ALMA Data Products



Adele Plunkett NRAO





This talk is for you if...

- You are PI or co-I of an ALMA project with data in hand, or soon to be delivered (i.e. Cycle 5 and/or Cycle 6).
- You want to start searching the ALMA archive for data.
- You have a fabulous science case that will be essential to follow-up with ALMA in future cycles.

This talk will be available online for reference after this workshop.



The condensed version

- Data are delivered after passing Quality Assurance (QA)
- Download data from Archive Query and Request Handler tools on the ALMA Science Portal
- Delivered data include:
 - Calibration tables and diagnostics
 - Preliminary images (better products may be possible with more careful continuum identification & interactive cleaning)
 - —Instructions on how to proceed (weblog!)
- See Sections 11- 13, and Appendix C of ALMA Technical Handbook (Cycle
 6)



The reference I often use (or refer others to):

https://help.almascience.org/index.php?/Knowledgebase/Article/View/375/

Knowledgebase:

What Cycle 4 or Cycle 5 Calibration and Imaging products will be delivered to me?



Posted by Catarina Ubach, Last modified by Sarah Wood on 24 May 2018 03:12 PM

What you see in your data package when you download it from the archive will depend on whether the imaging and calibration have been performed by the pipeline, or one or both via a manual process. Except for datasets that were put through the imaging pipeline, the data products are similar to those for Cycle 3, discussed in detail in http://almascience.org/documents-and-tools/cycle3/ALMAQA2Products3.0.pdf.

For Cycle 4 Imaging pipeline data and all of Cycle 5 data, the products are discussed in detail in http://almascience.org/documents-and-tools/cycle5/ALMAQA2Productsv5.1.pdf

In this article we give a quick guide to the data packages, and more detailed information on the imaging pipeline products for data processed after October 1 2016.



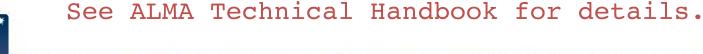
More details here:

https://almascience.nrao.edu/documents-and-

coole/evole5/ALMAOAODroductov5-1-redf

Goals of Quality Assurance (QA) Process

- Ensure reliable final data product
 - Desired sensitivity (as specified by PI)
 - Desired resolution (as specified by PI)
- Ensure calibration and QA imaging free from major artifacts
- Warning: Errors in PI-supplied parameters are outside scope of QA process, including:
 - –Incorrect source coordinates
 - Inadequate frequency specification
 - Inadequate sensitivity limits





During Observations – QA0

- Monitoring of on-the-fly calibration and system performance
- Rapidly-varying parameters (~SB/EB timescales)
 - –Atmospheric effects
 - –Antenna issues
 - —Front-end issues
 - –Connectivity issues
 - —Back-end issues
- Tolerances for each are explicitly laid out
 - –No fewer than 40 antennas in 12m array
 - Bandpass calibrator is strong enough
- Quick reduction may be run to check flux measurements and phase stability





Between Observations – QA1

- "Regular array maintenance" timescales
- Slowly Varying Parameters (~MOUS timescales)
- General array calibration
 - -Baseline measurements
 - –Delays
- Antenna Calibrations
 - –All-sky pointing
 - –Focus curves
 - -Beam patterns, etc.
- Observatory Calibrator Surveys
 - -Solar-system and quasar flux monitoring



After Observations – QA2

- Calibration by pipeline (~70%) or DA/staff.
- Final QA checks include
 - —RMS of complex antenna-based gains
 - Absolute flux calibration scale
 - -T_{svs} within acceptable range
 - Proper phase transfer cadence
 - Proper bandpass corrections
- Assessment of Imaging Products
 - —Signal-to-noise and angular resolution
 - –No strong artifacts
 - —Performed on the reference source/spectra
- Information about QA review is aggregated for delivery in the QA2 Report



After Delivery – QA3

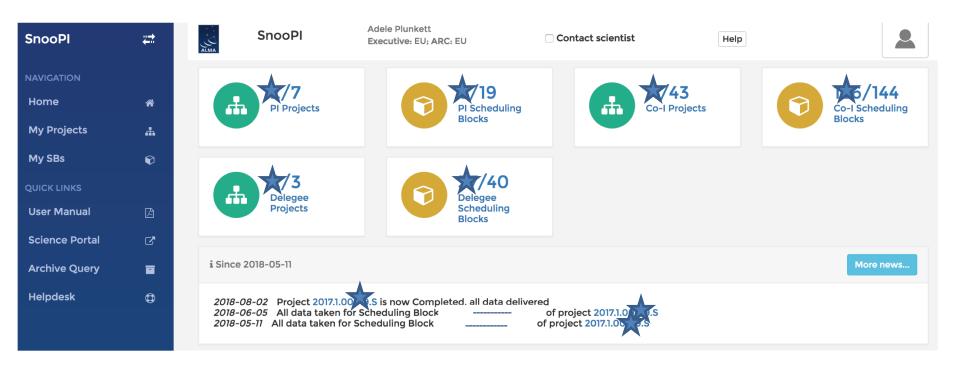
- Additional QA stage possibly triggered by PI reporting any issues underlying:
 - Data, observing procedure, calibration
- Re-evaluation of calibrated data products
 - -Only occurs if QA0 -> QA2 miss something
- Likely results in fix being implemented and products re-ingested into ALMA archive
- Proprietary period extension (within two months of delivery)



