

Feed Analysis ngVLA - SKA and DVA Optics

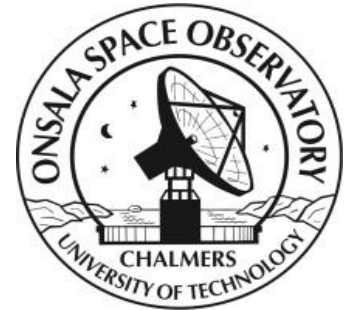
Jonas Flygare

Onsala Space Observatory, Department of Space, Earth and Environment
Chalmers University of Technology

ngVLA Optics Workshop, 19-20 June, 2018
Caltech, Pasadena, CA, US



CHALMERS

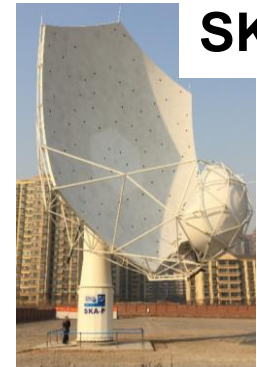
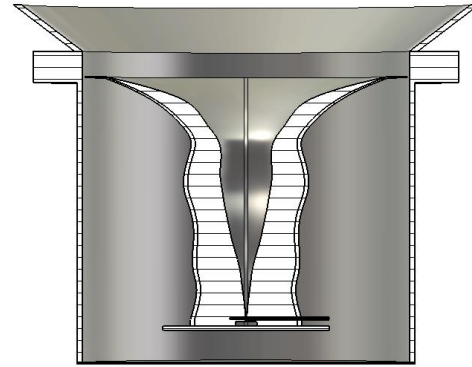


Outline

- Background
- Models
- Modelling
- Some Results
- Summary



Credit: Sander Weinreb



Credit: SKA Organization



VGOS (Geodesy)

**BL: 2-14 GHz
(- 18 GHz OTT)**

Waveband	Default central frequency
18 cm	1664 MHz
13 cm	2268 MHz
6 cm	4992 MHz
5 cm	6668 MHz (Methanol), 6030 MHz (OH)
4 cm	8418 MHz
1 cm	22230 MHz

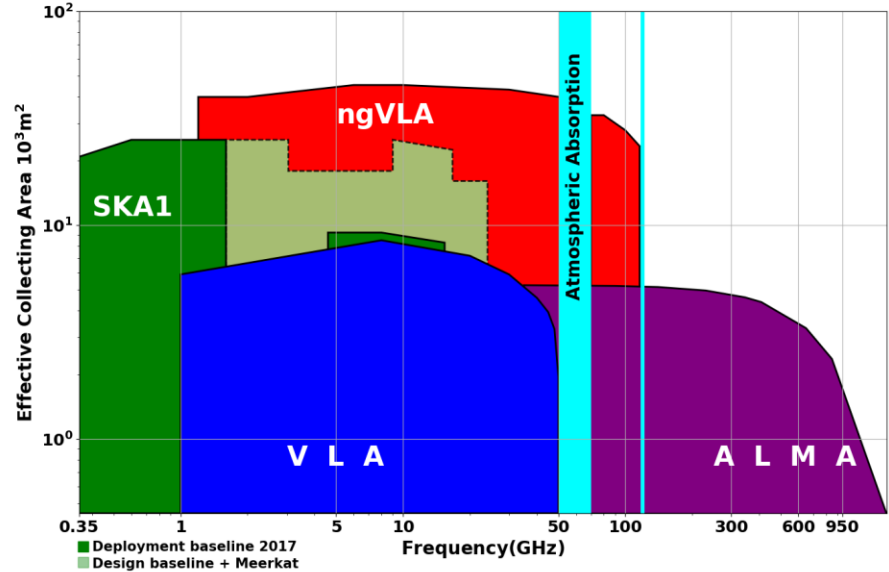
Waveband	Default Central Frequency
90 cm	327 MHz
50 cm	610 MHz
21 cm	1416 MHz
2 cm	15362 MHz
7 mm	43214 MHz

BRAND EVN

(UWB 10:1, 1.5-15.5 GHz)

ngVLA, 214 telescopes

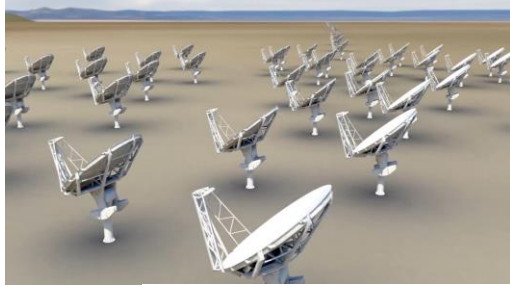
BL: 1.2 - 116 GHz (excl. atmos. absorption)



SKA1 MID, ~130 telescopes

BL: 0.35 – 15.4 GHz

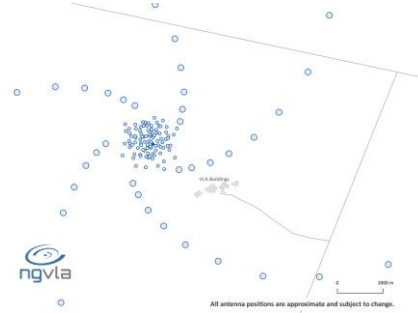
Credit: Roger Hammargren,
ngVLA/NRAO, EVN



ngVLA

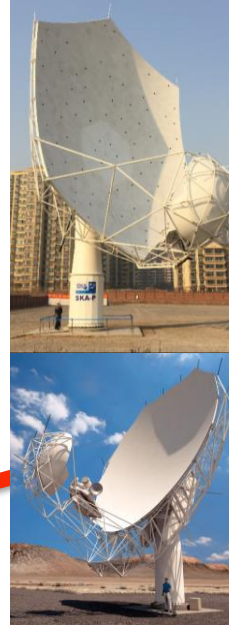


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Onsala Space Observatory, Department of Space, Earth and Environment
Chalmers University of Technology

SKA



MeerKAT

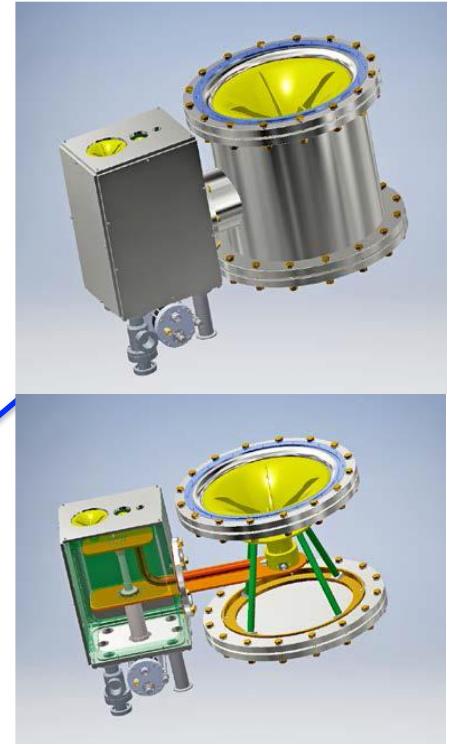
Credit: SKA Organization, ngVLA/NRAO,
Leif Helldner/EMSS/MeerKat

ngVLA 1.2 – 116 GHz

- 2 Dewars, sharing 1 cryo-vacuum system
 - 4-feed system
 - **1.2 – 4.2 GHz (~3.5:1) Prot.**
 - **4.2 – 15 GHz (~3.5:1)**
 - **15 – 50 GHz (~3.5:1)**
 - **70 – 116 GHz (1.65:1)**



Credit: Leif Helldner



"Cryogenic 1.2 to 116 GHz Receiver for Large Arrays"
S. Weinreb, H. Mani, W. Zhong, J. Flygare, B. Billade, A. Akgiray, L. Dong.,
12th European Conference on Antennas and Propagation (EuCAP),
London 2018 April.

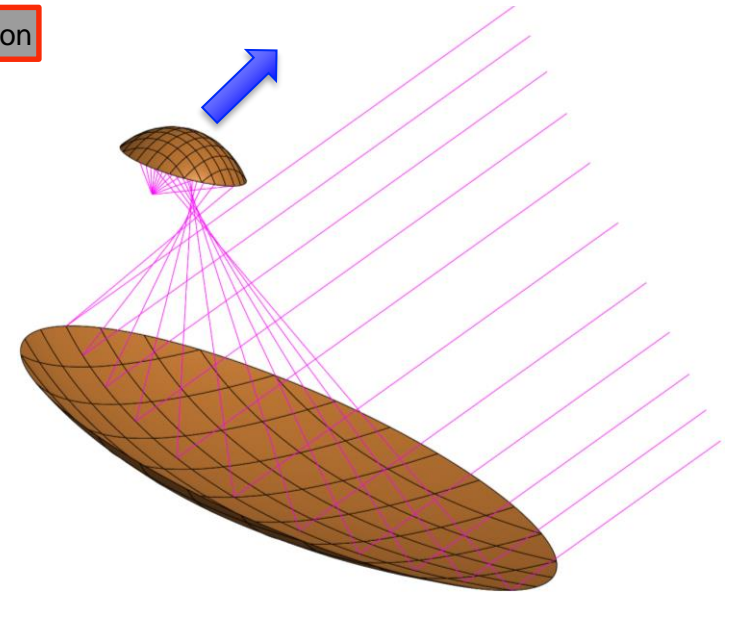
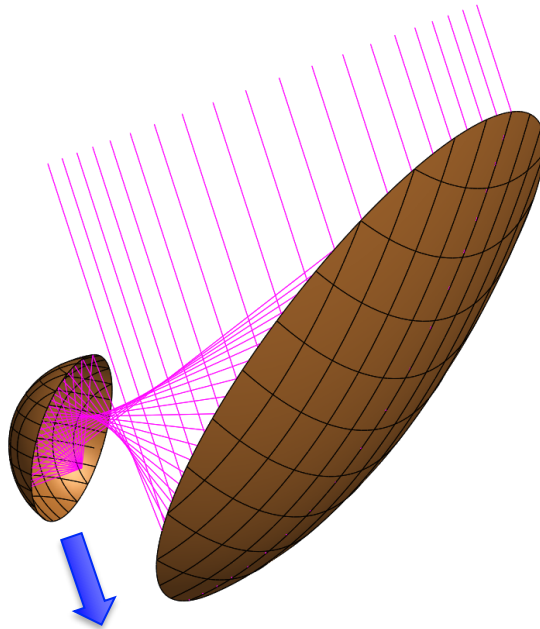
SKA Dish

