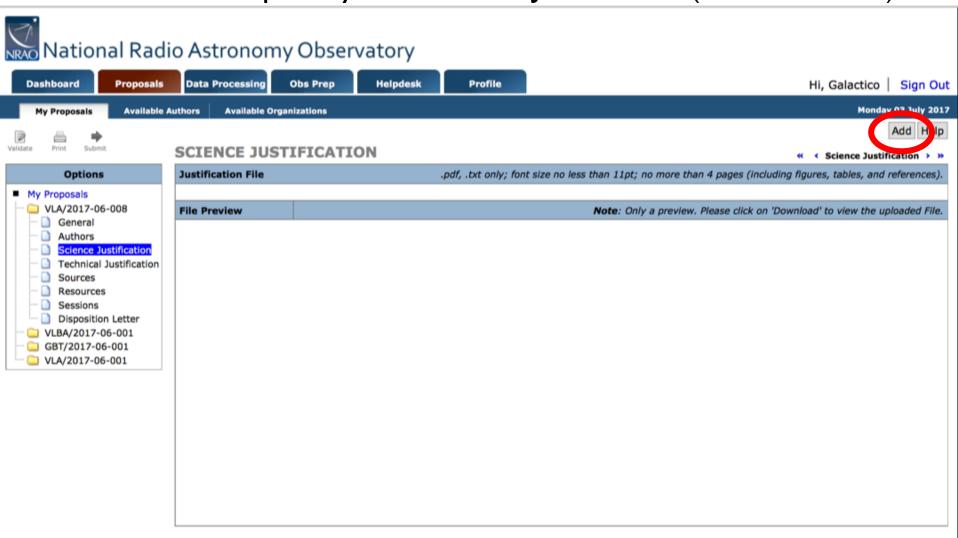
- Will automatically have you as PI and Contact author.
- Add authors with Add button
 - Co-authors need to be in the my.nrao.edu database.
- Move authors up and down on list by using up/down
- Can reassign Pl and Contact author





Scientific Justification Section

Click "Add" to upload your Scientific Justification (PDF or txt file)





Sources Section



SOURCES

Proposers must specify their source lists (or potential targets) in full with the exception of Triggered proposals where the targets are unknown a priori.
NOTE: If you update a source group after it has been attached to a session, you may have to un-attach and re-attach the source group for the changes to take effect.

Order Name Position Velocity

Sources can be added in three ways

- Creating "New Source Group" then:
 - add manually
 - search in NED/SIMBAD
- Copy from old proposal ("Copy Sources")
- Load from local data file ("Import")





Sources Section



Once you create a group, you can add sources manually or by searching.



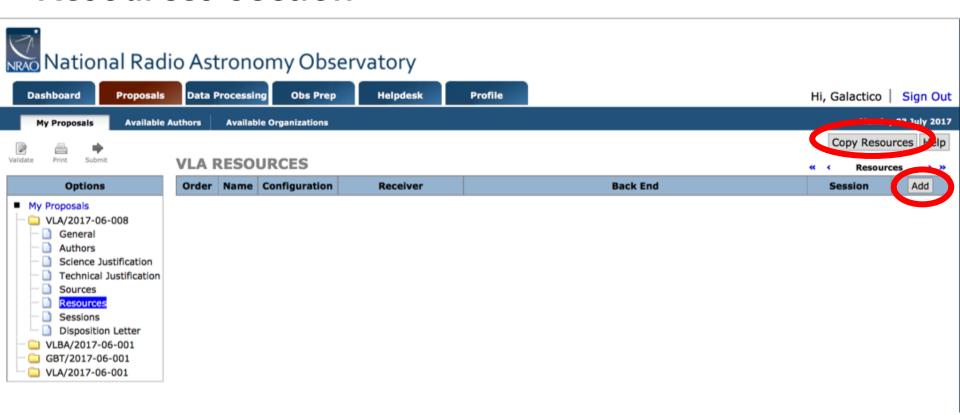


Sources Section

Order	Name		Position			Velocity	Add
		Coordinate System		Equatorial	Convention	Optical	
		Equinox		J2000			
		Right	Value:	16:42:58.80	Ref. Frame	Barycentric	
p / down	3C345	Acsension	Range(±):	00:00:00.0			
		Declination	Value:	+39:48:36.9	Redshift	0.59280	
		Decimation	Range(±):	00:00:00.0			
		Calibrator		N			
		Coordinate System		Equatorial	Convention	Optical	
		Equinox		J2000			
		Right	Value:	12:18:57.50	Ref. Frame	Barycentric	
up / down NGC4258	Acsension	Range(±):	0.00:00:00				
	Declination	Value:	+47:18:14.3	Redshift	0.001541		
	Deciliation	Range(±):	00:00:00.0				
		Calibrator		N			