After two months, extension only until fix is delivered

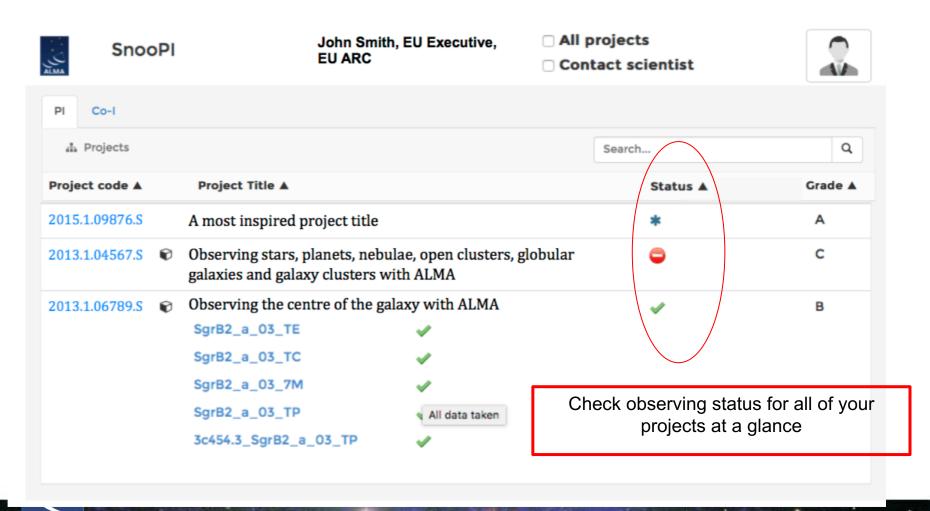


https://almascience.nrao.edu/observing/snoopi





Listing of PI (or co-PI) projects





Listing of PI (or co-PI) projects

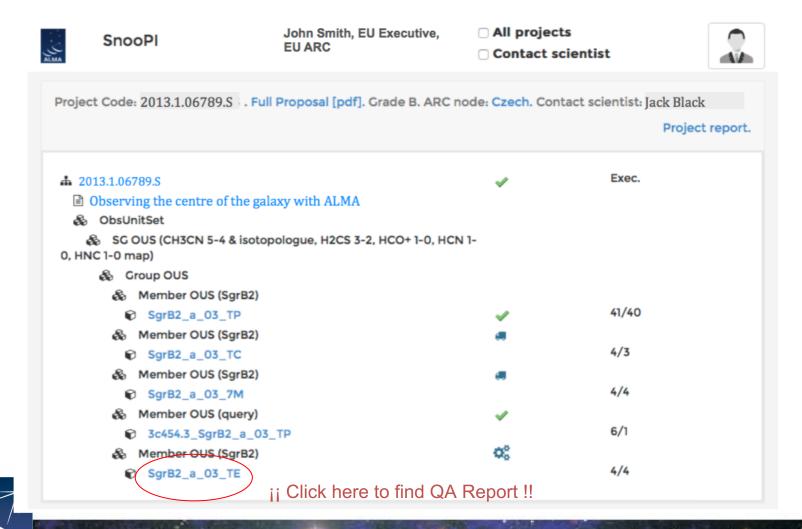
- Approved but SBs not yet prepared
- SBs prepared but are not yet in the observing queue SBs are in the observing queue but not yet taken
- Some data has been taken
- All the data has been taken
- Completed and delivered
 - Project is timed out
- Rejected at proposal review stage
- Unknown status







Single project view



Single project view

- a set of gears indicate that the MOUS is being processed;
- a smiling face shows that the MOUS are ready to be delivered;
- a truck indicates that the MOUS has been delivered

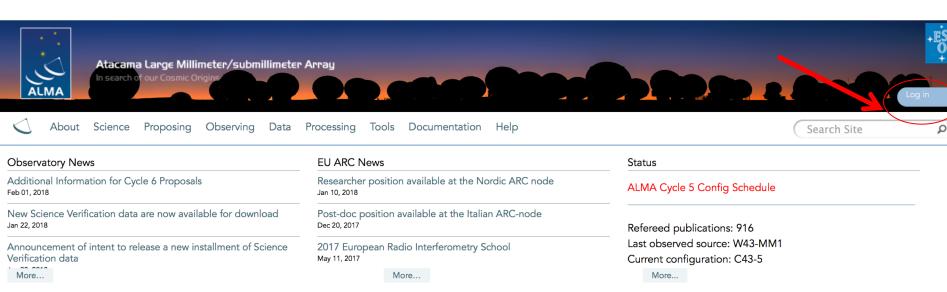


Monitor Project Status:

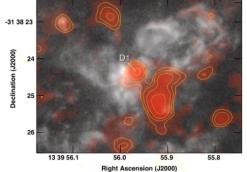
Optional Emails

- Subscribe to email notification for updates on changes to project status through your Science Portal user profile
 - **—**...
 - -Phase2Submitted
 - -Running
 - Partially Observed
 - –Fully Observed
 - —Pipeline Processing
 - **—**...
- With or without optional emails, PIs always receive notification when new data are available





Science Highlights - Molecular Gas Within the Supernebula of the Dwarf Galaxy NGC 5253



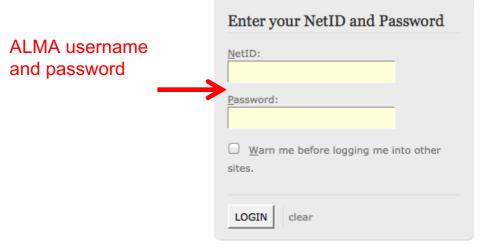
One of the areas of extragalactic research which makes great use of ALMA's resolution and sensitivity is the study of the molecular gas properties of dwarf galaxies. In a <u>recent study</u> by Dr. Jean Turner and her collaborators, they make use of Band 7 ALMA observations to detect warm ¹²CO(3-2) and ¹³CO(3-2) emission (Cloud D1) from the core of a giant star-forming region, in the dwarf galaxy NGC 5253. This "supernebula" is the source of one-third of the galaxy's infrared luminosity and is in proximity to optical clusters with measured stellar ages of ~ 1 Myr. From radio recombination line analysis, the region is estimated to have 1400-1800 O stars..