(Feed down, shaped, spill-over shield
Gregorian Offset, $\theta_e = 58^\circ$)

DVA Dish

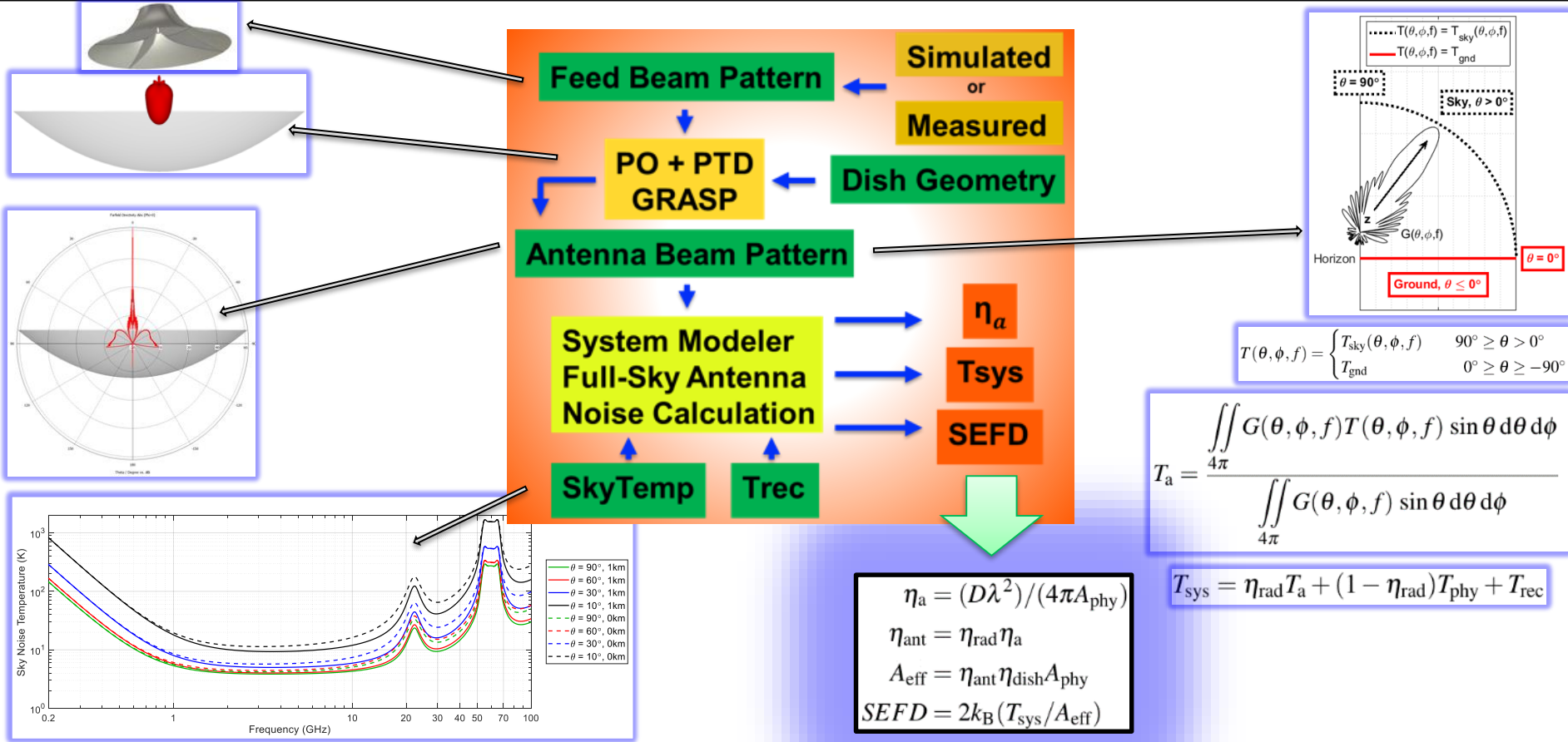
(Feed up, shaped
Gregorian Offset, $\theta_e = 55^\circ$)

Zenith Direction



	SKA	DVA
Shaped	Yes	Yes
Tipping Conf.	Down	Up
SpillOver Ext.	40°	No
θ_e	58°	55°
MR Long Ax.	18m	18m
MR Short Ax.	15m	15m
Proj. (MR) D.	15m	15m
SR D.	~5.2m	~4m

Simplicity: assuming
perfect dish in simulations



➤ **Spill-over shield, feed down**

- **MeerKAT 20°**
- **SKA Dish 40°**

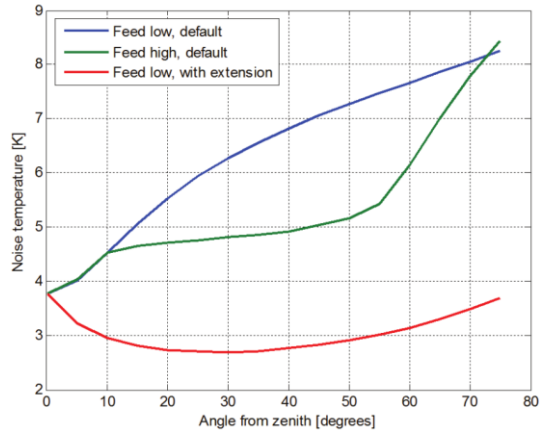
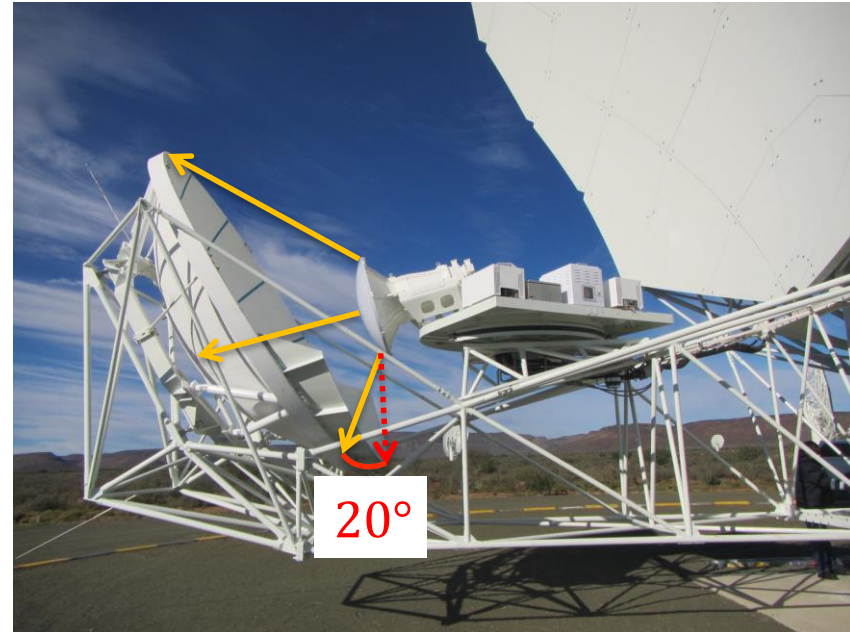


Figure 6: Spill-over tipping curve for the different configurations.

Credit: "The design of the MeerKAT dish optics"
I. P. Theron, R. Lehmensiek, D.I.L. de Villiers
ICEAA'12

