

Archive Access / Data Inspection Post-Observing, Pre-Calibration

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### Outline

- After the observations:
  - Obtaining your data from the archive
    - Which CPU processes the data? (Home or NRAO)
  - Examine your data
    - Structure and potential issues: all expected data present, RFI, calibrators, reference antenna...
  - Prepare for calibration steps
    - Use examination to flag bad data upfront
       Better preparation eases the process!



# Assumptions (for all these lectures)

This presentation assumes that you are familiar with **the basics of**:

- radio interferometry
- flux density calibration, *antenna-based* calibration (complex gain, bandpass), and self-calibration
- imaging and deconvolution

For references on the above, please check: (these 5 are web links...)

- The lectures of the synthesis imaging workshops: <u>https://science.nrao.edu/science/meetings/2014/14th-synthesis-imaging-workshop/</u> <u>https://science.nrao.edu/science/meetings/2016/15th-synthesis-imaging-workshop/</u> <u>https://science.nrao.edu/science/meetings/2018/16th-synthesis-imaging-workshop/</u>
- <u>Synthesis Imaging for Radio Astronomy II</u> (eds. Taylor, Carilli, and Perley).
- Interferometry and Synthesis in Radio Astronomy (by Thompson, Moran, and Swenson).



## **NRAO versus Local/home computing**

- Note that NRAO offers computing facilities for demanding projects upon request
  - Registered user (portal <u>https://my.nrao.edu</u>)
  - Limited capacity, compete with others, no guarantee
  - See computing policy page <u>https://info.nrao.edu/computing/guide/cluster-processing/</u>
- Here assume **processing at home institute** 
  - Data transfer over internet (up to couple of 100 GB)
  - Data shipped on disk (purchase, up to 1.8 TB/disk)



## **Observing operator logs are available**

### **Sent by email to proposers directly after observation** Stored on web servers:

#### http://www.vla.nrao.edu/cgi-bin/oplogs.cgi

👻 VLA Operators Logs - Mozilla Firefox 💦 📃 🗙		
VLA Operators Logs × +		
②	Date	Tir
National Radio Search NRAO Astronomy Observatory NRAO Home VLA - Tools for Array Operators > Operator Logs	2018-10-04	13:
The array operator logs are being merged with the new	2018-10-04	13:
e2e archive system, this lookup tool has been provided as a stopgap measure to tide us over until the work is done. The logs here are from October 2003 onwards,	2018-10-04	12:
older logs can be found using the previous tool, <u>here</u> .	2018-10-04	09:
To use this tool, select the range of dates you wish to see logs for and hit the `Show Logs' button. By default it will display logs for the last week.	2018-10-04	08:
Start Year 2018 V Month Oct V Day 04 V	2018-10-04	05:
Stop Year 2018 - Month Oct - Day 04 -	2018-10-04	04:
Show Logs	2018-10-04	03:
Staff   Contact Us   Careers   Directories   Site Map   Help   Policies   Diversity   Search	2018-10-04	03:
Copyright © 2009 Associated Universities, Inc. The National Radio Astronomy Observatory is a facility of the National Science Foundation operated under cooperative agreement by Associated Universities, Inc.	2018-10-04	02:

Date	Time	Code	File
2018-10-04	13:57	SOFTWARE	<u>pdf</u>
2018-10-04	13:21	18A-498	<u>pdf</u>
2018-10-04	12:03	18A-342	<u>pdf</u>
2018-10-04	09:32	18A-131	<u>pdf</u>
2018-10-04	08:32	18A-389	<u>pdf</u>
2018-10-04	05:41	TDRW0001	$\operatorname{pdf}$
2018-10-04	04:59	18A-146	<u>pdf</u>
2018-10-04	03:41	18A-342	<u>pdf</u>
2018-10-04	03:12	STARTUP	<u>pdf</u>
2018-10-04	02:12	18A-354	<u>pdf</u>

NRAO

DS

### **Observing operator logs are available**

### **Sent by email to proposers directly after observation** Stored on web servers:

http://www.vla.nrao.edu/cgi-bin/oplogs.cgi

VLA OBS	ERVING	LOG			2018	3-10-04	_0541	L_TD	RW0001	L							
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Secons sioned:	27	SBID(s):	35624494										intermittent2		1	1 1	
		Source File(s):	TDRW0001_sb39	5524494_1_1									If possible, a spectrum of the RPI should be includ	ed in the e-mail	-		
		Observer E-mail:	emornijan@mao	.edu									Thanks very much for your support; this information			+ +	
		Operator(s):	Kenneth Gibson			· · · · · · · · · · · · · · · · · · ·							updated on the EVLA science pages at:	will be contributivy	-		
													https://sdence.nrao.edu/facilities/via/docs/manuals/	obstauluie (mondes.)cfl/	-		
		ttp://www.via.wae.ec								ł	040ct 5:41:30		Antennas in the D-array may be shadowed at low el				
salvitty data is updat	of each a sector (ar) or	relidelight and is availa	the from the online a	archive at the	parchive.mag.ed						0.001 0.000		occurs, sensitivity will be affected.				
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											01001 3:11:30	0100 0:32:31	C-band receiver warm for cold head replacement	CKHOdena	1991	0.05	3/0
				-							040ct 5:57:00		Your new operator(s) Is(are): Sam Gimore				
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# **Obtaining data from the NRAO archive** The current archive tool and the new archive tool

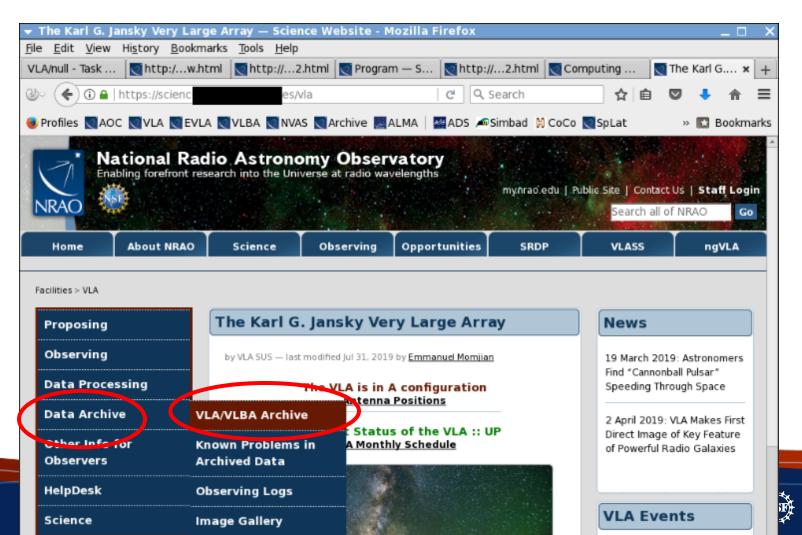
(improvements in progress)



### The NRAO Data Archive Tool

### https://science.nrao.edu/facilities/vla/

### → Data Archive (left menu), VLA/VLBA Archive



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### → Data Archive (left menu), VLA/VLBA Archive

Facilities > VLA > Data Archive > VLA Data Archive

#### Proposing Observing Data Processing Data Archive Other Info for Observers

HelpDesk

Science

#### VLA Data Archive

by <u>Castaal Van Moorsel</u> – last redified Feb 25, 2021 by <u>Frank Schinzel</u> Locating and unlocking data

Newly observed data are transferred to the archive and available for retrieval to those with the appropriate privileges approximately 10

minutes after the end of the observations. The archive content can be

generation AAT. On request, NRAO will ship data on physical hard disks,

accessed via the "legacy" Archive Access Tool (AAT) or via a next

subject to the conditions of our Data Shipment Policy. NOTE: The

NRAO has temporarily suspended the option to ship data on hard

#### Contents

- 1. <u>Locating and unlocking</u> <u>data</u>
- 2. <u>Data Formats and Data</u> <u>Retrieval</u>
- 3. New Archive Access Tool
- <u>Creating UVFITS</u> Formatted Files

Instructions to create UVFITS from the SDM or CASA format are given below.