Full Summary...





ALMA Central Authentication Service (CAS)



For security reasons, please Log Out and Exit your web browser when you are done accessing services that require authentication!

If you don't have an account, you can create one in the following link: Registration web form

If you forgot you account ID, you can go to the following link: Forgot account ID page

If you want to reset your password, you can go to the following link: Reset password page

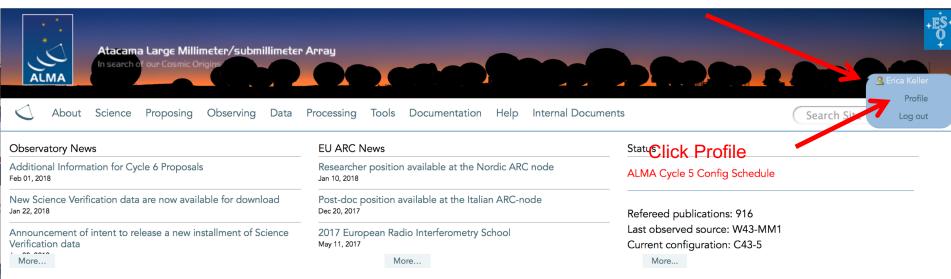
You may find a solution to your problem in the Support Center/Knowledgebase: $\underline{\mathsf{Helpdesk}}$

Copyright © 2011 Atacama Large Millimeter/submillimeter Array (ALMA). All rights reserved.

Powered by JA-SIG Central Authentication Service 3.4.10

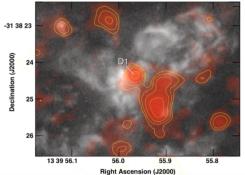


Click Name



Science Highlights - Molecular Gas Within the Supernebula of the Dwarf Galaxy NGC 5253

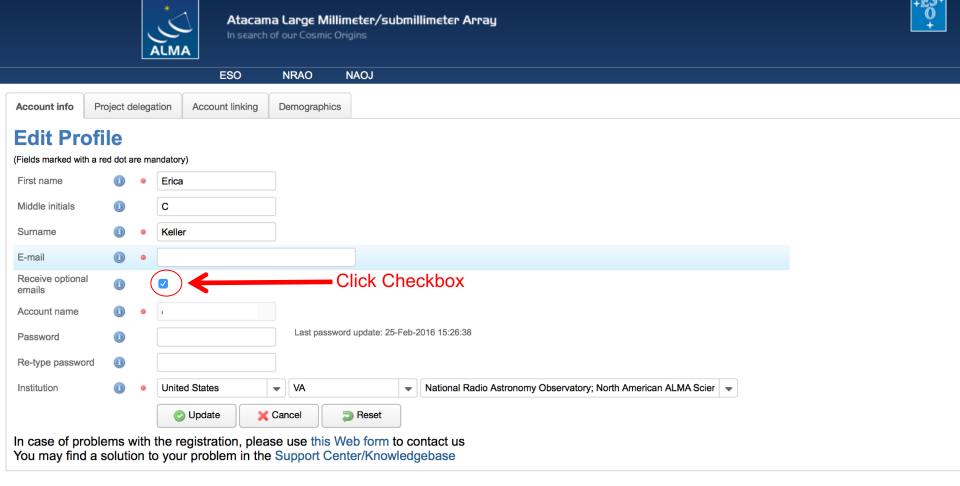
by Portal Admin — last modified Nov 30, 2017 09:38 PM



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Full Summary...





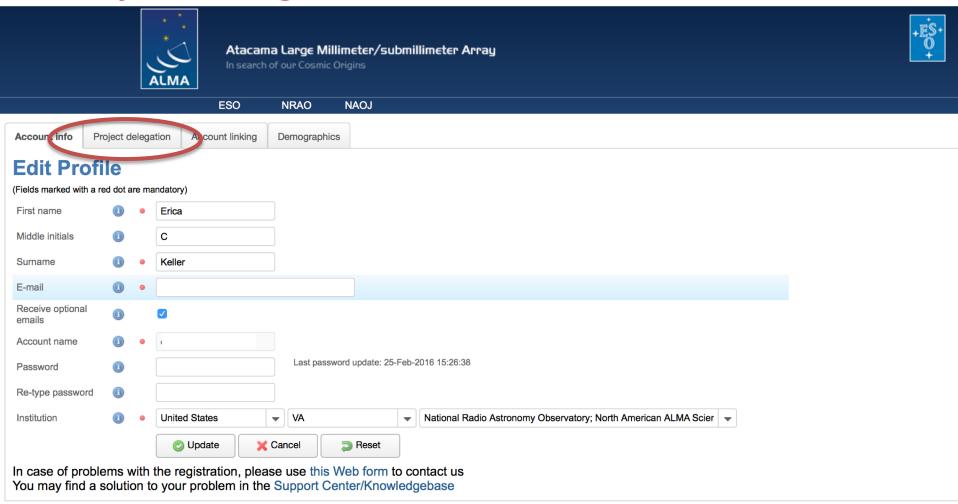




Data Delivery Email

- Sent when an individual MOUS passes QA2
- Data are ingested into the archive and made available at all Regional Centers
- This email only Sent to PI
- Included Metadata:
 - -MOUS ID, Scheduling Block (SB) name, project title
- Included Instructions:
 - Downloading data
 - Delegating access for registered ALMA users
- Included Descriptions:
 - Proprietary period
 - -Note: Triggers start of proprietary period (Usually 12 months)