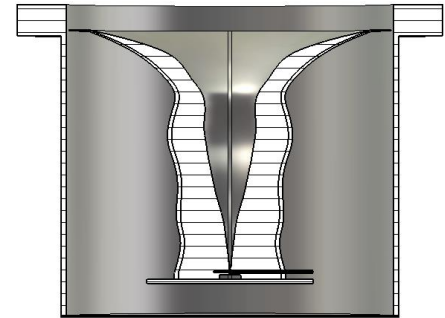
SKA Band 1 on MeerKAT



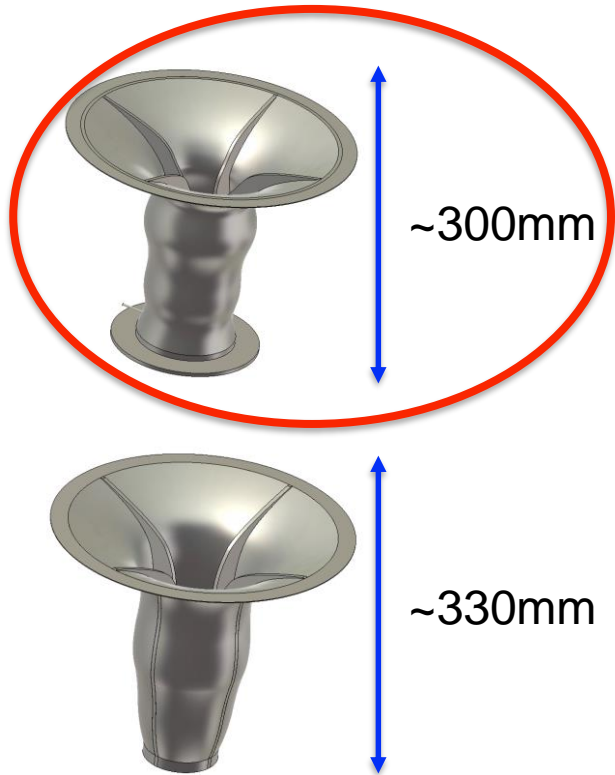
Credit: Leif Heldner, OSO

ngVLA 1.2- 4.2 Prototype

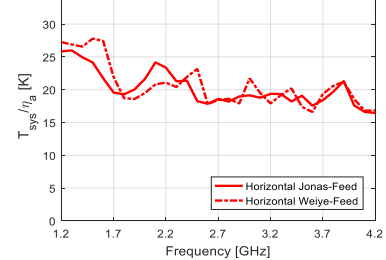
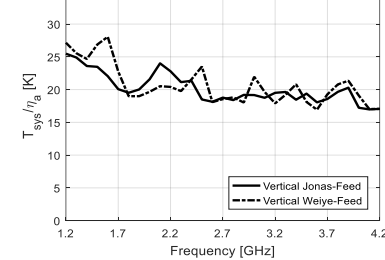
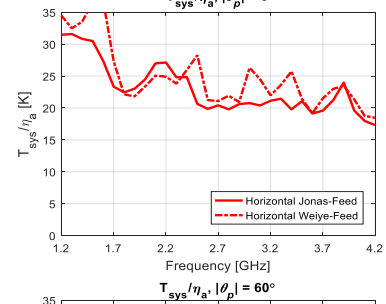
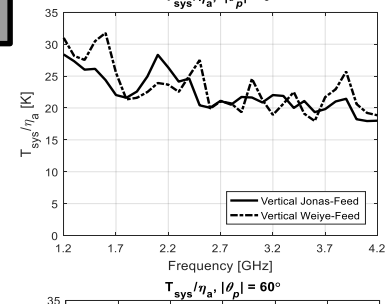
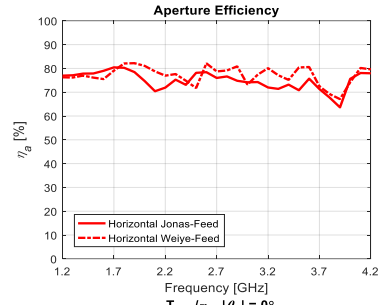
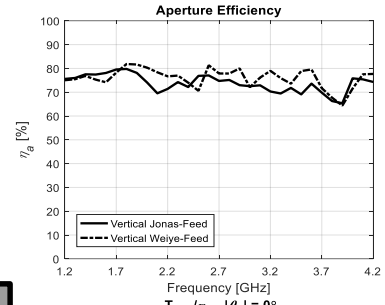
- Optimized for SKA Dish, 12 dB @ 58°
 - NOT optimized for DVA = 55°, 16 dB!
 - Keep in mind when looking at data
- Goal $\eta_a > 75\%$ ave. (achieve $> 77\%$ on SKA, $> 72\%$ on DVA)
- S11 < -10 dB (achieve < -12 dB, average -15 dB).
- Dewar dimensions constraints



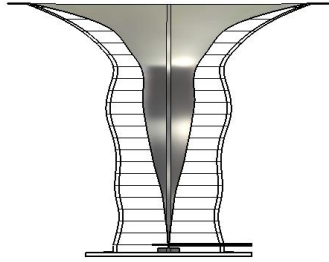
Pre-dewar Int., parallell design work



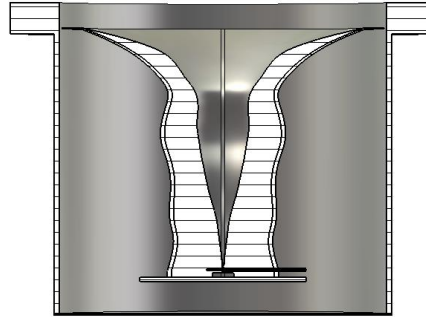
Smaller footprint



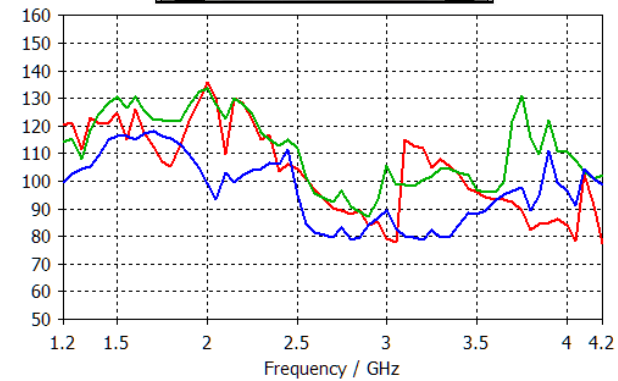
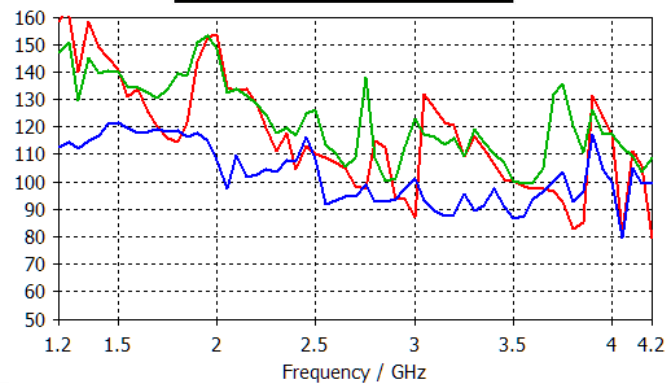
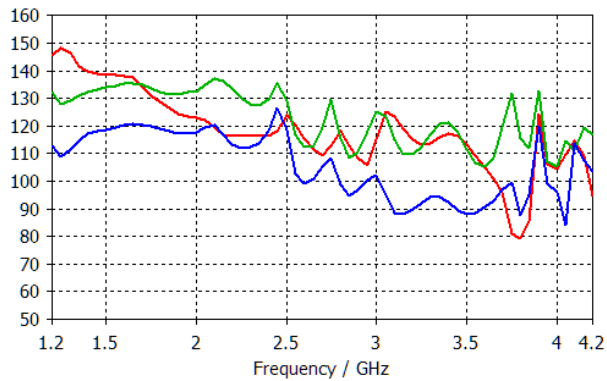
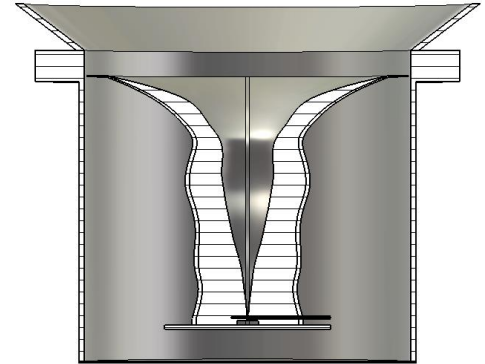
Feed



Feed in dewar

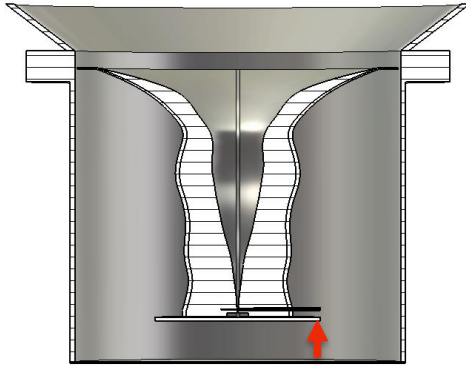


Feed in dewar + cone

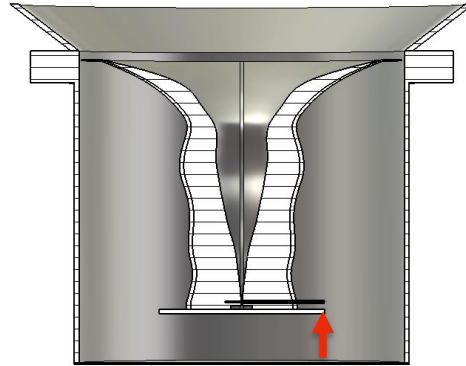


Beamwidth @ 12 dB - φ=0°, φ=45°, φ=90° - Goal: 2 x 58° = 116° (SKA)

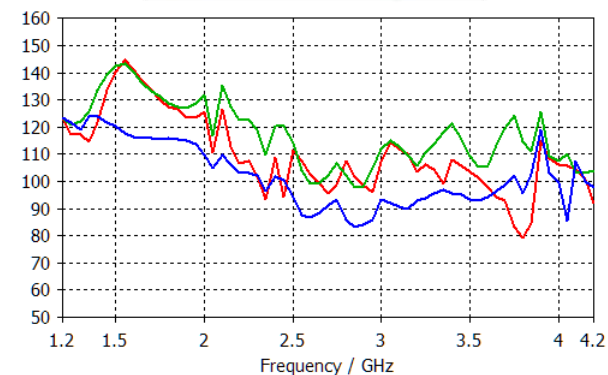
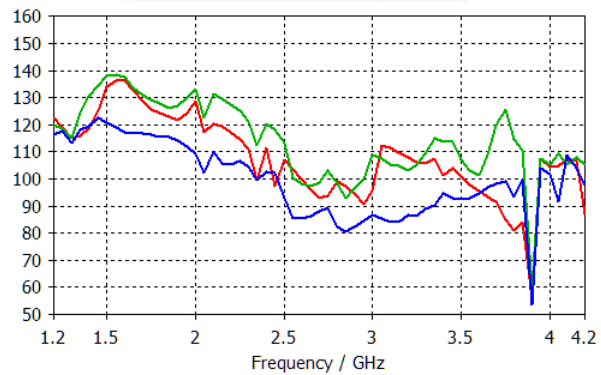
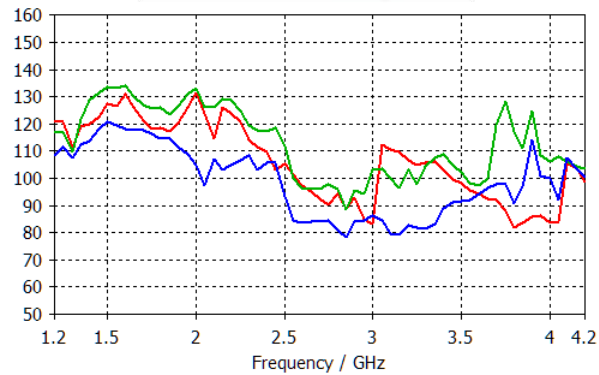
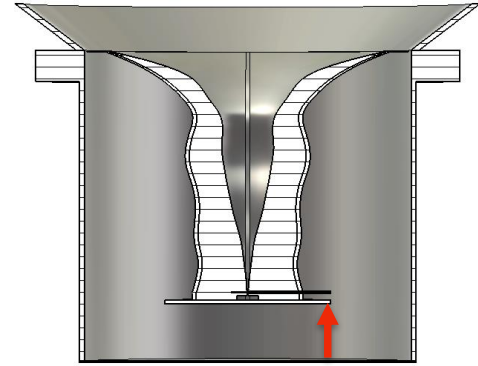
Feed (+10mm) in dewar + cone