#### **New Archive Access Tool**

disks during the COVID-19 pandemic.

NRAO continues to recommend using the "<u>legacy archive</u>" to retrieve data from all VLA observations. However, pipeline and SRDP data products can only be accessed through the <u>new archive access tool</u>.

The NRAO is currently developing a new portal for accessing data from VLA and VLBA observations. Observers are welcome to try using the "<u>new archive</u>" to retrieve their VLA data.

The <u>new archive</u> will eventually become the only tool for accessing all VLA and VLBA data. Any observers who decide to try using the new archive while it is under development and find problems are asked to try the legacy archive and report their issues to the NRAO via the <u>helpdesk</u>.

Current known limitations/features include:

All visitor computing account holders can no longer request data directly to be deposited in their

# https://data.nrao.edu

	VRAO	ALL	tio Astronomy Observatory search into the Universe at radio wavelengths			
Arc	hive Acce			( • <b>)</b>	Log In 🛛 🖉 Le	gacy Archive 🔒 🕈 About
Q						
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View	Projecta	View Observations	View Images Page 1 •		S	how 25 of 8251 Projects
	11 Project	11 Instrument	Title	11 First Obs	It Last Obs	
+	BF119	VLBA	An Optimal Search Strategy For FRBs	2016-08-18 20:51	2017-06-05 16:34	274 execution blocks
+	UF001	VLBA	No title found	2017-01-16 23:28	2017-10-22 13:48	20 execution blocks
+	BD192	VLBA	MSPSRPI: VLBI astrometry of millisecond pulsars	2016-08-03 00:09	2018-01-09 01:07	415 execution blocks
+	DQ718	VLBA	No title found	2017-05-04 07:45	2017-05-04 09:11	1 execution blocks
+	BT139	VLBA	Magnetic Fields and Exoplanets in AGB Environments: Deep Observations with H	SA 2017-05-21 20:15	2017-07-09 02:05	4 execution blocks
+	BB379	VLBA	SN 2016coi: A Rare Chance to Resolve a Nearby Supernova	2017-05-22 06:30	2017-05-22 14:29	1 execution blocks
+	UD001	VLBA	No title found	2017-01-08 23:57	2018-07-23 10:35	24 execution blocks
+	TR031	VLBA	No title found	2017-04-26 18:00	2017-09-27 20:35	(5 execution blocks)
+	TC026	VLBA	No tille found	2017-05-25 13:00	2017-05-25 13:39	1 execution blocks
+	DQ720	VLBA	No title found	2017-05-18 06:30	2017-05-18.07:55	(1 execution blocks)
+	BP221	VLBA	Mrk1018: Binary black hole interaction of transient jet ejection?	2017-05-12 16:19	2018-03-02 00:03	12 execution blocks



# https://data.nrao.edu

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TDRW000	.sb35624494.eb35628826.58395.23	23719237269 TDRW0001	VLA 18-10	0-04 05:41:34	18-10-04 08:32:43	12.446 GB D	J	C, S	visibility		25	
TDRW000	.sb35624494.eb35625702.58394.22	22234046296 TDRW0001	VLA 18-10	0-03 05:22:58	18-10-03 05:54:22	567.555 MB D		C, S	visibını,		8	

### https://data.nrao.edu



#### Observation ID: TDRW0001.sb35770743.eb35774735.58429.44719293981

Obs ID: TDRW0001.sb35770743.eb35774735.58429.44719293981 Project Code: TDRW0001 Estimated Size: 15.927 GB Obs Release Date: 2018-11-07T13:43:32.001Z Data Product Type: visibility Receiver Band: Ka, X Array Configuration: D



Longitude	Latitude	Target Name	Min Frequency	Max Frequency	Scan Intent	Polarizations	Temporal Res
13h31m8.288s	30°30'32.959"	1331+305=3C286	8.3320000 GHz	8.4600000 GHz	["CALIBRATE_POINTING"]	["RR, RL, LR, LL"]	1.044
9h54m56.824s	17°43'31.222"	J0954+1743	8.3320000 GHz	8.4600000 GHz	["CALIBRATE_POINTING"]	["RR, RL, LR, LL"]	1.053
9h47m57.382s	13°16'40.660"	IRC+10216	36.3115447 GHz	36.8677209 GHz	["OBSERVE_TARGET"]	["RR, RL, LR, LL"]	3.053
9h47m57.382s	13°16'40.660"	IRC+10216	36.3115447 GHz	36.8677209 GHz	["OBSERVE_TARGET"]	["RR, RL, LR, LL"]	3.022
9h54m56.824s	17°43'31.222"	J0954+1743	36.3115447 GHz	36.8677209 GHz	["SYSTEM_CONFIGURATION"]	["RR, RL, LR, LL"]	3.322
9h47m57.382s	13°16'40.660"	IRC+10216	36.3115447 GHz	36.8677209 GHz	["OBSERVE_TARGET"]	["RR, RL, LR, LL"]	3.022
9h54m56.824s	17°43'31.222"	J0954+1743	36.3115447 GHz	36.8677209 GHz	["CALIBRATE_AMPLI","CALIBRATE_PHASE"]	["RR, RL, LR, LL"]	3.095
9h47m57.382s	13°16'40.660"	IRC+10216	36.3115447 GHz	36.8677209 GHz	["OBSERVE_TARGET"]	["RR, RL, LR, LL"]	3.022
9h47m57.382s	13°16'40.660"	IRC+10216	36.3115447 GHz	36.8677209 GHz	["OBSERVE_TARGET"]	["RR, RL, LR, LL"]	3.053
9h54m56.824s	17°43'31.222"	J0954+1743	36.3115447 GHz	36.8677209 GHz	["CALIBRATE_AMPLI","CALIBRATE_PHASE"]	["RR, RL, LR, LL"]	3.095
9h54m56.824s	17°43'31.222"	J0954+1743	36.3115447 GHz	36.8677209 GHz	["CALIBRATE_AMPLI","CALIBRATE_PHASE"]	["RR, RL, LR, LL"]	3.097
9h47m57.382s	13°16'40.660"	IRC+10216	36.3115447 GHz	36.8677209 GHz	["OBSERVE_TARGET"]	["RR, RL, LR, LL"]	3.053

#### https://data.nrao.edu



Project Code:	Archive Filename:	PI Name:	Title Text:
tdrw0001			
Abstract Text:			
	vive F		
	Launch Workflow Task o	on: TBRWoodi	×
	User Email (required):	Isjouwer@nrao.edu	
View Projects View Observations View Images	Request Description:	VLA Processing Request	
11 Project 11 Instrument			
- TDRW0001 VLA	Destination Directory:	□ Specify directory (must be logged in & s	taff)
Title: No title found Abstract: No abstract found		/lustre/	
PI: Emmanuel Momjian Co-Authors: Frank Schinzel, Emmanuel Momjian	<b>—</b>		
Observations Images	Create tar file:	Return results as a tar file	
	Choose download data	<ul> <li>SDM taples only (metadata only)</li> </ul>	
	form <mark>:</mark> t:	SDM-BDF dataset (metadata + visibilitie)	es)
■ 0/10: selected 0/10.0 TB		⊖ Basic Neasurement Set (uncalibrated)	
♥ View Selection(s) ★ Clear All ♥ [	Downle	O Calibrated Measurement Set	
It Archive File	Apply telescope flags:	Apply flags generated during observing	
Add to elipboard	Scan Intents:	Click to Select	59
TDRW0001.sb35624494.eb35628826.58395.23			25
TDRW0001.sb35624494.eb35625702.58394.22	CASA Version:	5.4.2-8 (recommended) <del>-</del>	8
	TD	Cancel 🥄	Submit Request 🕗
	ТВ)		