Project delegation: how to



Site Map Accessibility Contact Privacy Statement



Data Delivery Email (cont'd)

- Includes Links:
 - Archive query for MOUS package
 - —Fully-calibrated MS (North America Only)
 - –CASA download and mailing lists
- Publication Requirements:
 - -ALMA acknowledgement
 - ARC specific acknowledgement
- Additional Support:
 - -Funded face-to-face reduction visits to your home ARC
 - -Contact info for your ARC Helpdesk



After untarring the processed data we have a directory tree like this:

Project code

- -- Science goal
 - -- Group OUS: Combination of member OUS's
 - -- Member OUS: May contain 12m, 7m, or Total Power observations
 - | -- DATA DELIVERY PRODUCTS

For example:

Note, the exact structure of the data delivery products (especially README) depends on cycle and dataset (i.e. pipeline or manual).

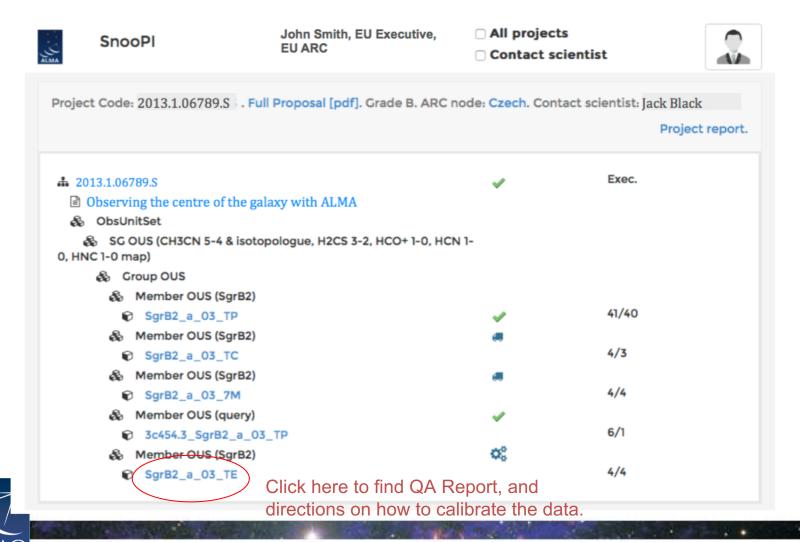


The log and product directories

```
Log of equivalent CASA commands
                                                      (non-executable)
                  A001 X1299 X39.hifa calimage.casa commands.log
    member.uid
              A001 X1299 X39.README.txt Directions to access QA comments and
member.uid
product
                                         restoration instructions. Check SnooPl.
    member.uid
                  A001 X1299 X39.SOURCE sci.spw25 27 29 31.cont.I.pb.fits
                  A001 X1299 X39.SOURCE sci.spw25 27 29 31.cont.I.pbcor.fits
    member.uid
    member.uid
                  A001 X1299 X39.SOURCE sci.spw25.cube.I.mask.fits
                  A001 X1299 X39.SOURCE sci.spw25.cube.I.pbcor.fits
    member.uid
                  A001 X1299 X39.SOURCE sci.spw25.cube.I.pb.fits.gz
    member.uid
                  A001 X1299 X39.J0117p1418 ph.spw31.mfs.I.pbcor.fits
    member.uid
                  A001 X1299 X39.J0117p1418 ph.spw31.mfs.I.pb.fits.qz
    member.uid
Calibration and Target images produced during reduction
(may be representative)
```



Remember the project view



The calibration directory

```
Contains manual flagging commands, continuum
                                     selection, flux measurements for calibrators
calibration
                   A001 X1299 X39.hifa calimage auxproducts.tgz
    member.uid
                   A001 X1299 X39.session 1.auxcaltables.tgz
    member.uid
    member.uid
                   A001 X1299 X39.session 1.caltables.tgz
                                                        Calibration tables
    uid A002 Xc8ed15 X1a9.ms.calapply.txt
                                                        generated by the
           A002 Xc8ed15 X1a9.ms.flagversions.tgz
                                                        pipeline
           A002 Xc8ed15 X1a9 target.ms\auxcalapply.txt
                       All flags will be restored during calibration
```



The weblog, and calibration scripts

(and get calibrated measurement set data

in directory called "calibrated")

```
Weblog contains plots and images from reduction and
                      imaging. Unpack this for lots of information!
   member.uid
                 A001 X1299 X39.hifa calimage.weblog.tgz
script
                 A001 X1299 X39.calimage.pipeline manifest.xml
   member.uid
                 A001 X1299 X39.calimage.product rename.txt
   member.uid
                 A001 X1299 X39.hifa calimage.casa piperestorescript.py
 -- member.uid
                 A001 X1299 X39.hifa calimage.casa pipescript.py
   member.uid
                 A001 X1299 X39.hifa calimage.pprequest.xml
   member.uid
                                                        commands to re-run
                 A001 X1299 X39 scriptForPI.py
    member.uid
                                                               the pipeline
    Run scriptForPl.py to restore calibration
```