Feed (+20mm) in dewar + cone

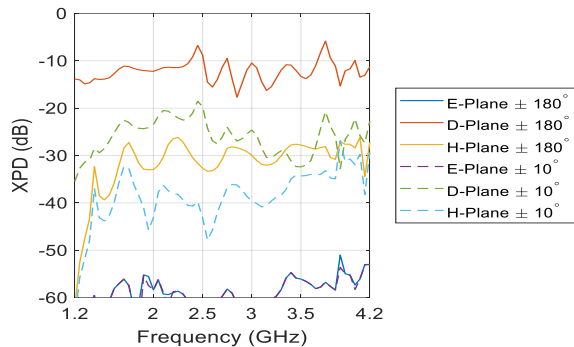


Feed (+30mm) in dewar + cone

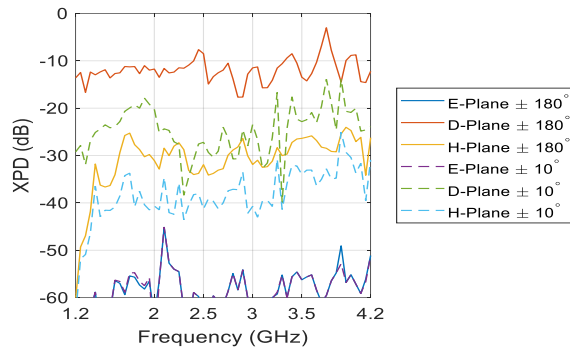


Beamwidth @ 12 dB - $\phi=0^\circ$, $\phi=45^\circ$, $\phi=90^\circ$ - Goal: $2 \times 58^\circ = 116^\circ$ (SKA)

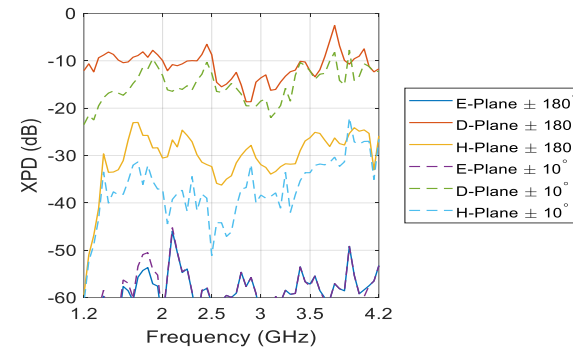
Feed



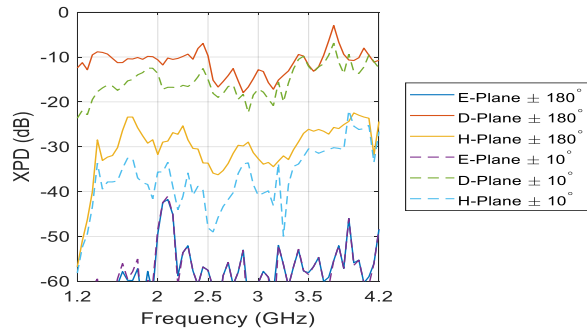
Feed in dewar



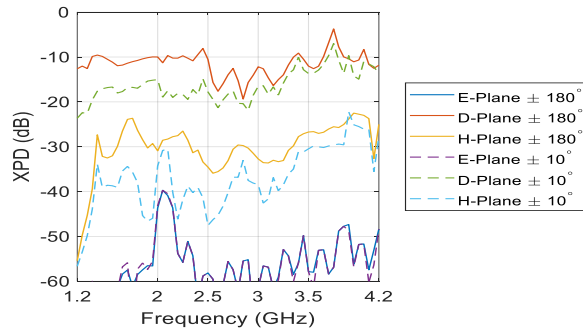
Feed in dewar + cone



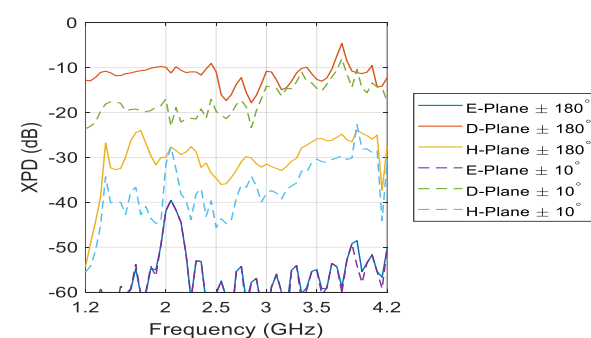
Feed (+10mm) in dewar + cone



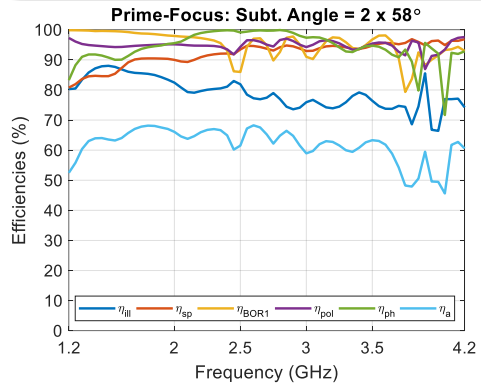
Feed (+20mm) in dewar + cone



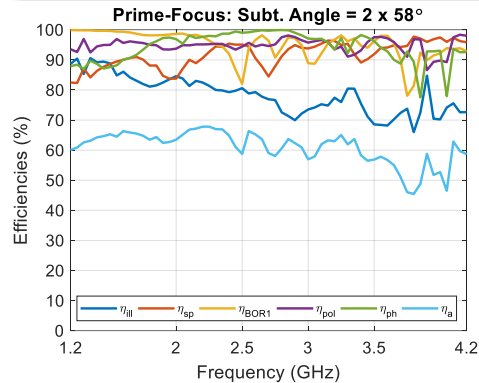
Feed (+30mm) in dewar + cone



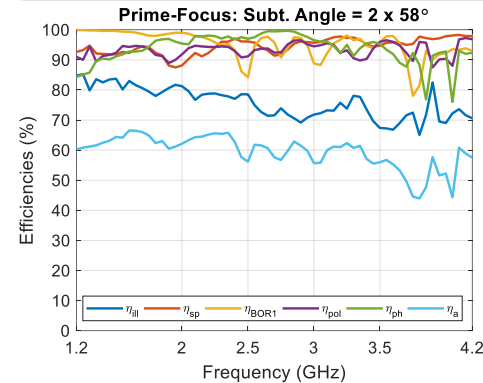
Feed



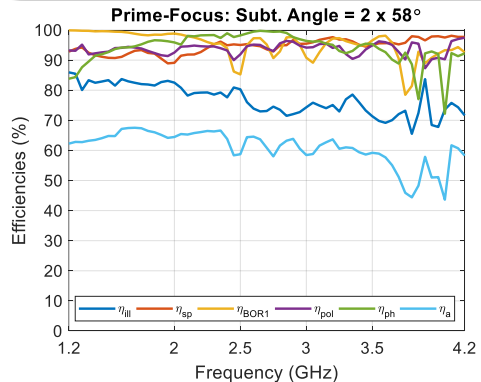
Feed in dewar



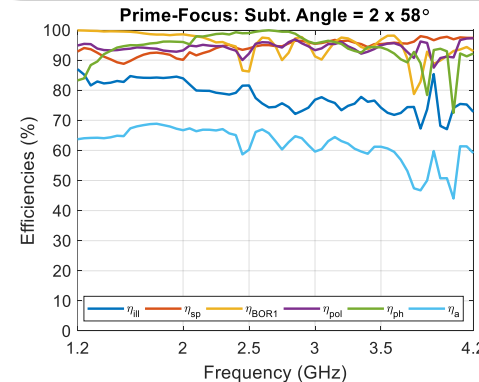
Feed in dewar + cone



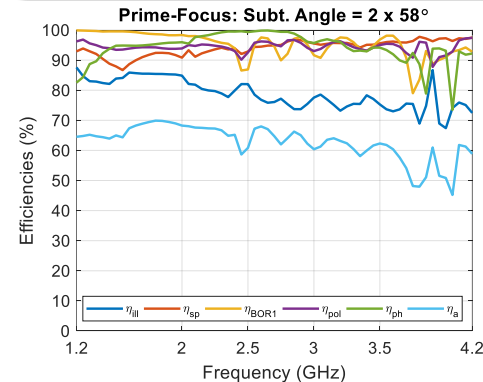
Feed (+10mm) in dewar + cone

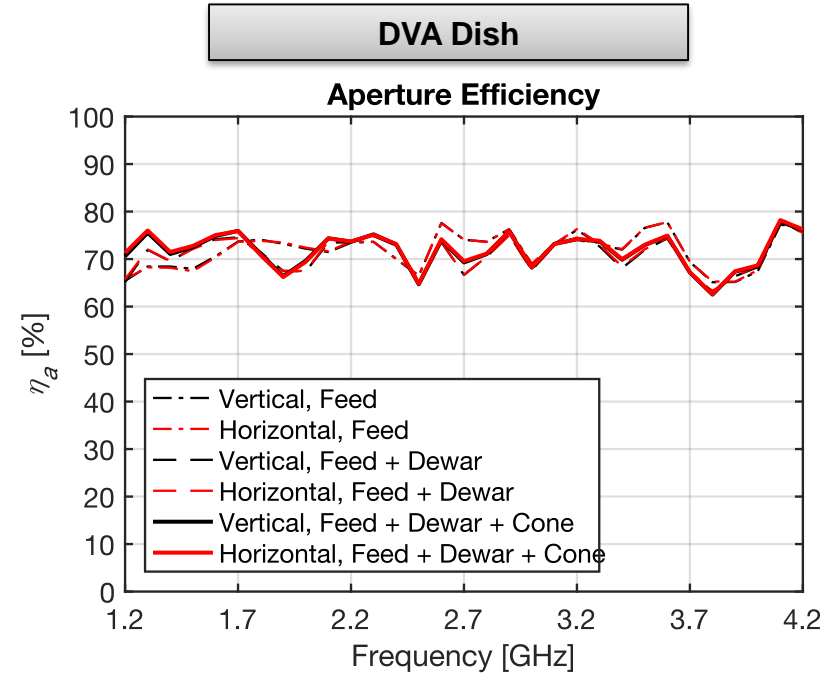
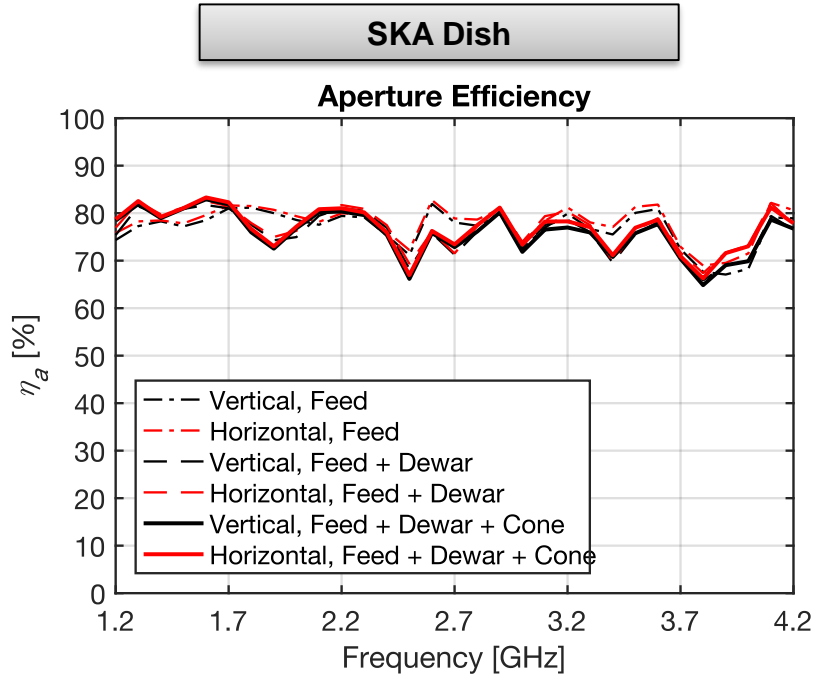