## Science Data Model

- SDM native VLA data format
  - Header tables (XML, meta)
  - Scan data (BDF, binary)
- CASA Measurement Set (MS)
  - Basic (SDM  $\implies$  MS, with flags?) formation
    - Online, Shadow, Zero
  - Calibrated (VLA CASA pipeline)
    - CASA version dependent !
- SDM/MS are data directories
  - "tar" for downloading over internet (but requires twice the disk space)
  - Alternatively, use "wget"

User Email (required):	1
Request Description:	EVLA Processing Request
Destination Directory:	Specify directory (must be logged in & staff)
	/lustre/
Create tar file:	Return results as a tar file
Choose download data	SDM tables only (metadata only)
gs:) format:	SDM-BDF dataset (metadata + visibilities)
	Basic Measurement Set (uncalibrated)
	Calibrated Measurement Set
Apply telescope flags:	Apply flage generated during observing
CASA Pipeline Version:	6.1.2-7   2020.1 0.36 (recommended) -
Restore previous CMS	19B-225_2020_01_07_T17_01_05.711.tar+
internet	Estimated Processing Time: 1
	Request Description: Destination Directory: Create tar file: Choose download data format:

Launch Workflow Task on: 19B-225



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### **The Archive Tool**

### https://archive.nrao.edu/

National Radio Astronomy Observatory	Log in for proprietary data here	Unlock my data : Legin to My NRAO.edu   Logoff
Archive Ho	me   Basic Search   Advanced Search   Image Search   Description   Archive Policy   Archive Status   Archive Tools   Future Goals   VLA Images   VLBA Sources   Downloads   In order to unlock your proprietary data and have access to other archive tools, you must log in to your My.NRAO account.	Hard Disks
	NRAO Science Data Archive : Advanced Search Tool Historical VLA, Jansky VLA, VLBA and GBT Data Products	
	Submit Query Check Query Clear Form	
	Output Control Parameters :         Choose Query Return Type :         © Download Archive Data Files         O'LA Observations Summary         List of Observation Scans         List of Projects         General Search Parameters :         Telescapes € AllJansky VLAHistorican VLAVLBAGBT         Information Control Parameters :         GBP ACGRT12A_055         Project Session         Dates From         Observer Name         Archive File ID         (partial strings allowed)         (2010-06-21 14:20:30)	
	Position Search : Target Name         Search Type SIMBAD or NED \$ Min. Exposure         RA or Longitude       DEC or Latitude       Equinox 12000 \$         (04h33m11.1s or 68.29d)       DEC or Latitude       Equinox 12000 \$         Search Radius 1.0' (1d0000' or 0.2d)       - OR -       Check for automatic VLA field-of-view, freq. dependent.??         Observing Configurations Search : Telescope All A AB BnA B BC CnB Config C CD DnC D DA       Observing Bands X U K Ka Q W         Sub_array All 1 2 3 4 5       Frequency Range (In MiHz : 1665.401 - 1720.500)         Polarization ALL \$       Polarization ALL \$	



### **Query return**

- For each match, the archive query return presents per observation (i.e. per row):
  - The observing run identifier (i.e., the SB name)
  - Any data quality issues (highlighted in yellow/red)
  - The SDM-BDF set (content of the SDM directory)
  - The individual scans with their details
  - The operator log (usually, also sent by email)

												· · · · · · · · · · · · · · · · · · ·	
Archive File	Status	Project	Seg	Obs. Data Starts	Obs. Data Stops	File Size	Telescope: config:sub"	Bands	Format	Туре	Data Qual	View Scans	.ogs etc.
11A-291.sb4911125.eb4924302.55782.00136674769	public	11A-291	x	11-Aug-09 00:02:01	11-Aug-09 01:01:45	42.46GB	VLA:A:0	L	SDMset	raw	ок	Scans	Logs
11A-291.sb4911125.eb4944094.55784.99251239583	public	11A-291	x	11-Aug-11 23:50:07	11-Aug-13 02:14:44	30.29GB	VLA:A:0	L	SDMset	raw	ок	Scans	Logs
11A-291.sb4910900.eb4947827.55787.6933925	public	11A-291	x	11-Aug-14 16:39:27	11-Aug-14 18:39:07	78.96GB	VLA:A:0	L	SDMset	raw	info	<u>Scans</u>	Logs
11A-291_sb4911125_2.55795.922649976856	public	11A-291	x	11-Aug-22 22:08:44	11-Aug-22 23:08:30	36.44GB	VLA:A:0	L	SDMset	raw	ок	Scans	Logs
11A-291_sb4911125_3_000.55804.894766516205	public	11A-291	x	11-Aug-31 21:28:29	11-Aug-31 22:28:18	39.47GB	VLA:A:0	L	SDMset	raw	ОК	<u>Scans</u>	Logs





### Scan listing:

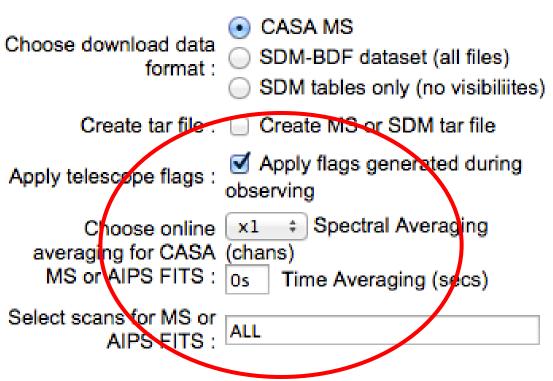
### Scan details (source, date, setup, etc)