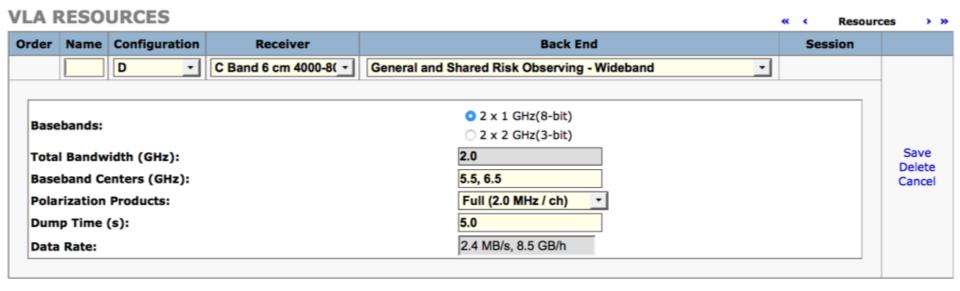


- Click on "Copy Resources" if you want to copy from another proposal, or
- Click on "Add".



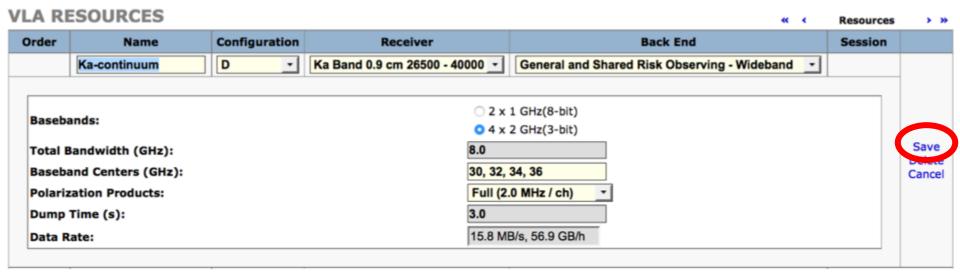


Adding a resource for continuum science:





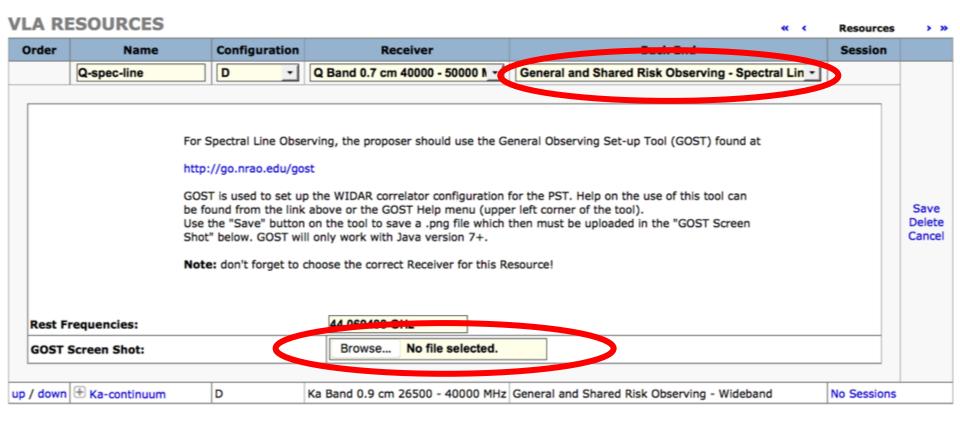








Adding a resource for spectral-line science:

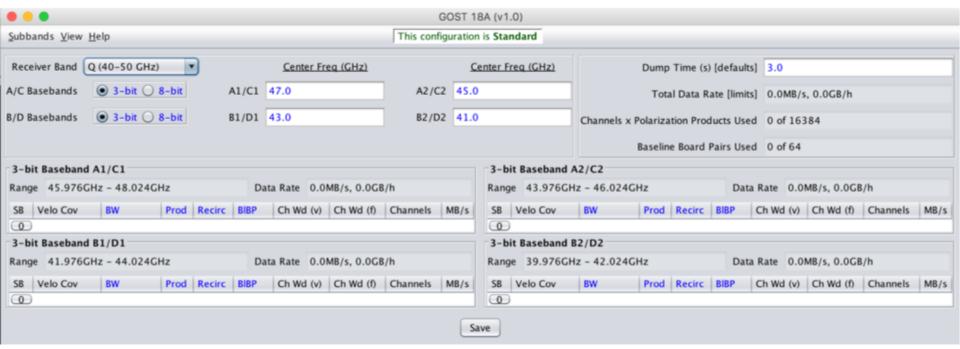






GOST (General Observing Setup Tool)

go.nrao.edu/gost



- click Subbands → Template... then adjust subband size, polarizations desired, recirculation and baseband pair stacking (BIBP) – recirculation and BIBP can be used to increase spectral resolution
- click Subbands → Fill 16 Subbands (or Fill 32 Subbands for 3 bit) → All A/C (or All B/D...)
- Refer to GOST section of "VLA Proposing Guide" at go.nrao.edu/vla-prop, especially "GOST Usage Hints"