Feed (+20mm) in dewar + cone



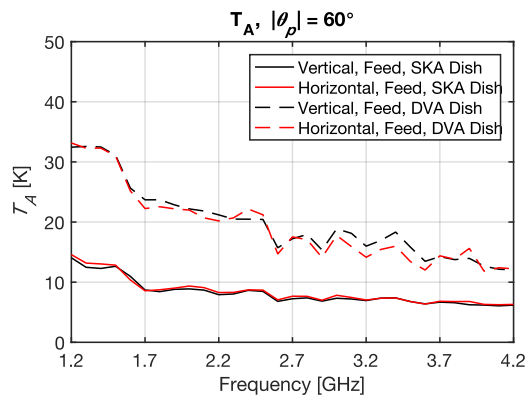
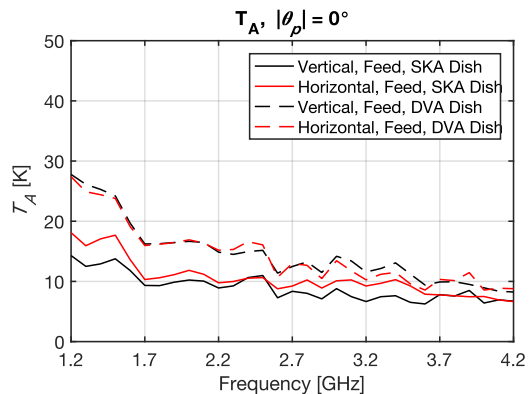
Feed (+30mm) in dewar + cone



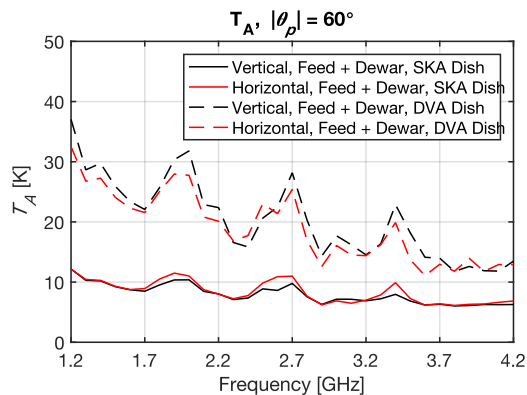
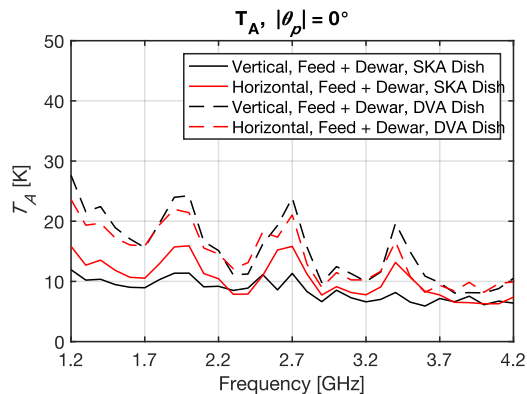


	Feed	Feed + Dewar	Feed + Dewar + Cone
SKA	77,4 %	76,5 %	76,4 %
DVA	71,9 %	71,0 %	71,8 %

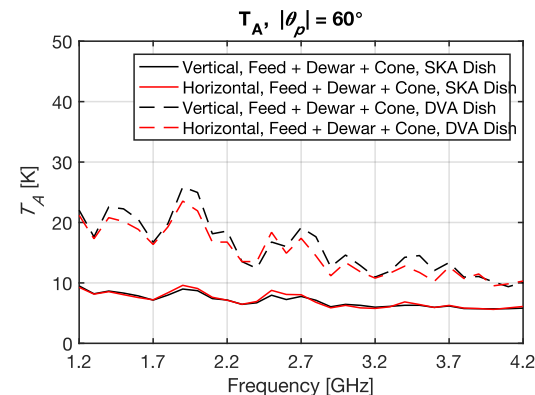
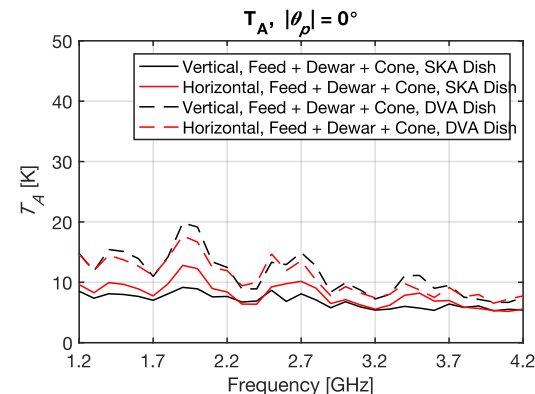
Feed



Feed in dewar

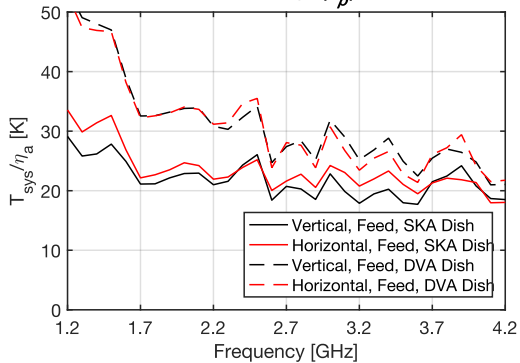


Feed in dewar + cone

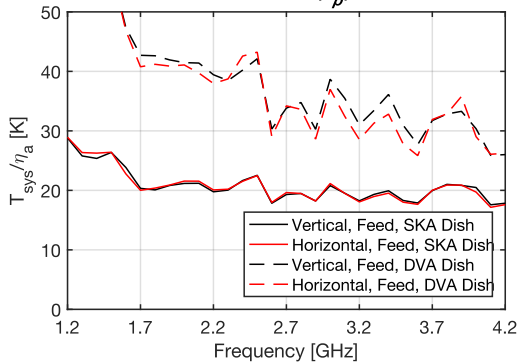


Feed

Alt. Sensitivity, $|\theta_\rho| = 0^\circ$

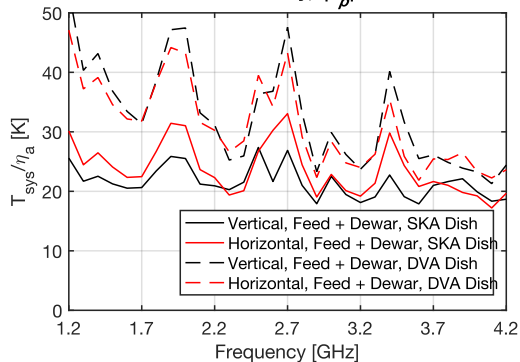


Alt. Sensitivity, $|\theta_\rho| = 60^\circ$

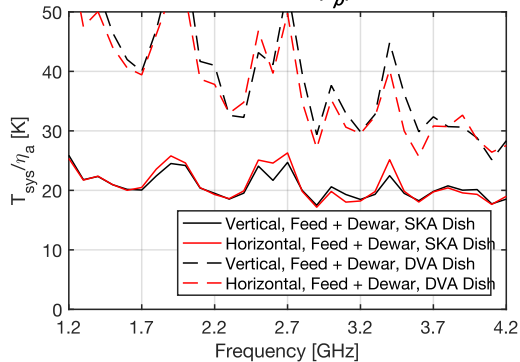


Feed in dewar

Alt. Sensitivity, $|\theta_\rho| = 0^\circ$

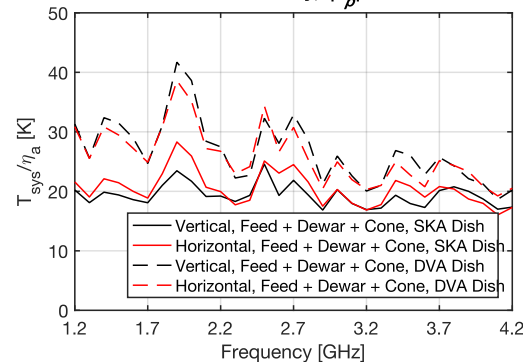


Alt. Sensitivity, $|\theta_\rho| = 60^\circ$

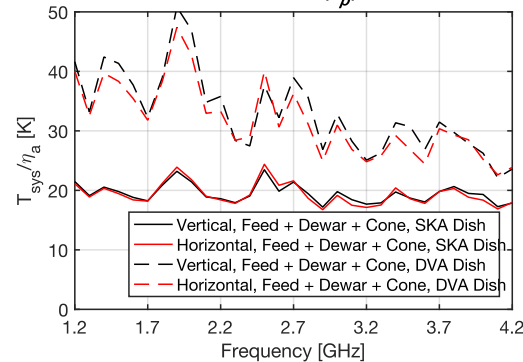


Feed in dewar + cone

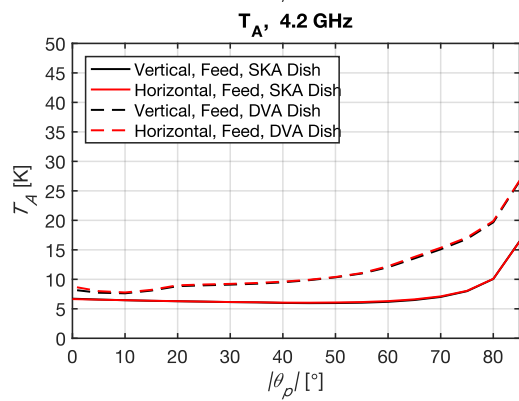
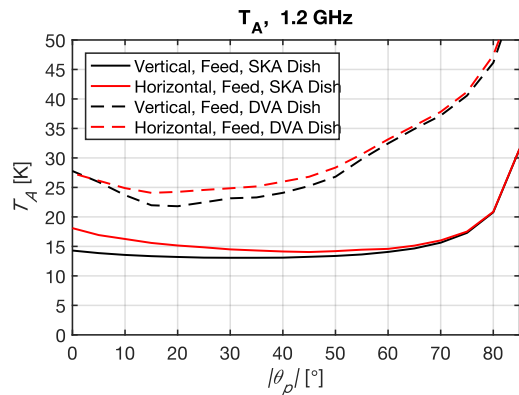
Alt. Sensitivity, $|\theta_\rho| = 0^\circ$