				-						_									
Project	Scan	Source	Cal	Start Time	Stop Time	Sys	тоя	Intrvl	Scan	Spect	Obs_Freq		Polar	-		Tele:config	RA(J2000)	DEC(J2000)	Archive File
J. J.	:sub		Code		<b>r</b>	-0-	(sec)	(sec)	Intent	Win	(MHz)	(MHz)		chans	Mode	:sub:nants			
11A-291		J1120+1420		11-Aug-09 00:02:01	11-Aug-09 00:02:54	UTC	53.5		OBS	CD_0:SW_0 CD_0:SW_1 CD_0:SW_2 CD_0:SW_3 CD_0:SW_4 CD_0:SW_5 CD_0:SW_6 CD_0:SW_6 CD_0:SW_7 CD_0:SW_8 CD_0:SW_9 CD_0:SW_10 CD_0:SW_11 CD_0:SW_11 CD_0:SW_11 CD_0:SW_13 CD_0:SW_14	998.00000 1062.000000 1126.00000 1126.00000 1254.00000 1318.00000 1318.00000 1506.00000 1570.00000 1634.00000 1634.00000 1634.00000 1826.000000 1826.000000	64.000 64.000 64.000 64.000 64.000 64.000 64.000 64.000 64.000	RR,LL RR,LL RR,LL RR,LL RR,LL RR,LL RR,LL RR,LL RR,LL RR,LL RR,LL RR,LL	128 128 128 128 128 128 128 128 128 128	WIDR WIDR WIDR WIDR WIDR WIDR WIDR WIDR	EVLA:A:1:27	11h20m27.807s	+14d20754.99*	11А-291.sb4911125.eb4924302.55782.00136674769 uidevla_bdf_1312848123251.bdf
114-291	2:1	J1120+1420		11-Aug-09 00:02:54	11-Анg-09 00:03:54	urc	59.8	1	CAL	CD_0:SW_0 CD_0:SW_1 CD_0:SW_2 CD_0:SW_3 CD_0:SW_4 CD_0:SW_4 CD_0:SW_5 CD_0:SW_6 CD_0:SW_7 CD_0:SW_8	998.00000 1062.00000 1126.00000 1254.00000 1318.00000 1382.00000 1382.00000 1506.00000 1570.00000 1634.00000 1634.00000 1698.00000 1826.000000 1890.00000	64.000 64.000 64.000 64.000	RR,LL RR,LL RR,LL RR,LL RR,LL RR,LL RR,LL RR,LL RR,LL RR,LL	128 128 128 128 128 128 128 128 128 128	WIDR WIDR WIDR WIDR WIDR WIDR WIDR WIDR	EVLA:A:1:27	11h20m27.807s	+14d20/54.99*	11A-291.sb4911125.eb4924302.55782.00136674769 uidevla_bdf_1312848123257.bdf
114-291	3:1	J1120+1420		11-Aug-09 00:03:54	11-Aug-09 00:05:24	итс	89.8	1	CAL	CD_0:SW_7 CD_0:SW_8 CD_0:SW_9 CD_0:SW_10 CD_0:SW_11 CD_0:SW_12 CD_0:SW_13 CD_0:SW_14	1126.00000 1190.00000 1254.00000 1318.00000 1382.00000 1546.00000 1570.00000 1634.00000 1638.00000 1658.00000 1826.00000	64.000 64.000 64.000 64.000 64.000 64.000 64.000 64.000 64.000 64.000 64.000	RR,LL RR,LL RR,LL RR,LL RR,LL RR,LL RR,LL	128 128 128 128 128 128 128 128 128 128	WIDR WIDR WIDR WIDR WIDR WIDR WIDR WIDR	EVLA:A:1:27	11h20m27.807s	+14d20'54.99"	11A-291.sb4911125.eb4924302.55782.00136674769 uidevla_bdf_1312848174961.bdf

### **Download options: data format**

#### Jansky VLA datasets

- Same data formats:
   SDM or MS, tar?
- Flagging and averaging options only apply to CASA MS format
  - Decreases data size
  - Apply the flags!
  - May cause coherence loss: See discussion in VLA Observational Status Summary (<u>OSS</u>)







### Some final archive notes

- The native SDM-BDF data is always good:
  - May take a while & disk space to convert to MS at home
  - Should be usable for any CASA version available
  - Can also be used for AIPS
- Archive or Pipeline (averaged/flagged/calibrated) data and/or products must use the same CASA version to proceed
  - Calibration tables specific to CASA version
  - Version used should be listed in a file in the download directory in \*\_asdm2MS.log or \*\_casalog.log



### Requesting data on a hard disk

- NRAO can ship data on hard disks upon request, e.g.
  - when the size of the data is large (over a few 100 GB)
  - when the internet connection cannot handle the request
- Disk-ordering process is done through the archive/helpdesk
- Data is shipped on a 2 TB disk (which holds I.8 TB of data)
- Cost: USD 125 per disk, potentially plus shipping cost
- Disk shipment information and policies are posted at <u>https://science.nrao.edu/facilities/vla/archive/shipment</u>



### **Getting CASA Pipeline Calibrated Data**

- Upcoming VLA CASA pipeline talks...
- Note that VLA CASA calibration pipeline products are only available through the new archive
- Request pipelined data products through the VLA Pipeline department of the NRAO help desk (https://help.nrao.edu/)

Download through the internet or ask for a hard disk (purchase)



### Loading data into AIPS

- Conversion from the native SDM into UV FITS format is no longer supported through the archive
- Download the native SDM-BDF from the archive.
- Use OBIT to load into AIPS using task 'bdf2aips'.
  <u>http://www.cv.nrao.edu/~bcotton/Obit.html</u>
- For more details on the VLA data archive, see <u>https://science.nrao.edu/facilities/vla/archive/index</u>



### Examine the visibility data (in CASA)



### CASA

- See previous lecture..
- Documentation is available at <u>http://casa.nrao.edu/</u> → 'Using CASA'
- Training material is available at <a href="http://casaguides.nrao.edu">http://casaguides.nrao.edu</a>
- For help, use the NRAO help desk at: <u>http://help.nrao.edu</u>

### CASA 6.1.2-7 will be used at this workshop

- Type "casa —ls" to find the exact string
- Then "casa [-r <version-string>]" to start





### Loading The Data: importasdm

If one chooses to download the SDM-BDF (not CASA MS)

- Task importased converts the SDM-BDF to MS
- importasdm only understands VLA online flags:
  - It converts the data into a MS while applying flags.
- > default importasdm
- > inp
- > asdm
- > vis
- > ocorr\_mode
- > applyflags
- > inp
- > go

- = 'archive\_sdm\_directory\_name'
- = 'output MS name'
- = 'co' (or load ca, ao)
  - = True



# Loading The Data: flagdata

### Additional flags for VLA data:

- Note that if applyflags = False (the default in importasdm), the flags are written to a FLAG\_CMD MS table. They can be examined (listed, plotted) and applied by using the task flagcmd
- VLA data needs additional flags for shadowing and pure zero data; these flags are applied using the *flagdata* task:
  - > default flagdata
  - > vis = 'drw.ms'
    > mode = 'shadow'
    > inp
  - > go
  - > mode = 'clip'
  - > correlation = 'ABS\_ALL'
  - > clipzeros =
  - > go



True



## **Examining Your Data**

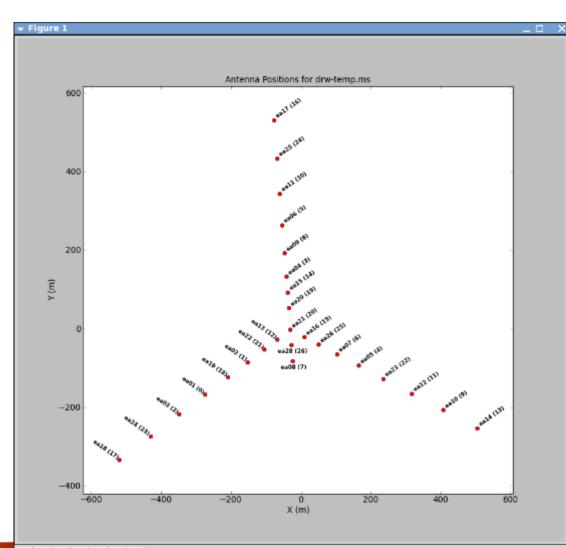
- Operator observing log (email, posted on web)
- Plotting the antenna positions: plotants (potential reference antennas)
- Observing summary: listobs (sources, scans, spectral windows, antennas, etc...)
- Plotting/displaying/editing data: *plotms* Examine your data carefully before flagging: That is, know your data content



## **Plotting the antennas: plotants**

- > default plotants
- > inp
- > vis = 'drw.ms'
- > antindex = True
- > inp
- > go

# Find a reference antenna for use in data processing



Post-Observing, Pre-Calibration 8th VLA DRW2021 LOS

x=318.726 y=-227.629

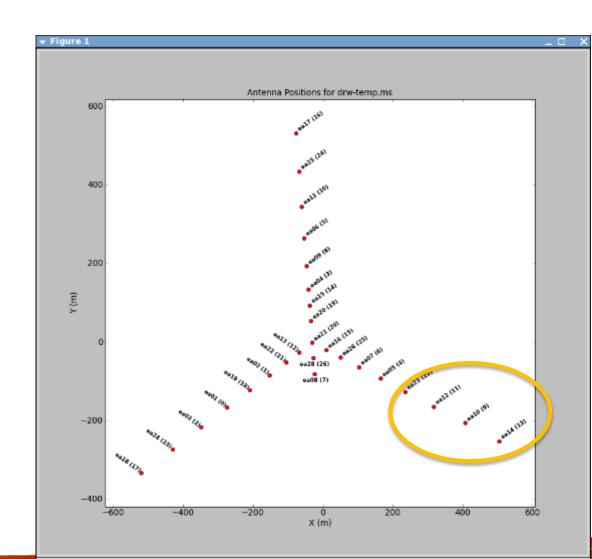
## **Choosing the reference antenna**

#### Hints:

- Must have good data on all calibrator sources (target scans irrelevant)
  - No antenna "issues"
  - Clean from RFI
- Baselines not too long
  - Near physical center of array
    - (use listobs/vishead)
  - e.g. 08-pad [WNE]
     A: inner antennas
     D: no shadowing
     Also don't move..