Processed ALMA Data

Summary:

- -Check the README, access SnooPI
- –Image products are delivered to you in the "product" directory
- -Calibration/flagging files are saved in the "calibration" directory, but these you do not need to change or access directly.
- -You can run "script/scriptForPI.py" to obtain the *.ms calibrated data, for further imaging (** NA delivers the calibrated *.ms to PI)
- If the **imaging pipeline** has been run, the weblog includes the imaging steps. For **manual calibrations**, the QA information is in several files that together constitute a QA summary for the data.

increased the processing capacity of the QA2 team significantly. We refer to data processed by the imaging section of the Science Pipeline as "pipeline-imaged". If it was processed by an analyst using CASA directly without the help of the pipeline, we call it "manually imaged". In Cycle 5, ca. 60% of all science data is pipeline-imaged. The approval of the imaging products remains in the hands of the analysts and Data Reduction Managers. Note that manual imaging is sometimes performed on pipeline-calibrated datasets, occasionally in addition to pipeline-imaging. See https://almascience.org/processing/science-pipeline and chapter 13 of the Cycle 5 ALMA Technical Handbook for more details on the pipeline.

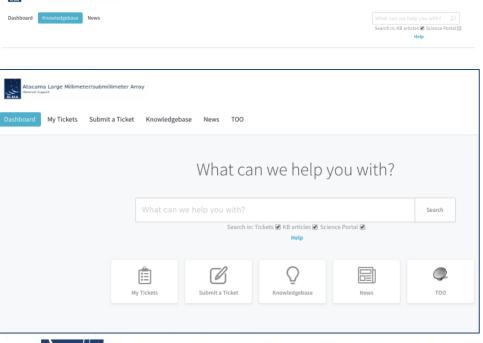
At the beginning of Cycle 4, the Science Pipeline was enabled to also perform imaging. This



Resources After Delivery

HelpDesk & Knowledgebase:

help.almascience.org



Face-to-Face visits in Charlottesville:

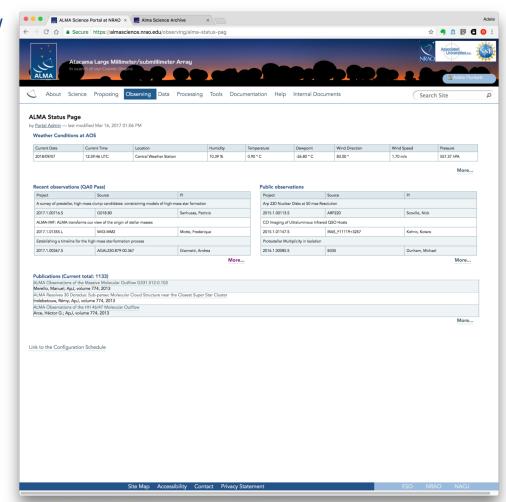
https://science.nrao.edu/facilities/alma/visitors-shortterm





Remember, you can monitor the current observations at ALMA.

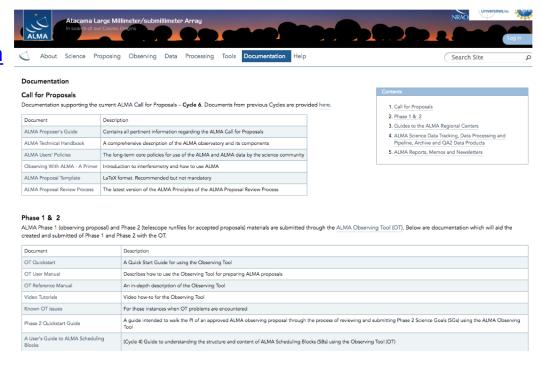
https://almascience.nrao.edu/ observing/alma-status-pag





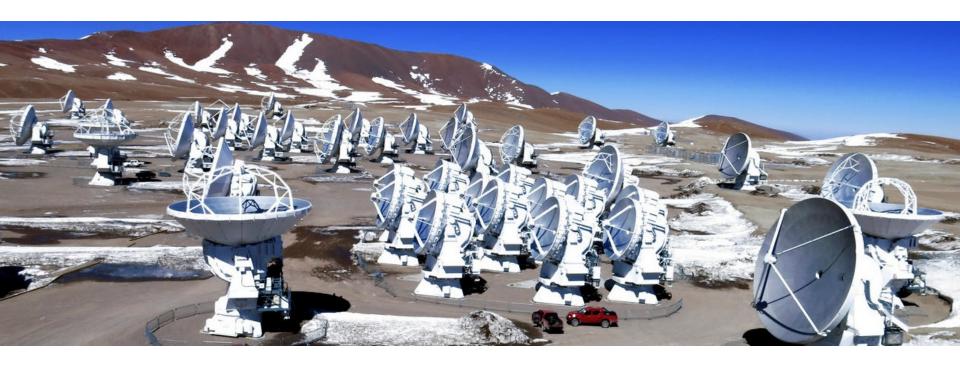
Wait, what about setting up my proposal/observations?