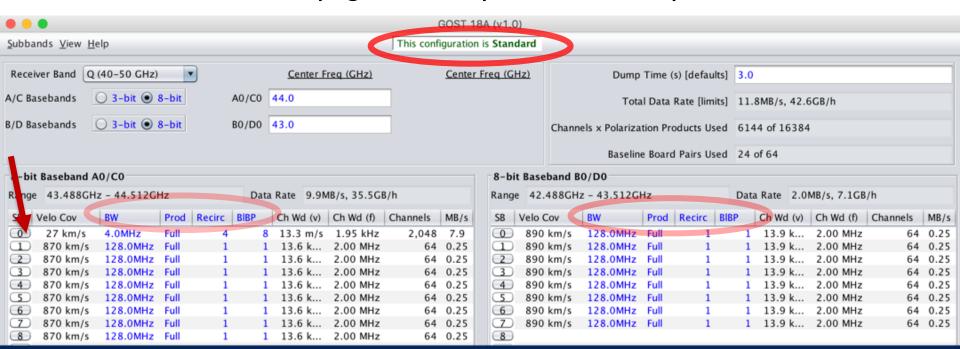


GOST (General Observing Setup Tool)

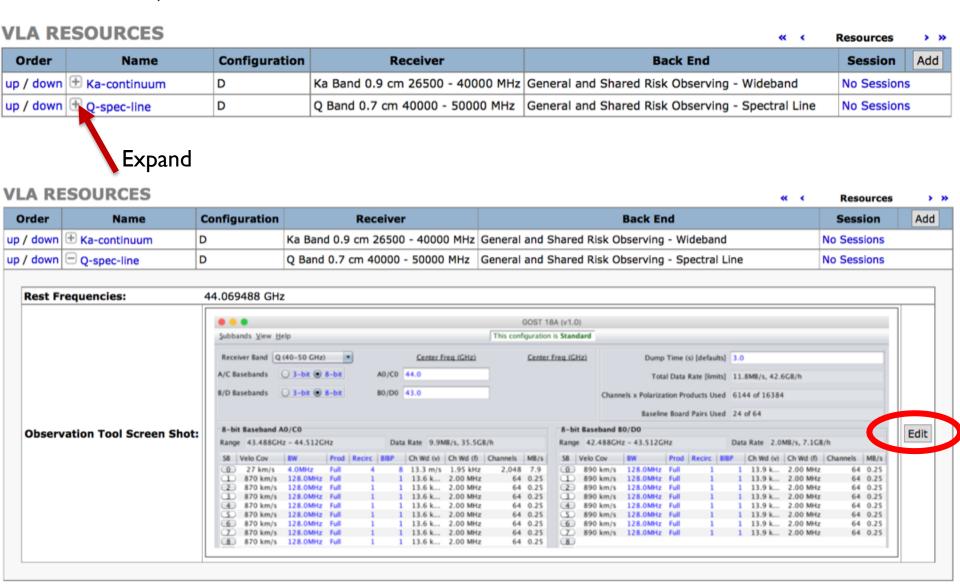
go.nrao.edu/gost

For the spectral line of interest at Q-band, using:

- 8-bit samplers
- 4 MHz subband with 2048 channels (use Recirc and BIBP)
- Full polarization products
- May add wide-band subbands for continuum (based on the science)
- Note: blue color fields give drop-down menus.
- Once done, save as a 'png' file and upload it to the spectral line resource



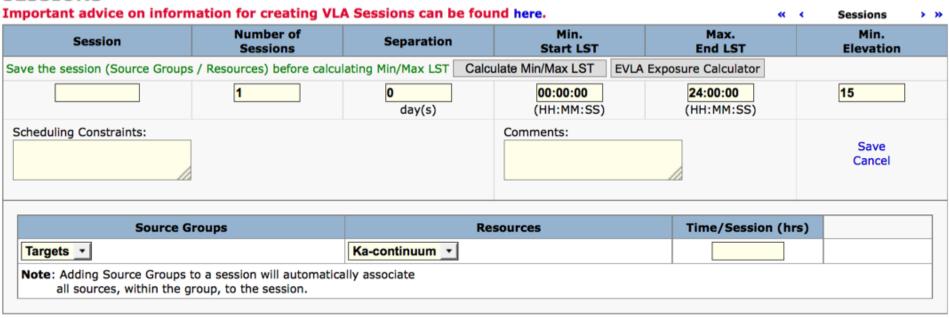
Once done, it will look like:



Sessions-- connecting the Sources and Resources

...and some other details

SESSIONS



- Click "New Session" on the top right
- Enter name and number and separation of epochs, and LST range on top line
- Click Add
 - Select a source group and a resource
 - Enter time per session