• Here: **ea10** 

(D array, E08, #=9)



Post-Observing, Pre-Calibration 8th VLA DRW2021 LOS

x=318.726



y=-227.629

### **Observing summary: listobs**

#### > Default listobs

#### > vis = `drw-temp.ms'

	VID			CCmp													
				lonID = 0		ArrayID =											
>	ınp	Date		Timerange			Scan		d FieldName		nRows	SpwIds		-	Inte		
	<b>_</b>	04-0	ct-201	18/05:41:35.	0 - 05	:42:31.0	1	(	0137+331=3C48	8	39317		[1, 1]	-	_	_	
>	ao			05:42:32.	0 - 05	:47:30.0	2	(	0137+331=3C48	8	20919 <mark>6</mark>	[0,1]	[1, 1]	[SY	STI M	CON	FIG
	go			05:47:35.			3		0137+331=3C48		3088		,5,6,7,				
				05:48:35	0 - 05	:49:00.0	4	(	0137+331=3C48	8	14040	[2,3,4	,5,6,7,	8,9]	[5,	, 5,	5,
				05:49:05.	0 - 05	:53:25.0	5	(	0137+331=3C48	8	146016	12.3,4	,5,6,7,	8 -1	[5	, 5,	5,
				05:53:30.	0 - 05	:57:55.0	6	1	J2355+4950		148824	[2,3,4	, 5, 0, 7,	8,9]	[5	, 5,	5,
Nc	ote:			05:58:00.	0 - 06	:03:55.0	7	2	2 J0259+0747		199368	[2,3,4	,5,6,7,	8,9]	[5,	, 5,	5,
TAC				06:04:00	0 - 06	:18:55.0	8	3	3 3C75		502632	[2,3,4	,5,6,7,	8,9]	[5	, 5,	5,
~	1 + ~+	£ ; ] ~		06:19:00.	0 - 06	:20:10.0	9	2	2 J0259+0747		39312	[2,3,4	,5,6,7,	8,9]	[5,	, 5,	5,
>	list	ТТТе	= =	06:20:15	0 - 06	:35:05.0	10	3	3 3 2 7 5		499824	[2,3,4	,5,6,7,	8,9]	[5,	, 5,	5,
				06:35:10.	0 - 06	:36:20.0	11	2	2 J0259+0747		39312	[2,3,4	,5,6,7,	8,9]	[5	, 5,	5,
				06:36:25	0 - 06	:51:20.0	12		3 3 2 7 5		502632	[2,3,4	,5,6,7,	8,9]	[5,	, 5,	5,
				06:51:25.	0 - 06	:52:30.0	13	2	2 J0259+0747		36504	[2,3,4	,5,6,7,	8,9]	[5	, 5,	5,
				06:52:35	0 - 07	:07:30.0	14	3	3 3C75		502632	[2,3,4	,5,6,7,	8,9]	[5	, 5,	5,
				07:07:35.	0 - 07	:08:45.0	15		2 J0259+0747		39312	[2,3,4	,5,6,7,	8,9]	[5]	, 5,	5,
				07:08:50	0 - 07	:23:40.0	16	3	3 3C75		499824		,5,6,7,			, 5,	5,
				07:23:45.	0 - 07	:26:25.0	17	2	2 J0259+0747		89856	[2,3,4	,5,6,7,	8,9]	[5]	, 5,	5,
				07:26:30	0 - 07	:41:25.0	18		3 3C75			[2,3,4				, 5,	
				07:41:30.	0 - 07	:42:40.0	19	2	2 J0259+0747		39312	[2,3,4	,5,6,7,	8,9]	[5]	, 5,	5,
				07:42:45	0 - 07	:57:35.0	20	3	3 3C75		499824	[2,3,4	,5,6,7,	8,9]	[5	, 5,	5,
				07:57:40.	0 - 07	:58:50.0	21	2	2 J0259+0747		39312	[2,3,4	,5,6,7,	8,9]	[5]	, 5,	5,
				07:58:55.	0 - 08	:13:50.0	22	3	3 3C75		502632	[2,3,4	,5,6,7,	8,9]	[5	, 5,	5,
				08:13:55.	0 - 08	:15:05.0	23	2	2 J0259+0747		39312	[2,3,4	,5,6,7,	8,9]	[5]	, 5,	5,
				08:15:10	0 - 08	:30:00.0	24	3	3 3 2 7 5		499824	[2,3,4	,5,6,7,	8,9]	[5,	, 5,	5,
				08:30:05	0 - 08	:32:45.0	25	2	2 J0259+0747		89856	[2,3,4	,5,6,7,	8,9]	[5]	, 5,	5,
				nRows = Tot	al num	ber of ro	ows per	scan)									
		Fields															
		ID	Code	Name		RA			Decl	Epoch	SrcId	nRows					
		0	NONE	0137+331=30	248	01:37	41.299	431 +3	33.09.35.13299	J2000	0	439452					
		1		J2355+4950					19.50.08.34001		1	148824					
		2	NONE	J0259+0747		02:59	27.076	633 +0	07.47.39.64322	J2000	2	651456					
		3	NONE	3C75		02:57	42.630	000 +0	06.01.04.80000	J2000	3	4512456					
		Spectra	al Win	(10	-	-	windo	ws and	i 1 unique pola	arizatio	n setups)						
		SpwIl	D Nan	ne	#Chans	Frame	Ch0 (	MHz)	ChanWid(kHz)	TotBW (k)	Hz) CtrFre	q(MHz) B	BC Num	Cor	rs		
		0		LA_C#A0C0#0	64		4832		2000.000	12800		.0000	12		RL		LL
		1		A_C#B0D0#1	64		4960		2000.000	12800		.0000	15	RR		LR	$\mathbf{L}\mathbf{L}$
		2		LA_S#A0C0#2	64		2488		2000.000	12800		.0000	12	RR		LR	LL
		3		A_S#A0C0#3	64			.000	2000.000	12800		.0000	12				$\mathbf{L}\mathbf{L}$
								000	0000 000	10000		0000		-			

# **Observing summary: listobs**

Summary, for example:

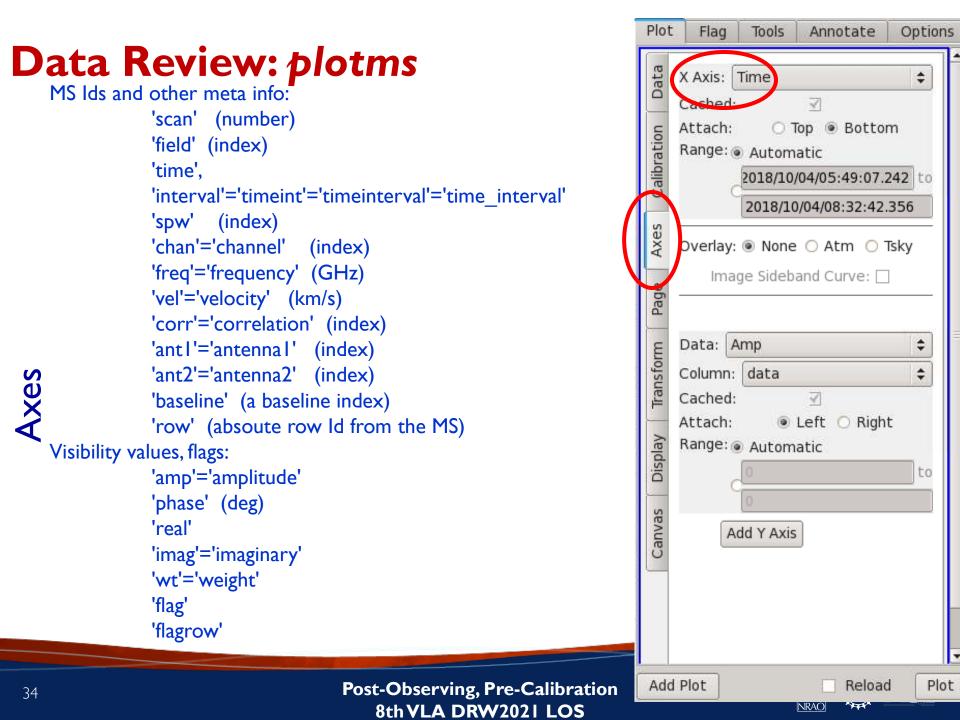
- Four sources:
  - 0137+331=3C48
    - 5 scans of which 4 are setup, i.e. 1 scan on Flux calibrator
  - J2355+4950, 1 scan, some calibrator
  - J0259+0747, 10 scans, Gain/phase calibrator
  - 3C75, 9 scans, source with the most time on source, has the most visibilities so this would be the target source
- 8 spectral windows (spw), full polarization products (RR,RL,LR,LL)
  - Ranging 2488 to (3384+128=) 3512 MHz, centered at 3000 MHz
  - 64 channels per spw/pol, each 2 MHz wide
- Using 27 antennas (ea27 is not used)
  - Highest numbered antenna pad is E09/N09/W09: D-array



### Data Review: plotms

33

Carvas Display Transform Page Axes Data	PlotMS Graphics Panel	
		pols Panel



### **Data Review:** *plotms*

#### Unnagged Po Spw Display None Antenna1 Antenna2 Data Baseline Style: 2 Colorize by: Channel Corr Axes Fill: 0000 Time Scan Outline: 💿 None 🔘 Default Field Page Flagged Points Symbol Spw None $\sim$ ansform Antennal Custom Style: 2 ‡px, circle Antenna<sup>2</sup> Fill: ff0000 .... **Baseline** isplay Outline: None O Default Channel Correlation Canva Time Observation Intent

**Post-Observing, Pre-Calibration** 8thVLA DRW2021 LOS

Add Plot

000 Plot

Flag

Tools

Colorize V Scan

Y Axis Data: Amp: corrected

Field

Annotate

Options

÷

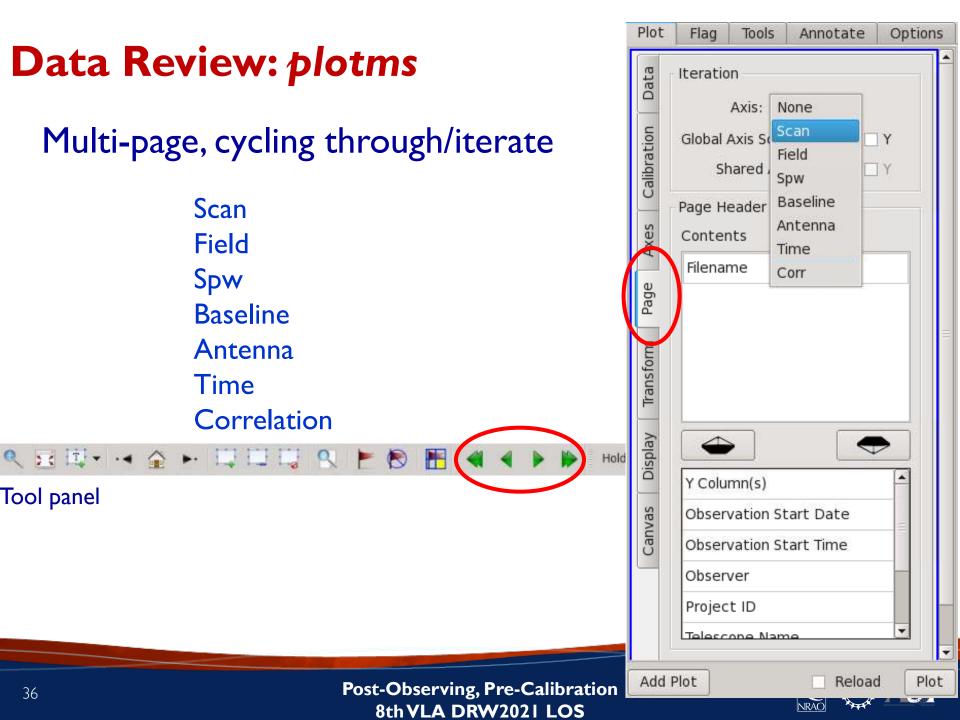
alin ‡

Default

fill

÷

**Reload** Plot



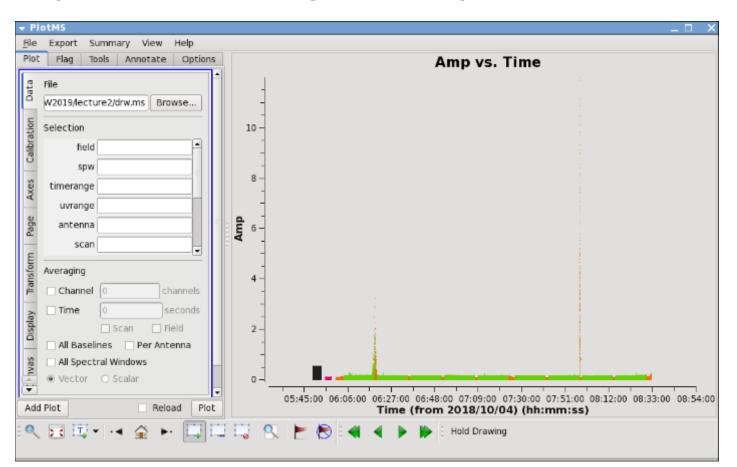
### What are we looking for?

- A feel of the overall structure of the data (see also the OPT schedule):
  - Calibrators and target visibilities, frequency setup
  - Observing conditions, instrumental response
- Where to expect bad data
  - Specific ill-performing antennas/baseline(boards)
  - In time
    - Start of scans
    - Bad weather/pointing/RFI (observing conditions)
  - In frequency
    - Bandpass, subband edges
    - RFI not your line!