- Check out other ALMA Community Day events, includes Proposal Preparation Process and Strategies + ALMA Data Products: https://science.nrao.edu/facilities/alma/naasc-workshops/nrao-cd-chile18/program
- Search documentation online: https://almascience.nrao.edu/proposing/documents-and-tools





Retrieving ALMA Archival Data

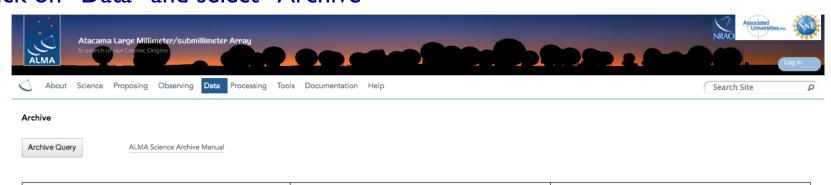


** In case of lack of time, this will be available online.



How to find the archive

Go to the science portal: https://almascience.nrao.edu
(Later, it will prompt you to log-in to download your proprietary data.)
- Click on "Data" and select "Archive"

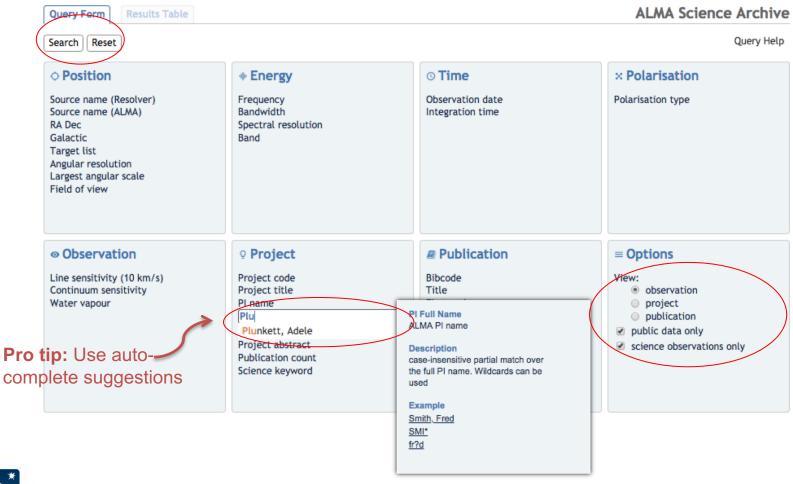


News	Did you know	Known issues
Piles from Cycle 5 onwards can be downloaded either in the usual tar-files or individually directly from the ALMA Request Handler. Click on the small black triangles in front of the tar-file names to open up the file contents and use either the "Save As" of your browser or select the files for download with the script. The 'frequency support' column now also contains the line-sensitivities of the spectral windows at native resolution and at 10km/s. The QA2 status information in the result list has been corrected 2018 June improved calculation of values (e.g. sensitivities) for ObsUnitSets that have some of their data declared QA0 SemiPass. Science Verification data (SV) searchable in the Query Interface (downloads still need to be done through the SV data page). The products for Cycle 5+ data have been split into two tar files: the tar file containing the FITS files as well as an auxiliary tar-file with the remainder of the products. Users only interested in the weblog or QA2 report or in recomputing the calibrated measurement sets do not have to download the FITS products and users who are only interested in the FITS products do not	that you can search for public but not published data; Put 0 into Mike so. that queries and downloads can be done also from a script with astroquery? Cycle 5+ files can now be downloaded individually? that a full description of the possibilities is available in the ALMA Science Archive Manual?	Report a <u>new issue</u> .
Site Map Accessibil	ity Contact Privacy Statement	ESO NRAO NAOJ



Find data in archive:

Archive Query http://almascience.nrao.edu/aq/





Archive Query

Query Form Results Table Results Bookmark Export Table Results Help Submit download request Showing 30 rows (30 before filtering). More columns RA Integration Release date -**Velocity resolution Project code** Source name Dec Frequency support Filter: + m/s ☑ 2012.1.00090.S S2CLS UDS110 7 9.326 27236.96 02:18:48.44 -05:18:05.0 2014-11-07T09:35:00.000 ~ 7 2012.1.00090.S S2CLS UDS156 02:18:24.23 -05:22:53.4 8.836 2014-11-07T09:35:00.000 27236.96 ⋖ 2012.1.00090.S S2CLS UDS160 02:18:23.86 -05:11:36.2 7 8.842 27236.96 2014-11-07T09:35:00.000 2012.1.00090.S S2CLS UDS168 02:18:20.34 -05:31:41.6 7 8.843 2014-11-07T09:35:00.000 27236.96 336.00..351.99GHz ☑ 7 2012.1.00090.S S2CLS UDS199 02:18:07.38 -04:44:11.7 8.812 2014-11-07T09:35:00.000 27236.96 2012.1.00090.S S2CLS UDS204 02:18:03.01 -05:28:39.8 7 8.873 2014-11-07T09:35:00.000 27236.96 336.00..351.99GHz 2012.1.00090.S S2CLS_UDS216 -04:52:39.6 7 8.82 27236.96 02:17:56.80 2014-11-07T09:35:00.000 336.00..351.99GHz 2012.1.00090.S S2CLS_UDS252 -05:20:10.2 7 27236.96 02:17:37.79 8.827 2014-11-07T09:35:00.000 336.00..351.99GHz 2012.1.00090.S S2CLS_UDS286 -05:25:36.5 7 9.657 2014-11-07T09:35:00.000 27236.96 02:17:25.76 336.00..351.99GHz 2012.1.00090.S S2CLS UDS292 02:17:21.85 -05:19:03.3 7 8.815 2014-11-07T09:35:00.000 27236.96 336.00..351.99GHz 2012.1.00090.S S2CLS_UDS298 02:17:19.90 -05:09:36.4 7 9.55 2014-11-07T09:35:00.000 27236.96 336.00..351.99GHz 2012.1.00090.S S2CLS UDS334 02:17:02.81 -04:57:24.9 7 8.856 2014-11-07T09:35:00.000 27236.96 336.00..351.99GHz 2012.1.00090.S S2CLS UDS408 02:16:22.59 -05:11:06.0 7 8.819 27236.96 2014-11-07T09:35:00.000 336.00..351.99GHz 2012.1.00090.S S2CLS UDS421 02:16:17.62 -05:09:02.0 7 8.803 2014-11-07T09:35:00.000 27236.96 336.00..351.99GHz 2012.1.00090.S S2CLS UDS47 02:19:24.97 -05:09:19.9 7 8.785 2014-11-07T09:35:00.000 27236.96 336.00..351.99GHz