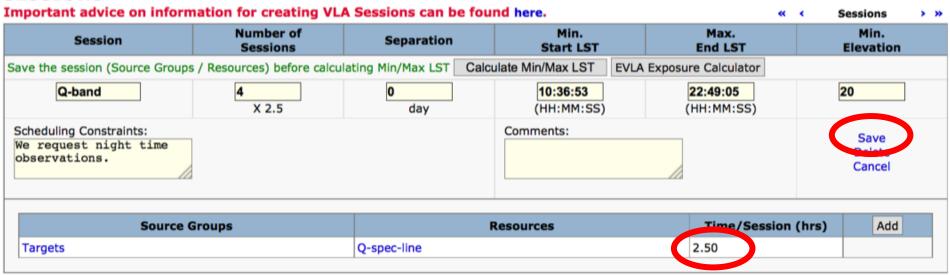
New Session

Help



Sessions

SESSIONS



Note that this number, and the number of sessions are based on the sensitivity calculations and overhead assumptions. Stay tuned....





Technical Justification — I

See 'Guide to Proposing for the VLA' for an example: go.nrao.edu/vla-prop

6	Save H II
TECHNICAL JUSTIFICATION	«
VLA Technical Justification	
Use this page to specify how the technical set-up requested for your proposal enables the scientific goals to relevant for your proposal then enter "NA" into the textbox. The links within each box provide information of	
Are the data to be combined with those from other configurations or radio telescopes, if so, please specify: http://go.nrao.edu/combine	
Explain the reason for the array configuration(s) requested. Include the angular extent of the source and the largest angular size (LAS) to be measured:	
http://go.nrao.edu/vla-res Give possible scheduling constraints. Issues that should be addressed: 1. Are targets nighttime/daytime for the configurations proposed (possibly important for low-frequency interference or high frequency phase stability)? 2. What will be the target elevation (possibly important for high-frequency calibration and overhead)? 3. What is the required date for coordinated or fixed-date observations? 4. Are there dates that should be excluded, e.g. for proximity to the Sun? 5. For Large projects, what is the total number of passes required at a given LST? http://go.nrao.edu/vla-plan	
If you entered a minimum / maximum LST different from the calculated value for any session, you must justify the values here:	
Explain choice of receiver(s) requested: http://go.nrao.edu/vla-frq	
Describe the choice of samplers and the correlator set-up(s) requested. For spectral line observations also provide details such as: - Velocity/frequency span of the line(s). - Velocity/frequency resolution needed. - Subband width(s) and channel numbers to be used for each line. - Rest and sky frequencies of interest. http://go.nrao.edu/vla-samplers	
map () gotting occupy the partiplets	

http://go.nrao.edu/vla-obsline

Technical Justification — I

go.nrao.edu/vla-res

Configuration	Α	В	С	D	Α	В	С	D
B _{max} (km ¹)	36.4	11.1	3.4	1.03	36.4	11.1	3.4	1.03
B _{min} (km ¹)	0.68	0.21	0.035 ⁵	0.035	0.68	0.21	0.035 ⁵	0.035
Band	Synthesiz	ed Beamv	vidth O _{HPBW} ((arcsec) ^{1,2,3}	Largest Angular Scale θ _{LAS} (arcsec) ^{1,4}			
74 MHz (4)	24	80	260	850	800	2200	20000	20000
350 MHz (P)	5.6	18.5	60	200	155	515	4150	4150
1.5 GHz (L)	1.3	4.3	14	46	36	120	970	970
3.0 GHz (S)	0.65	2.1	7.0	23	18	58	490	490
6.0 GHz (C)	0.33	1.0	3.5	12	8.9	29	240	240
10 GHz (X)	0.20	0.60	2.1	7.2	5.3	17	145	145
15 GHz (Ku)	0.13	0.42	1.4	4.6	3.6	12	97	97
22 GHz (K)	0.089	0.28	0.95	3.1	2.4	7.9	66	66
33 GHz (Ka)	0.059	0.19	0.63	2.1	1.6	5.3	44	44
45 GHz (Q)	0.043	0.14	0.47	1.5	1.2	3.9	32	32