#### **Data Review:** plotms

#### Example: xaxis='time', yaxis='amp,' coloraxis='field'





### **Radio Frequency Interference (RFI)**

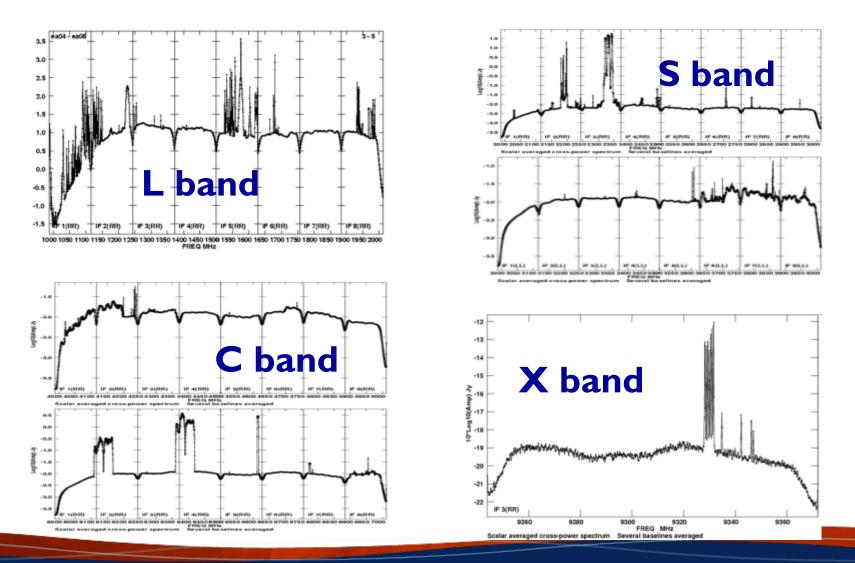
- I. VLA observations, particularly at the lower frequency bands, will be severely affected by RFI.
- 2. VLA RFI information is available at:

<u>https://science.nrao.edu/facilities/vla/</u>  $\rightarrow$  Observing  $\rightarrow$  VLA Observing Guide  $\rightarrow$  #5: Radio Frequency Interference

- RFI listings per frequency band.
- Spectra of various RFI sweeps between I-50 GHz.



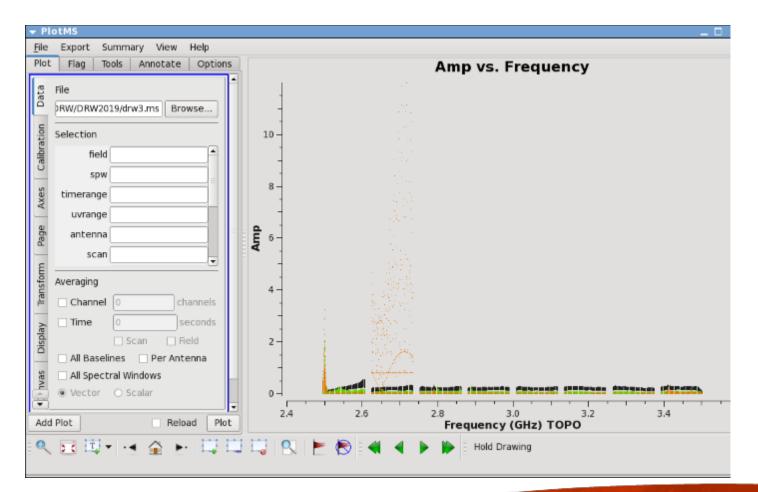
#### **RFI is present at lower frequency bands**





#### **Data Review:** plotms

#### Example: xaxis='frequency', yaxis='amp',coloraxis='field'





#### **Preparing for calibration: editing**



# Editing/Flagging (or unflagging) Data

- I. flagdata: All purpose flagging task based on selection.
  - Includes RFI flagging capabilities (RFLAG, TFCROP).
- 2. *flagcmd*: All purpose flagging task based on commands (alternative to *flagdata* for certain types of flagging).
- 3. plotms: Interactive flagging

Review the VLA operator's log carefully. Certain issues (e.g., antennas without receivers), do not end up in the online flags, and may need to be flagged manually.



#### Editing/Flagging (or unflagging) Data A few important notes

- Data in CASA are either flagged or not flagged.
  - Every MS has a flag column.
  - Every bit of data has its own flag (set either to True or False).
  - Applying flags means setting the flag column entries of the selected bits of data to True there is no record of what it was before setting to True.
  - Unflagging sets it to Flase, regardless it's immediate previous setting
- Most flagging tasks have the option to create a flag backup.
  - In particular *plotms* does not have this option!
- A flag backup is an MS table and contains the state of the flags before running the flagging task.
- With *flagmanager* flag back-ups can be restored (and made)



### Editing Data: flagdata - Modes

- list = apply a list of flagging commands
- manual = flagging based on specific selection parameters [default]
- *clip* = clip data according to values
- quack = remove/keep specific time range at scan beginning/end
- shadow = remove antenna-shadowed data
- elevation = remove data below/above given elevations
- *tfcrop* = auto identification of outliers on the time-freq plane
- rflag = auto detection of outliers based on sliding-window RMS filters
- extend = extend and/or grow flags
- Also summary (per antenna, correlation, field, scan, total), and unflag.
- Can also flag calibration tables.



# Editing Data: flagcmd

- It allows listing, plotting, saving, applying, or un-applying flags.
- Flagging modes (inpmode) are:
  - table: uses the FLAG\_CMD MS table (created by importasdm)
  - *list*: uses an ASCII file that contains a set of flagging commands.
  - *xml*: uses the online flags from Flag.xml in the MS.
- It allows the user to save the flag records in the FLAG\_CMD MS table or a file.