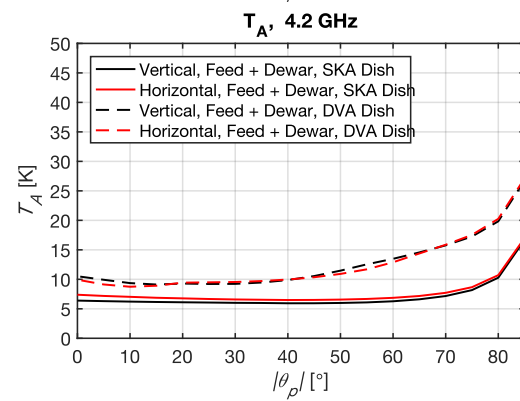
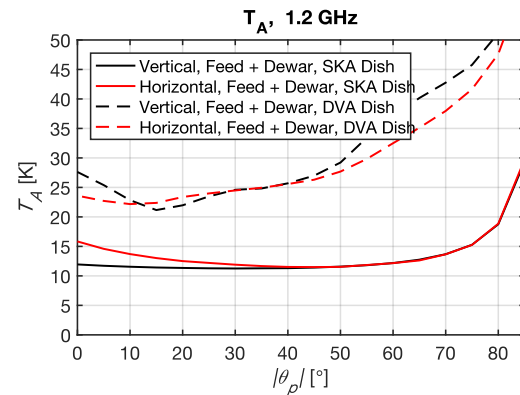
Alt. Sensitivity, $|\theta_\rho| = 60^\circ$



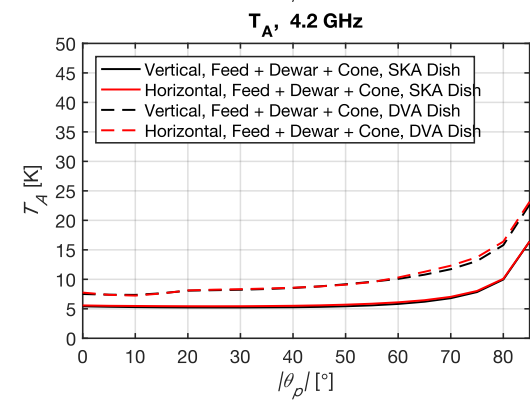
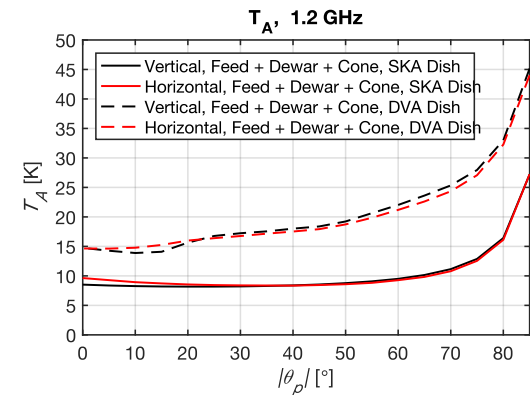
Feed



Feed in dewar



Feed in dewar + cone



SKA Dish

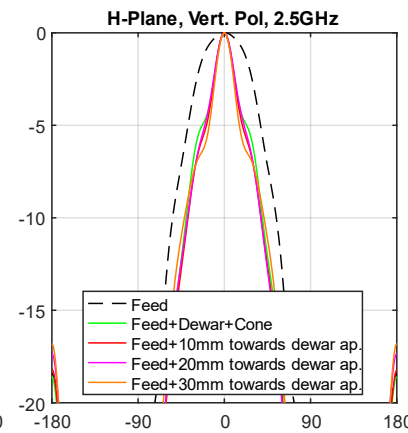
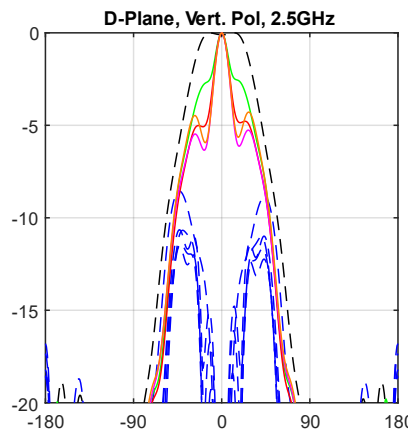
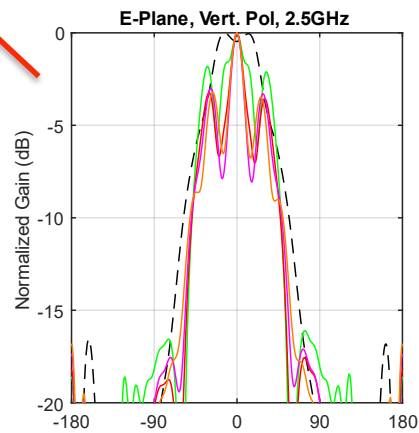
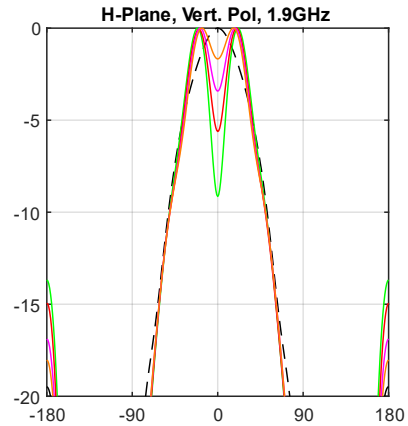
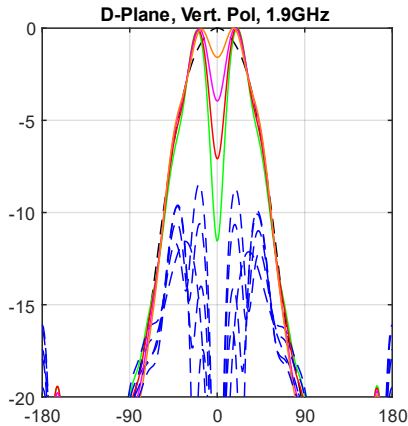
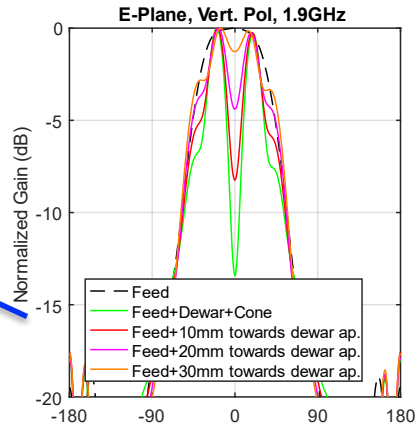
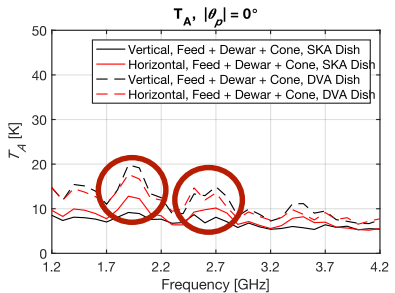
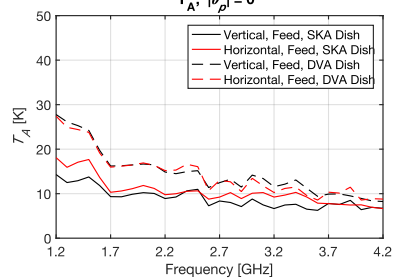
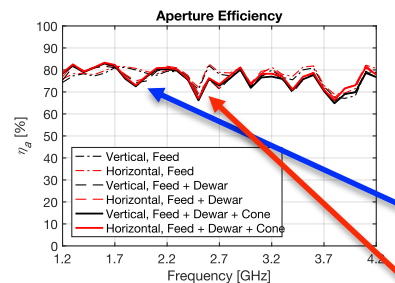
(Feed down, shaped,
spill-over shield
Gregorian Offset, $\theta_e = 58^\circ$)