Technical Justification – I

Save	Heln
Save	Help

TECHNICAL JUSTIFICATION	
VLA Technical Justification	
Use this page to specify how the technical set-up requested for your proposal enables the scientific goals to relevant for your proposal then enter "NA" into the textbox. The links within each box provide information of	
Are the data to be combined with those from other configurations or radio telescopes, if so, please specify: http://go.nrao.edu/combine	
Explain the reason for the array configuration(s) requested. Include the angular extent of the source and the largest angular size (LAS) to be measured: http://go.nrao.edu/vla-res	
Give possible scheduling constraints. Issues that should be addressed: 1. Are targets nighttime/daytime for the configurations proposed (possibly important for low-frequency interference or high frequency phase stability)? 2. What will be the target elevation (possibly important for high-frequency calibration and overhead)? 3. What is the required date for coordinated or fixed-date observations? 4. Are there dates that should be excluded, e.g. for proximity to the Sun? 5. For Large projects, what is the total number of passes required at a given LST? http://go.nrao.edu/vla-plan	
If you entered a minimum / maximum LST different from the calculated value for any session, you must justify the values here:	
Explain choice of receiver(s) requested: http://go.nrao.edu/vla-frq	
Describe the choice of samplers and the correlator set-up(s) requested. For spectral line observations also provide details such as: - Velocity/frequency span of the line(s). - Velocity/frequency resolution needed. - Subband width(s) and channel numbers to be used for each line. - Rest and sky frequencies of interest. http://go.nrao.edu/vla-samplers	

http://go.nrao.edu/vla-obsline

Technical Justification — I

- VLA has two separate sampler sets:
 - Two 8-bit sampler pairs, each pair covering 1024 MHz
 - Total of 2048 MHz BW, per polarization
 - Four 3-bit sampler pairs, each pair covering 2048 MHz.
 - Total of 8192 MHz BW, per polarization
- 8-bit paths primarily for low frequency bands (P, L, S)
 - 90cm, 20cm, 10cm bands
- 3-bit paths primarily for high frequency bands (C through Q)
 - 5cm, 3cm, 2cm, 1.3cm, 0.9cm, 0.7cm
- But: 3-bit samplers lose ~ 15% sensitivity.
 - If your science requires < 2 GHz BW, use the 8-bit samplers.







Technical Justification - II

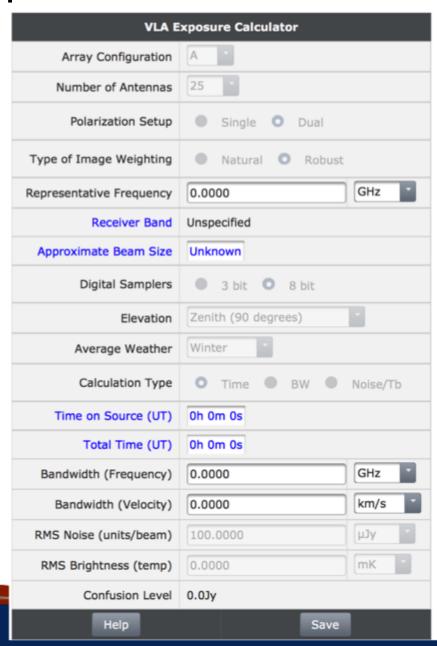
Give the sensitivity required to achieve the science goal; include frequency or velocity width assumed: http://go.nrao.edu/vla-rms					
ive the required on-source integration time to achieve the required sensitivity, and total time including overhead; include considerations such as source infusion in compact configurations, RFI in the geostationary satellite belt, self-noise for strong sources; if the overhead assumed is different from that given the exposure calculator, please explain:					
Please upload exposure calculator graphic(s). Multiple files should be uploaded if there are multiple resources. Use the "Save" button on the tool to save a pdf file which can then be uploaded using the browse/upload buttons to the right. http://go.nrao.edu/ect	Browse No file selected.			. [Upload
	File Name	Size			
	File Name	Size	delete	download	
Note correlator dump time, data rate, and total volume of all raw dat expected (not just the on-source fraction); for data rates in excess of 25 MB/s, please provide additional justification for why this data rate is required (for simple experiments, the data rates are calculated for each correlator setup in the GOST tool and for wide-band observations, the PST gives the rate when a Resource is set up): http://go.nrao.edu/tim-res					
Use this space to tell the technical reviewer what expected imaging problems you might expect to see, due, e.g. to wide fractional bandwidths, ionosphere, nearby strong sources, complex source structure, etc. Please also let us know how you plan to ameliorate these imaging problems. This might include using particular kinds of software and computing resources, either at NRAO or your home institution. Other information that might be useful to the reviewer are whether the target can be self-calibrated, whether or not the images will be dynamic range limited, etc. http://go.nrao.edu/imaging					
For polarimetric observations, note whether the observations require parallactic angle coverage, or whether an unpolarized source will be used to calibrate determine the D-terms:					
http://go.nrao.edu/vla-pol					
Note any potential problems with RFI in the proposed observations. Proximity to the geosynchronous satellite belt in the declination range from about 0 to -10 degrees should be noted.					
http://go.nrao.edu/vla-rfi					
If this is a joint external proposal (e.g., HST, Chandra, or Swift), please add any technical details about the external telescope here:					
Note any other special technical considerations with either the setup or the data processing. RSRO proposals should use this section to describe who will fill the residency requirements for the proposal, along with a description of their technical expertise.					
http://go.nrao.edu/vla-oss					
http://go.nrao.edu/vla-capabilities					
If you entered a minimum / maximum LST different from the calculated value for any session, you must justify the values here:					