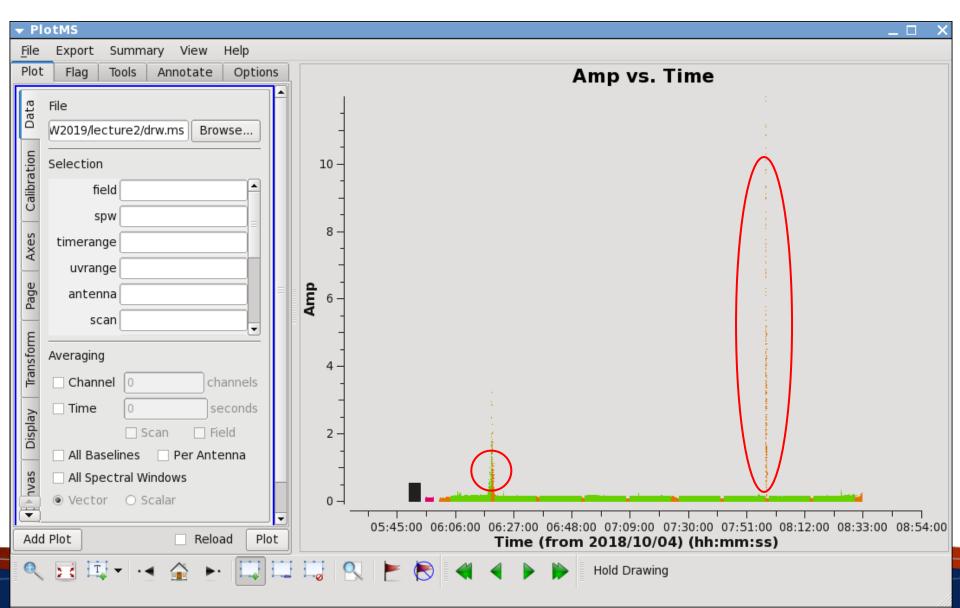
# Editing Data: flagdata vs. flagcmd

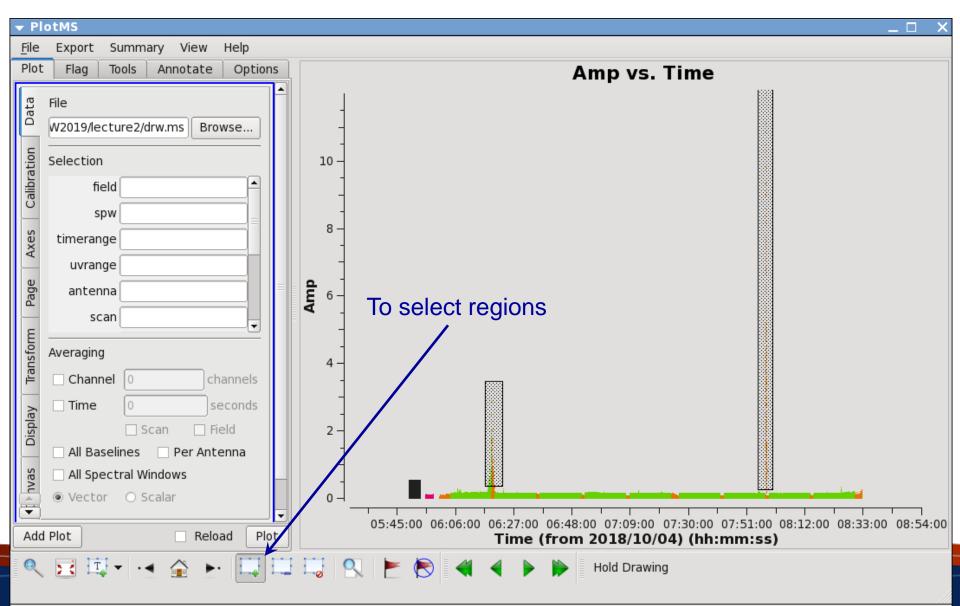
- Complementary flagging tasks.
- Have several common features.
- Some of the important differences:

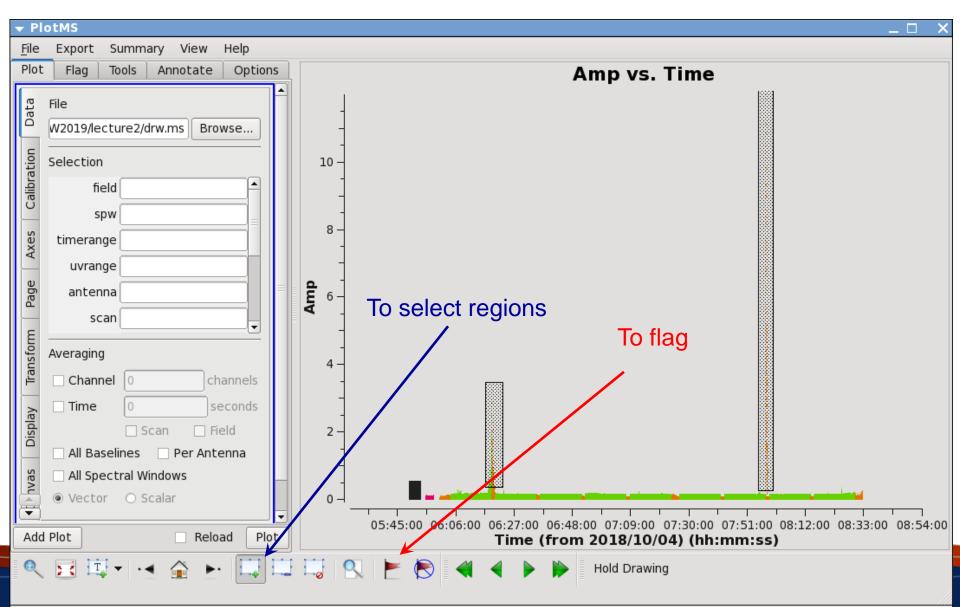
Flagdata	Flagcmd					
RFI flagging (tfcrop, rflag)*	Access to the Flag.xml					
Runtime displays* (before and after flagging)	Apply the online (and other) flags in FLAG_CMD MS table					
	Plot Flags					

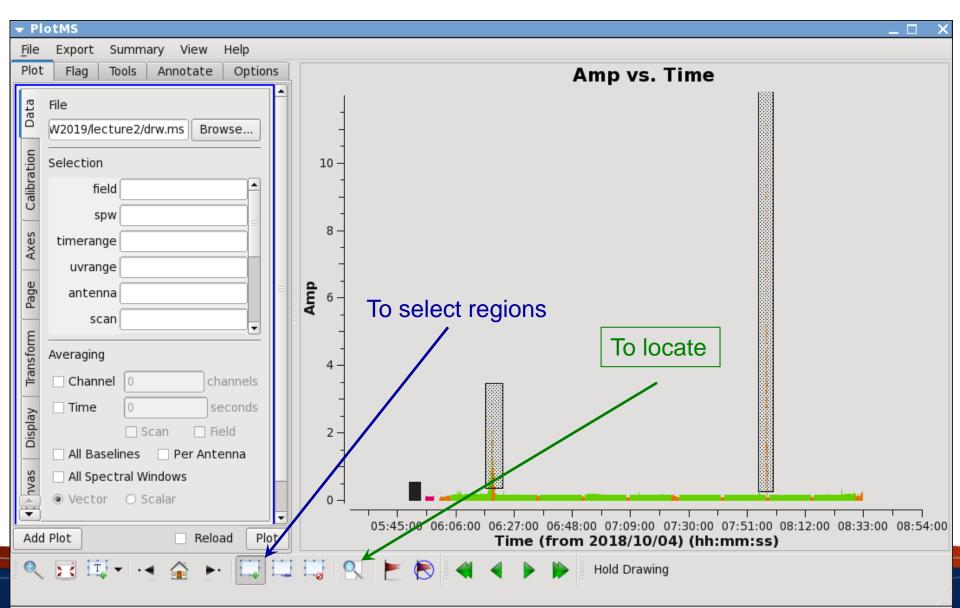
\* More details on Tuesday (RFI talk)











#### The output of "locate" in the logger – look for common lines

🔻 Log	j Me	ssage	s (euro:	:/lust	tre/	aoc/scio	ps/lsjou	wer/ca	isaDRW/	DRW20	)19/lec	:tu	re2/casa-	201909	24-14	4121.log	)	_ [	⊐ ×
<u>F</u> ile	<u>E</u> dit	<u>V</u> iev	v																
				K L	J	Search M	essage:				#6	T	Filter: Ti	me	\$			] 7	C
Messa	ige																		-
Scan=	21 F.	iel <u>d</u> -	10259+07	747 [	2]	Time=201	8/10/04	/07:58	:22.5000	BL=ea	03@ <b>W</b> 07	æ	ea1∠@E07	[2&11]	S⊦ v=1	Chan=40	Freq=2.	702	Corr
				-	-						-		a12@E07		-		-		
				=	-						-		ea12@E07		-		-		
	_			-							_		ea12@E07		-		-		
													ea12@E07						
	_												ea12@E07		-		-		
				-	-						-		ea12@E07 ea12@E07		-		-		
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4								111											
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#### Editing Data: plotms A few important notes

- Use plotms carefully for editing data.
- Keep in mind that editing data with *plotms* often requires extending the flags (through the Flag tab).
- plotms does not produce a flag backup (flagmanager has to be used).
- Use *plotms* to identify bad data (through the locate option). Then flag the bad data using *flagcmd* or *flagdata*.



#### **Ready to calibrate the data?**

- The data structure is understood, reference antenna picked
   Calibrators (flux density, bandpass, gain) are identified
- ✓ Bad antennas and bad basebands are flagged
- ✓ RFI is removed (as much as possible), Hanning smooth?
- ✓ Bad individual visibilities/baselines/times are flagged
- Maybe inspect (some parts of) the data again to make sure Likely more flagging may need to be done during/after calibration steps
- Ready to start with data calibration

Next lecture...





#### www.nrao.edu science.nrao.edu public.nrao.edu

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