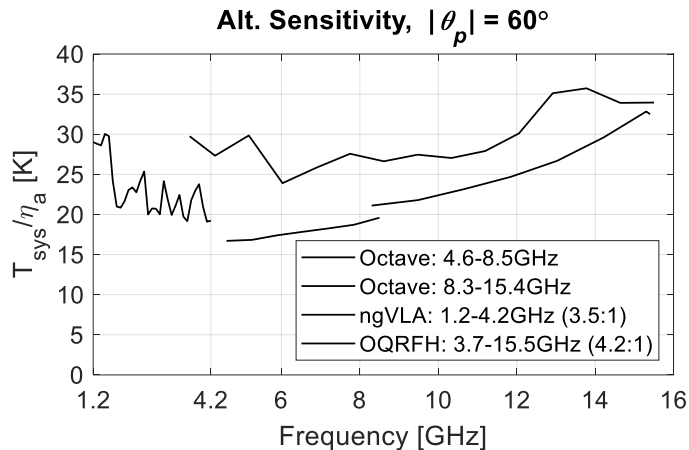
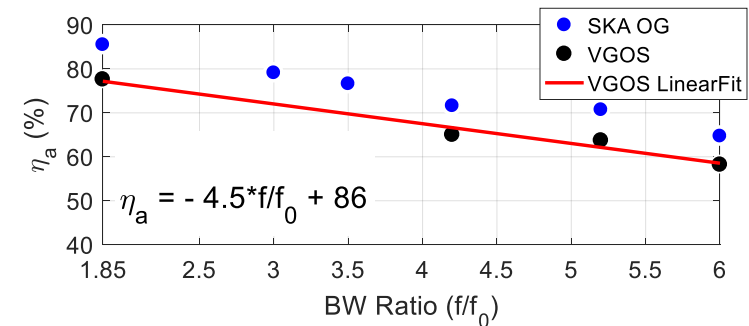
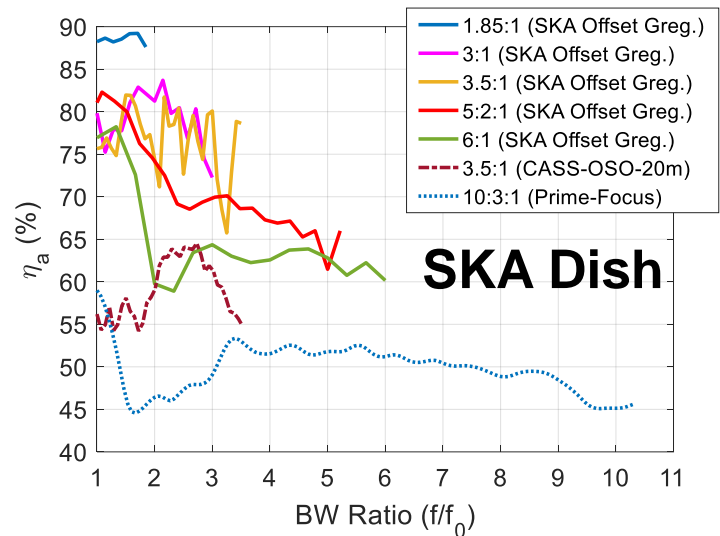
	ApEff	Ta 0	Ta 60	Tsys 0	Tsys 60	Tsys/Ap Eff 0	Tsys/Ap Eff 60	SEFD 0	SEFD 60	IXR
Feed	77,4	9,8	8,3	17,5	16,1	22,6	20,8	353,7	324,6	13,5
Feed + Dewar	76,5	9,5	8,2	17,3	15,9	22,6	20,9	353,2	326,1	14,7
Feed + Dewar + Cone	76,4	7,4	7,1	15,2	14,8	19,9	19,4	311,0	303,5	15,1

DVA Dish

(Feed up, shaped
Gregorian Offset, $\theta_e = 55^\circ$)

	ApEff	Ta 0	Ta 60	Tsys 0	Tsys 60	Tsys/Ap Eff 0	Tsys/Ap Eff 60	SEFD 0	SEFD 60	IXR
Feed	71,9	14,5	21,1	22,2	28,8	31,2	40,4	487,0	631,6	13,9
Feed + Dewar	71,0	14,8	21,2	22,5	29,0	31,9	41,1	498,5	641,8	14,3
Feed + Dewar + Cone	71,8	11,1	15,7	18,8	23,4	26,3	32,7	410,6	511,3	14,6





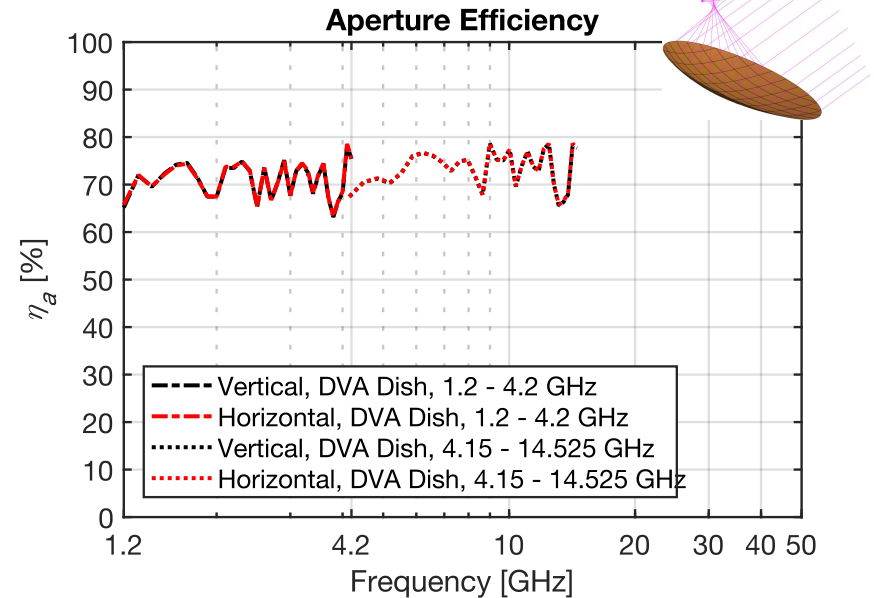
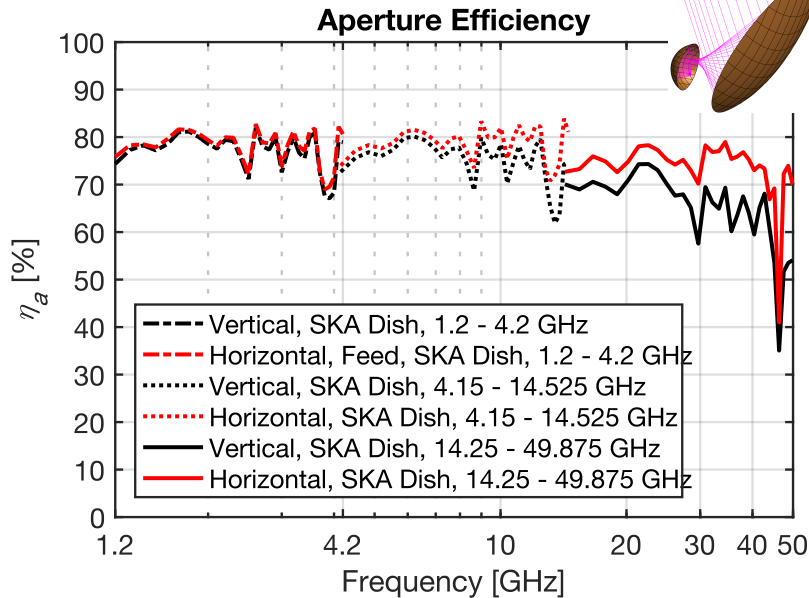
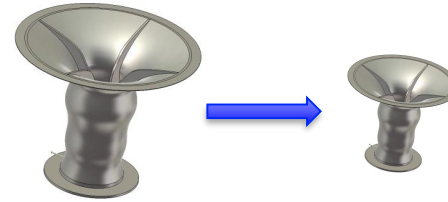
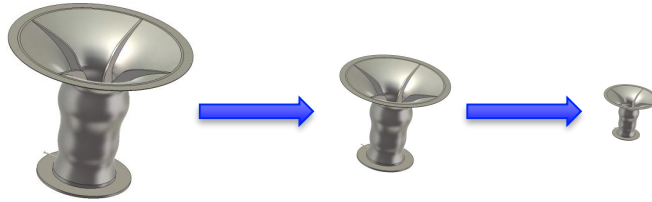
$$\eta_a = (D\lambda^2)/(4\pi A_{\text{phy}})$$

$$\eta_{\text{ant}} = \eta_{\text{rad}} \eta_a$$

$$A_{\text{eff}} = \eta_{\text{ant}} \eta_{\text{dish}} A_{\text{phy}}$$

$$SEFD = 2k_B(T_{\text{sys}}/A_{\text{eff}})$$

BW	η_a (%) (SKA OG)
1.85:1	85,5
3:1	79,1
3.5:1	76,6
4.2:1	71,6
5.2:1	70,7
6:1	64,7