VLA Exposure Calculator



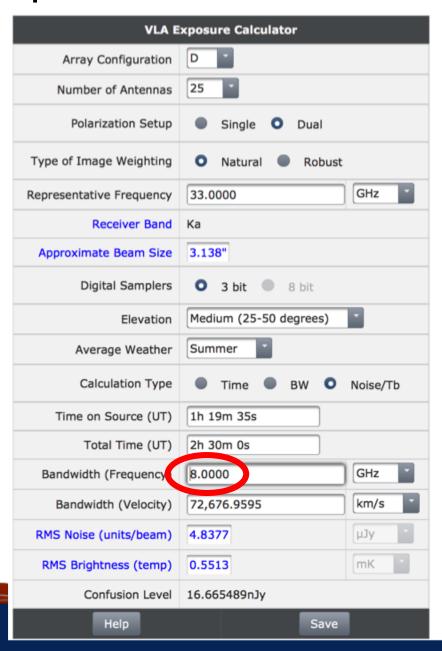
go.nrao.edu/ect

You must provide a value for Frequency. Press the <Tab> key afterwards to move to the Bandwidth field.

Input bandwidth
You must provide a value for Bandwidth. Press the <Tab> key to leave the field and perform the calculation.



VLA Exposure Calculator



go.nrao.edu/ect

Samplers have been switched to 3-bit

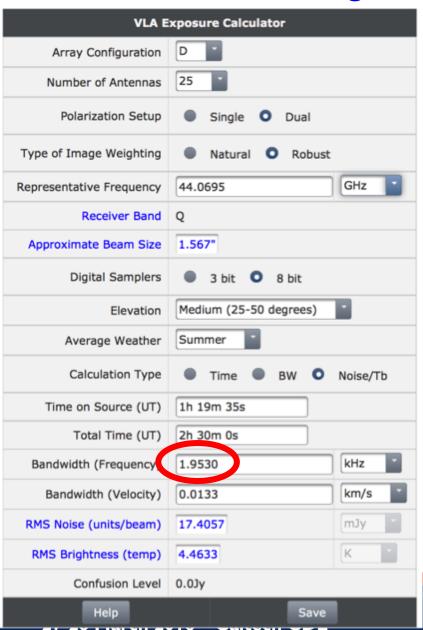
The calculations now reflect the use of the 3-bit digital samplers.

Compared to the 8-bit samplers there is about a 15% sensitivity penalty when using the 3-bit samplers.



VLA Exposure Calculator

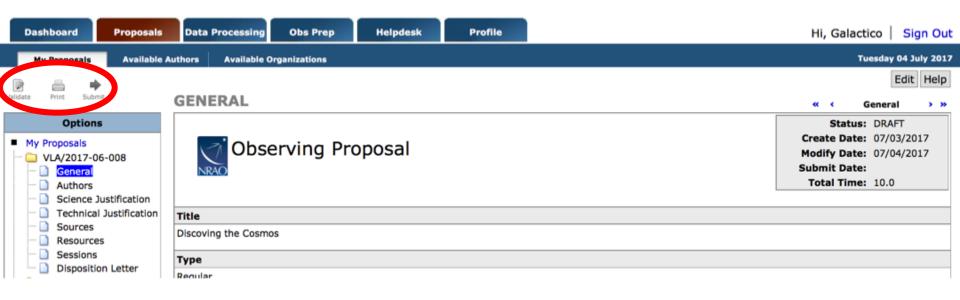
go.nrao.edu/ect







Go back to "General" and see that the total time is filled in after setting the sessions



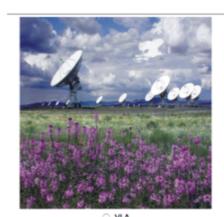
You may "Validate", "Print" and "Submit" when you are ready.



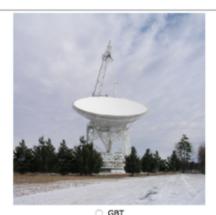


For VLBA Proposals

- Also use the PST
- Many elements are similar to the VLA's
- Will highlight the sections that are significantly different.