Archive Query: more columns

•	Project code	
Filter:		
	2012.1.00090.S	

and boolean OR expressions (" "), can be used. RA deg Right Ascension of the field pointing. Dec deg Declination of the field pointing. ALMA receiver band. Integration s Aggregated integration time for the field in the ASDM. Release date Velocity resolution m/s Estimated velocity resolution from all the spectral windows, from frequency resolution. Frequency support GHz All frequency ranges used by the field Spatial resolution Frequency resolution with Estimated frequency resolution from all the spectral windows, using median values of channel widths. Pol products Polarisation products provided. Observation date PI name case-insensitive partial match over the full PI name. Wildcards can be used PWV mm Estimated precipitable water vapour from the XML_CALWVR_ENTITIES table. MEMBER_OUSS_ID generating this ASDM. Asdm uid UID of the ASDM containing this Field.	Project code		Project code, in the form YYYY.NNNNN.C.AAA, where:
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PWV mm Estimated precipitable water vapour from the XML_CALWVR_ENTITIES table. Member ous id MEMBER_OUSS_ID generating this ASDM. UID of the ASDM containing this Field.	Spatial resolution Frequency resolution		Estimated frequency resolution from all the spectral windows, using median values of channel widths.
Member ous id MEMBER_OUSS_ID generating this ASDM. UID of the ASDM containing this Field.	Spatial resolution Frequency resolution Pol products		Estimated frequency resolution from all the spectral windows, using median values of channel widths.
Asdm uid UID of the ASDM containing this Field.	Spatial resolution Frequency resolution Pol products Observation date		Estimated frequency resolution from all the spectral windows, using median values of channel widths. Polarisation products provided.
	Spatial resolution Frequency resolution Pol products Observation date PI name	kHz	Estimated frequency resolution from all the spectral windows, using median values of channel widths. Polarisation products provided. case-insensitive partial match over the full PI name. Wildcards can be used
Project title Case-insensitive search over the project title	Spatial resolution Frequency resolution Pol products Observation date PI name PWV	kHz	Estimated frequency resolution from all the spectral windows, using median values of channel widths. Polarisation products provided. case-insensitive partial match over the full PI name. Wildcards can be used Estimated precipitable water vapour from the XML_CALWVR_ENTITIES table.
	Spatial resolution Frequency resolution Pol products Observation date Pl name PWV Member ous id	kHz	Estimated frequency resolution from all the spectral windows, using median values of channel widths. Polarisation products provided. case-insensitive partial match over the full PI name. Wildcards can be used Estimated precipitable water vapour from the XML_CALWVR_ENTITIES table. MEMBER_OUSS_ID generating this ASDM.
Project type	Spatial resolution Frequency resolution Pol products Observation date Pl name PWV Member ous id	kHz	Estimated frequency resolution from all the spectral windows, using median values of channel widths. Polarisation products provided. case-insensitive partial match over the full PI name. Wildcards can be used Estimated precipitable water vapour from the XML_CALWVR_ENTITIES table. MEMBER_OUSS_ID generating this ASDM.
Troject type	Spatial resolution Frequency resolution Pol products Observation date PI name PWV Member ous id Asdm uid	kHz	Estimated frequency resolution from all the spectral windows, using median values of channel widths. Polarisation products provided. case-insensitive partial match over the full PI name. Wildcards can be used Estimated precipitable water vapour from the XML_CALWVR_ENTITIES table. MEMBER_OUSS_ID generating this ASDM. UID of the ASDM containing this Field.

Frequency support
335.99351.99GHz

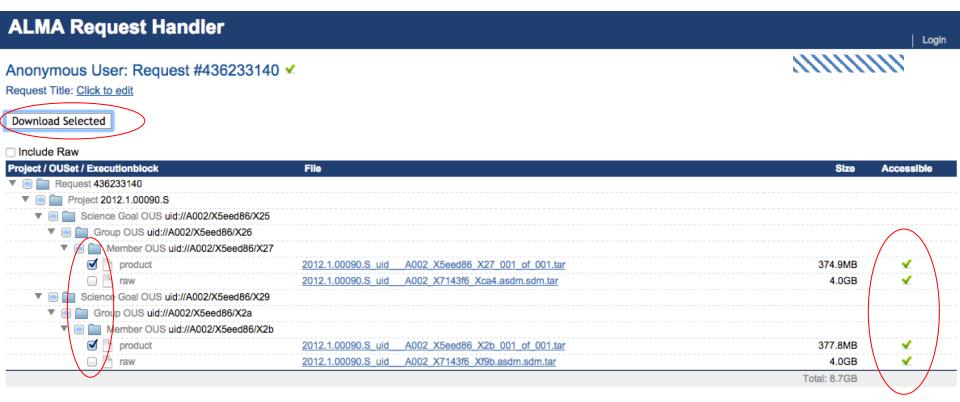


Downloading the data: Request Handler

- All data downloaded as tar files
- Large data sets may be broken into several pieces
 - -Name is [project_code]_[OUS_ID]_m_of_n.tar
 - Raw data packaged as one tar file per execution block (EB)
 - name is [project_code]_[EB_ID].asdm.sdm.tar
- For Cycle 0-5 projects, cannot directly download individual data products but potentially coming in Cycle 6...
 - —FITS images
 - -Diagnostic plots, etc.