- WBSPF SKA: 4.6 – 24 GHz
- Spillover improved with cone

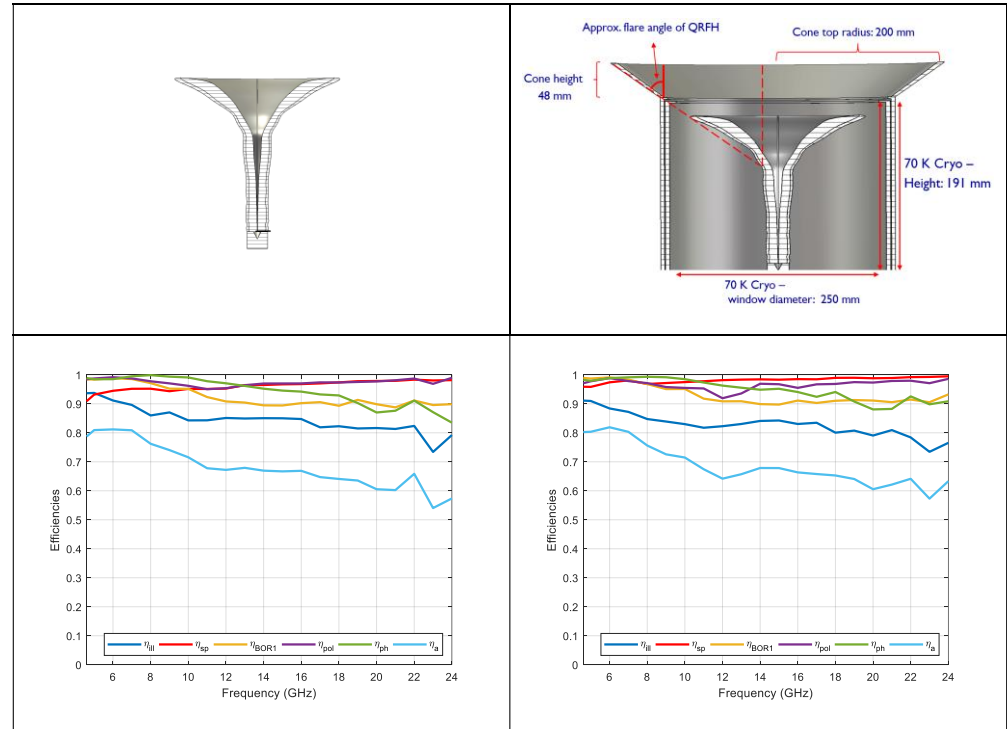
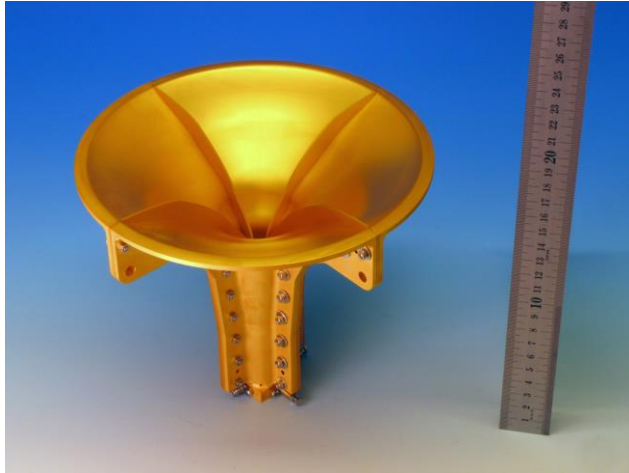
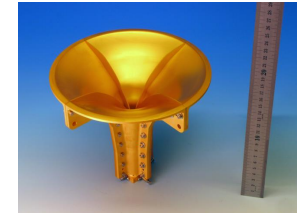
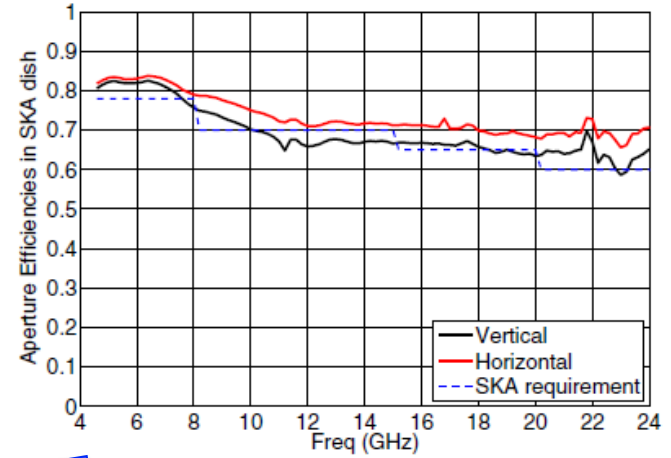
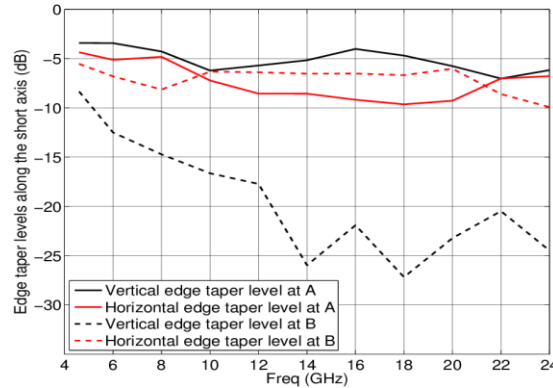
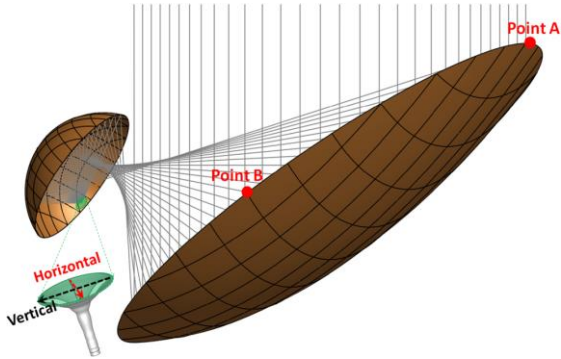


Figure 1: Aperture, spillover and polarization efficiency in the $\theta_o = 58^\circ$ SKA Dish for the Band B QRFH in free space (left) and the Band B QRFH inside the cryostat (250 mm diameter IR window) with spillover cone (right). Especially notice the improved spillover efficiency (red) with the cone mounted.

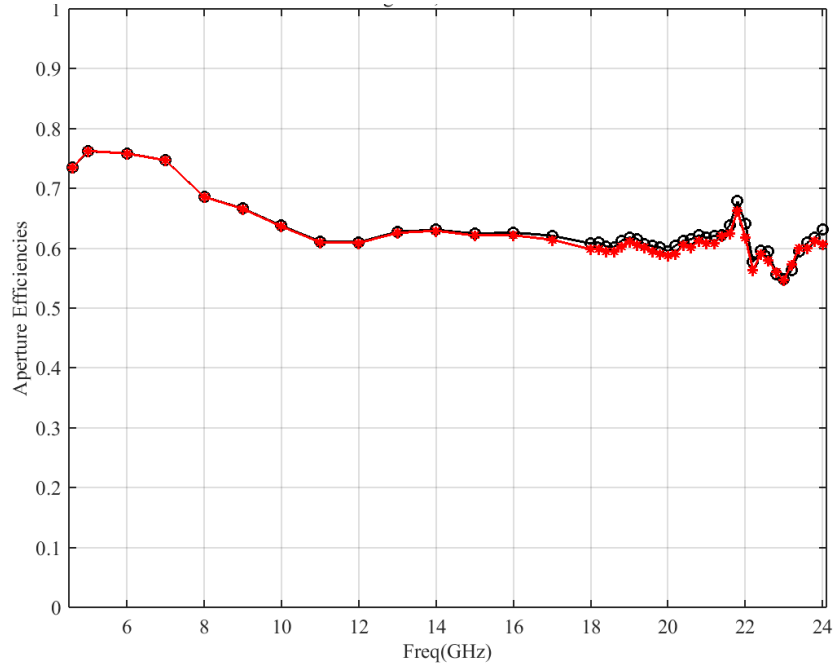
Credit: F. Mokhupuki

Polarization discrepancy – SKA Dish

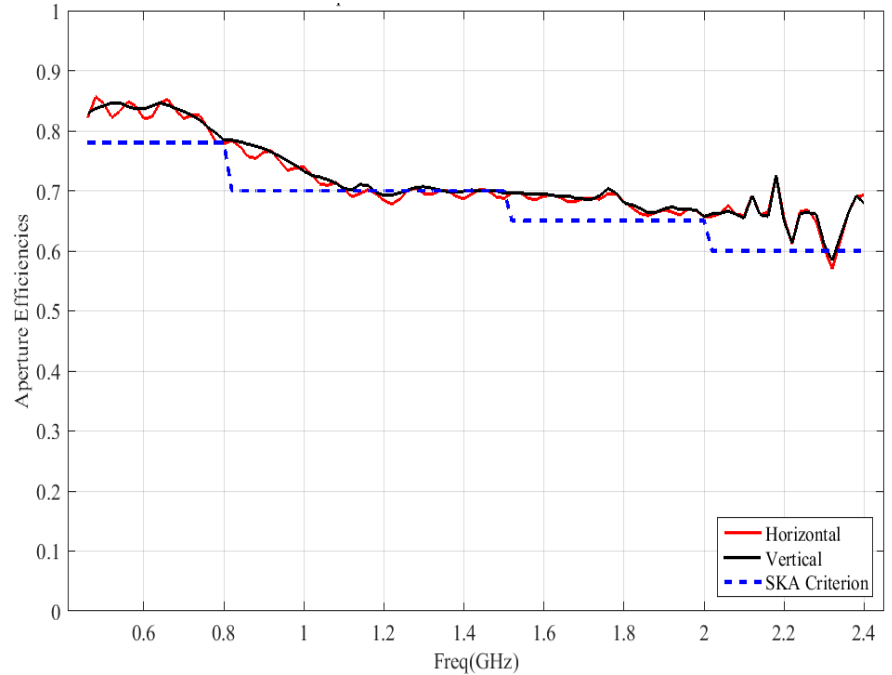
- High-frequency, 4.6 – 24 GHz.
- Vertical taper at point B decreasing with frequency
 - Efficiency follows



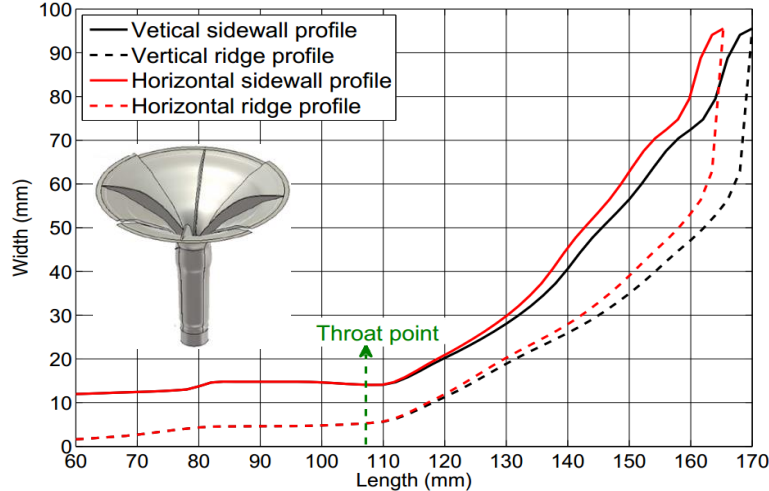
Band-B feed in a primary-dish:
No discrepancy



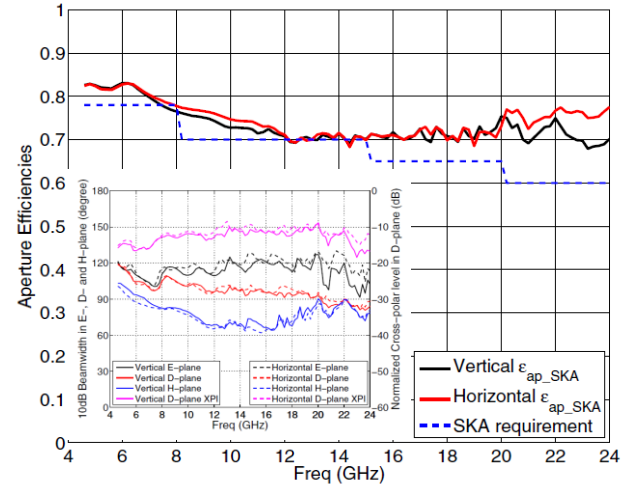
Scaled to 0.46-2.4GHz: No discrepancy



Assymmetric QRFH solution



- Profiles' difference begins only from the “throat point” where the feed opens up;



- Asymmetry caused **resonance** at high band.

- Discontinuity gives slightly worse S11



Solution is Elliptical QRFH

Future Work

- Feed was optimized for SKA Dish, should be re-optimized for new ngVLA dish
- Re-Opt for sensitivity (spill-over reduction)
- “Smoother” efficiency
- Optics specification 55° , feed up?
- Scale Band B feed to 4 – 21 GHz for midrange feed?
- Implement fixed dewar dimensions for optimization to mitigate split-beam effect
- Implementation of spill-over cone on dish? Effects on indexer rotation of feeds?

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