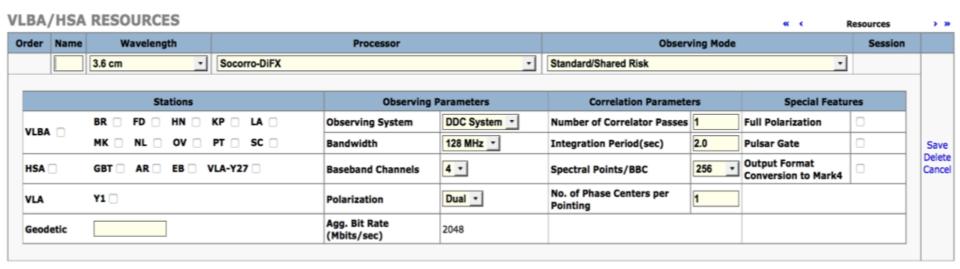








VLBA Resources

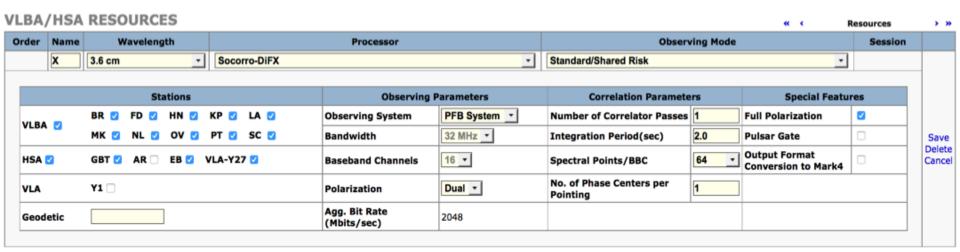


• Similar to VLA you can "Copy Resources" from another project or "Add" a resource.





VLBA Resources



- For VLBA proposals you choose which antennas you want/need.
- In this example, we chose all VLBA antennas + HSA (GBT, Y27, EB) to boost the sensitivity.
- Also chose maximum bandwidth (256MHz or 2048 Mbps)
 - 16 baseband channels with dual polarization means: 8 in RCP + 8 in LCP delivering 256 MHz total per polarization.





VLBA Technical Justification

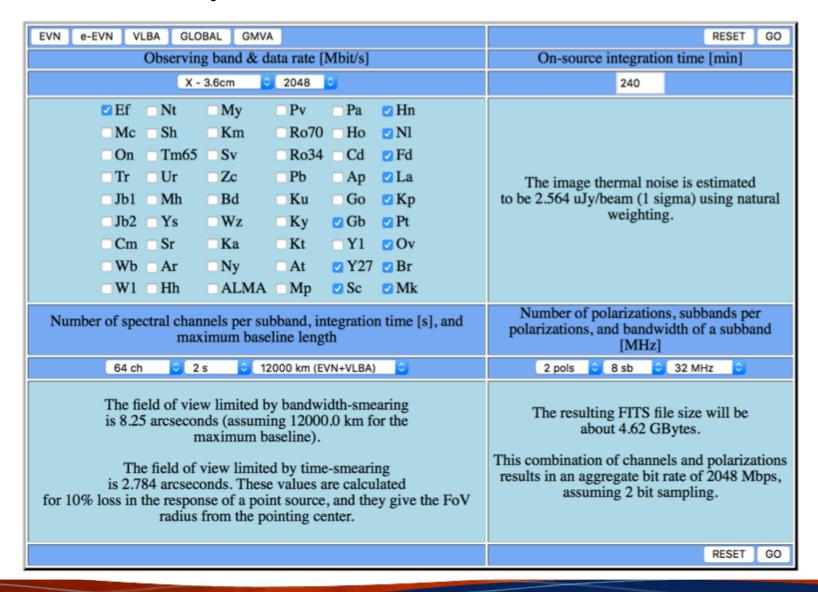
TECHNICAL JUSTIFICATION **Technical Justification VLBA Technical Justification** Use this page to specify how the technical set-up requested for your proposal enables the scientific goals to be met. Input is required for all fields. If a field is not relevant for your proposal then enter "NA" into the textbox. The links within each box provide information concerning these technical questions. Explain the reasons for the stations requested; specify minimum number acceptable, and note which stations are optional and/or required. If HSA observations are being requested, justify why the HSA is needed to achieve the science, and verify that all stations can sample/record with the same observing mode. https://science.lbo.us/facilities/vlba/docs/manuals/oss/ang-res https://science.lbo.us/facilities/vlba/docs/manuals/oss/vlba-plus Explain the choice of receiver(s) requested and whether or not dual polarization is required for each receiver: https://science.lbo.us/facilities/vlba/docs/manuals/oss/bands-perf Explain scheduling issues including requested weather conditions, dates, and length of scheduling blocks. Specify the weather suitable for a given frequency band. For example - 'I request weather suitable for the 2cm band'. Note that this is not necessarily the observing frequency (since one may request lower or higher frequency weather). Specify preferred dates, or excluded dates, and/or if a series of observations with specified cadence, specify that cadence. Specify minimum length of scheduling blocks (blocks of observing time, which may be different than sessions) that can be observed and a start-time range in Pt_LST; note that shorter blocks are, in general, easier to schedule; if 24-hour blocks are required, indicate whether of not break-points may be installed in the schedule to allow different start times. Describe correlator set-up requested. Correlation parameters beyond those required for narrow-field continuum or spectral line observing should be justified. For example, use of pulsar processing, multiple phase centers, multiple correlator passes or wide-field phase centers should be explained. These capabilities, used in isolation or in combination, may have an impact on correlator throughput. Also justify the number of multiple phase centers if > 100. https://science.lbo.us/facilities/vlba/docs/manuals/oss/correlator Note whether the target(s) can be self-calibrated and estimate their flux density. If phase-referencing is required, specify the phase-reference calibrators to be used and their expected flux densities, or whether extra time (on the VLBA or VLA) will be required to find calibrators: https://science.lbo.us/facilities/vlba/docs/manuals/oss/bsln-sens https://www.lbo.us/vlba/astro/callb Sensitivity required to achieve the science goal. Include frequency or velocity width assumed, for non-imaging experiments, justify the baseline sensitivity: Required on-source integration time to achieve the required sensitivity, and total time including overhead; include considerations such as uv-coverage needed for precision imaging, recording rate, etc., and assume the minimum acceptable number of stations in calculating the required integration time; please also verify that the time request on the cover page is consistent with that specified here: Please upload EVN exposure calculator graphic(s), if it was used to calculate the integration time needed. Please make sure that all 4 subpanels of the calculator are captured. Multiple files should be uploaded if there are multiple resources. Use your favorite utility (e.g., xv or gimp [linux]; grab or Browse... No file selected. Upload Command+Shift+4 [Mac]) to make a png file of the EVN exposure calculator graphic which can then be uploaded using the browse/upload buttons to the right. File Name Size http://www.evibi.org/cgi-bin/EVNcalc.pl File Name Size delete download https://science.lbo.us/facilities/vlba/docs/manuals/oss/bsln-sens https://science.lbo.us/facilities/vlba/docs/manuals/oss/bands-perf