Request Handler





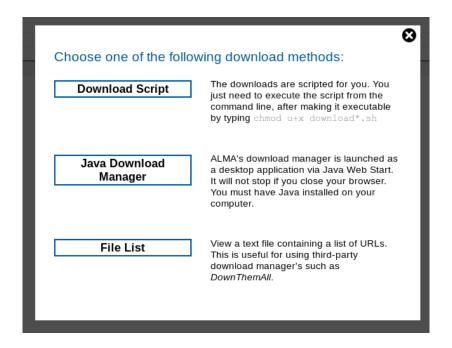
Request Handler prompts an email

From do-not-reply@nrao.edu 😭 Reply Reply Reply All Forward Archive Junk Delete Subject ALMA Archive at NRAO: Request 223292105 1:24 PM Other Actions Thank you for using the ALMA archive. Your data selection (4.3GB) is available from this link https://almascience.nrao.edu/rh/requests/nbrunett/223292105 We hope they meet your expectations and will lead to a successful completion of your scientific program. Publications making use of these data must include the following statement in the acknowledgment: "This paper makes use of the following ALMA data: ADS/JAO.ALMA#2012.1.00090.S. ALMA is a partnership of ESO (representing its member states), NSF (USA) and NINS (Japan), together with NRC (Canada) and NSC and ASIAA (Taiwan), in cooperation with the Republic of Chile. The Joint ALMA Observatory is operated by ESO, AUI/NRAO and NAOJ." Please submit your requests for help, for a visit to the ARC, or to report any problems discovered in your data through the ALMA Helpdesk at https://help.almascience.org. Best regards, The North American ALMA Archive at the NAASC Summary: Files available: 2 (4.3GB) Files under proprietary period: 0 (-) Files not available: 0 (-) Details: Files available: - 2012.1.00090.S_uid___A002_X5eed86_X2b_001_of_001.tar : AUTHORIZED 2012.1.00090.S uid A002 X7143f6 Xf9b.asdm.sdm.tar : AUTHORIZED



Files under proprietary period:

Request Handler: Download options

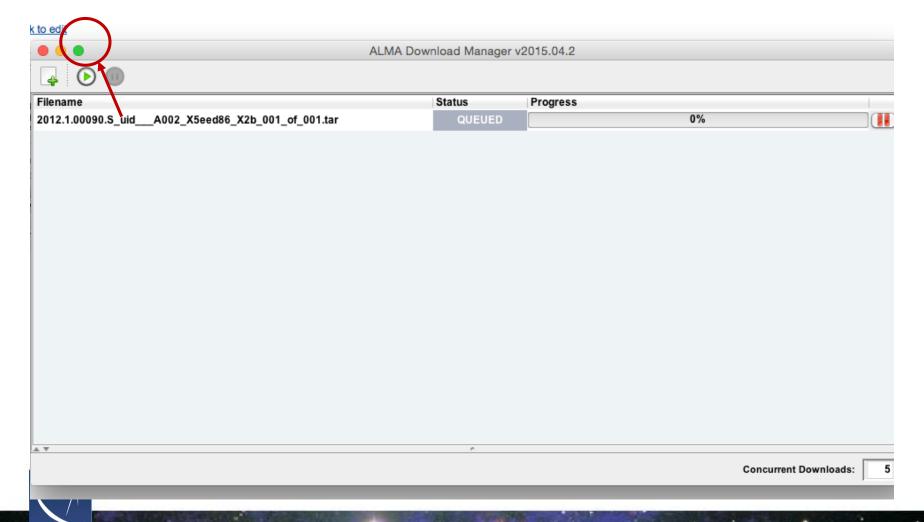




Request Handler: script



Request Handler: Java Download Manager



Resources related to the Archive

Check the science portal for possible maintenance message. https://almascience.nrao.edu

There are 3 versions of the ALMA archive. If one is down, it is possible I of the other two are available.

- NRAO: http://almascience.nrao.edu/aq/
- ESO: http://almascience.eso.org/aq/
- NAOJ: <u>almascience.nao.ac.jp/aq/</u>

Contact your local helpdesk and provide:

- Project ID
- SBname
- ASDM
- What method you are using to download?







For more info:

https://almascience.nrao.edu/

The Atacama Large Millimeter/submillimeter Array (ALMA), an international astronomy facility, is a partnership of Europe, North America and East Asia in cooperation with the Republic of Chile. ALMA is funded in Europe by the European Organization for Astronomical Research in the Southern Hemisphere (ESO), in North America by the U.S. National Science Foundation (NSF) in cooperation with the National Research Council of Canada (NRC) and the National Science Council of Taiwan (NSC), and in East Asia by the National Institutes of Natural Sciences (NINS) of Japan in cooperation with the Academia Sinica (AS) in Taiwan. ALMA construction and operations are led on behalf of Europe by ESO, on behalf of North America by the National Radio Astronomy Observatory (NRAO), which is managed by Associated Universities, Inc. (AUI), and on behalf of East Asia by the National Astronomical Observatory of Japan (NAOJ). The Joint ALMA Observatory (JAO) provides the unified leadership and management of the construction and operation of ALMA.





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