EVN Sensitivity Calculator www.evlbi.org/cgi-bin/EVNcalc







EVN Sensitivity Calculator www.evlbi.org/cgi-bin/EVNcalc

EVN e-EVN VLBA GLOBAL GMVA	RESET GO		
Observing band & data rate [Mbit/s]	On-source integration time [min]		
X - 3.6cm 2048 2	240		
☑Ef Nt My Pv Pa ☑Hn			
□Mc □Sh □Km □Ro70 □Ho ☑Nl			
□On □Tm65 □Sv □Ro34 □Cd ☑Fd			
□Tr □Ur □Zc □Pb □Ap ☑La	The image thermal noise is estimated to be 2.564 uJy/beam (1 sigma) using natural		
□Jb1 □Mh □Bd □Ku □Go ☑Kp			
□Jb2 □Ys □Wz □Ky ☑Gb ☑Pt	weighting.		
□Cm □Sr □Ka □Kt □Y1 ☑Ov			
□Wb □Ar □Ny □At ☑ Y27 ☑ Br			
□W1 □Hh □ALMA □Mp ☑Sc ☑Mk			

The tool gives the R.M.S. noise value for the full bandwidth.

For spectral-line observations:

- Get the value for wider bandwidth then multiply it by sqrt of # of channels.
- E.g. $\sigma = 2.564 \,\mu Jy/beam$ for 256MHz, so for 125 kHz channels (i.e., 2048 channels):

$$\sigma = 2.564 \times \sqrt{2048} = 116 \,\mu\text{Jy/beam}$$





Some final notes:

- The next VLA/VLBA/GBT proposal deadline is August 1st.
- If you have never proposed before please start early so there is time to get help from the NRAO helpdesk (help.nrao.edu).
- The next deadline is for A and B configurations of the VLA.
- The VLA and VLBA Call for Proposals are at:
 - VLA: https://science.nrao.edu/observing/call-for-proposals/
 - VLBA: https://science.lbo.us/observing/call-for-proposals/
- The VLA and VLBA Observational Status Summaries are at:
 - VLA: go.nrao.edu/vla-oss
 - VLBA: go.nrao.edu/vlba-oss



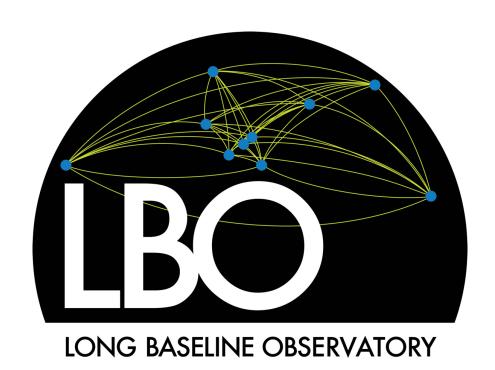






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www.lbo.us

The Long Baseline Observatory is a facility of the National Science Foundation operated under cooperative agreement by Associated Universities, Inc.





