

Visibility Data Inspection Loránt Sjouwerman, NRAO



Goal: prepare for (easier) calibration

Examine the visibility data, start from archive download

- Science Data Model (SDM) and Binary Data Format (BDF)
 - SDM = metadata; sources, antennas, correlator, weather, etc
 - BDF = single scan data ("missing BDF": scan lost)
- Convert to data reduction package format, here CASA
- Retrace observing schedule, observing conditions
 - Data structure, scan intents, antennas, operator logs (cf. OPT)
- Remove (flag) problematic data, items for watchlist
 - Known issues (off-source, unused antenna)
 - Unforeseen issues (RFI, unnoticed by operator)
 - Questionable data (may need flagging later, or not..)



CASA

- See previous lecture...
- Documentation is available at
 - <u>http://casa.nrao.edu/</u>
- Training material is available at
 - <u>http://casaguides.nrao.edu/</u>
- For help use the NRAO helpdesk at
 - <u>http://help.nrao.edu/</u>
- CASA 6.2.1 will be used at this workshop
 - Type "casa –ls" to find the exact version string
 - Then "casa [-r <version-string>]" to start





Loading the data: importasdm

Goal: convert SDM/BDF into CASA visibility data format

- Skip if CASA MS ("measurement set") format is downloaded
- Task importasdm converts SDM/BDF to MS for CASA processing
 - importvla for older data; more options like importfits ..
- *importasdm* understands and applies VLA online flags:

Visibility Data Inspection						
	go		execute task			
	inp		review parameter selections			
	applyflags		True			
	ocorr_mode	=	'co'			
Write:	vis	=	'CASA_MS_directory_string'			
Read:	asdm	=	'SDM/BDF_directory_string'			
	inp		review parameter options			
	default importasdm		set task and default parameters			

Loading the data: flagdata

If applyflags=False (default) a FLAG_CMD table is created

- Examine, list/plot (before applying by default) using *flagcmd*
 - Also used for other flagging operations, see later in this talk
- Additional useful a-priori flagging uses flagdata (also see later)

	default flagdata		set task and default parameters
Read/write:	vis	=	'CASA_MS_directory_string'
Shadowing:	mode	=	'shadow'
	inp	=	review parameter selections
	go	=	execute task
Pure zero data:	mode	=	ʻclip'
	correlation		'ABS_ALL'
	clipzeros		True
	go		execute task
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Examine your observation/data

Goal: know your data content and find/fix possible issues (that will give problems) before starting the calibration

- Operator observing log (email to observers)
 - <u>http://www.vla.nrao.edu/cgi-bin/oplogs.cgi</u>
- Plotting the antenna positions: *plotants*
 - Array configuration, possible reference antennas
- Observing summary: listobs
 - Calibrator sources, scan intents, spectral configurations, etc..
- Plotting/displaying/editing data: *plotms*
 - Examine, assess, act: (visibility) to be or not to be (retained)
- (pipeline operations weblog might be useful too if available)





Operator observing log

Goal: first impression about the observation

• Array status, weather, antenna problems etc, downtime measure

VLA OBS	FPV				201	8-10-04	_0541	L_TD	RW0001							
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Adobe PDF version of Visibility data is updat	this log is located at: ed each day at IAT/U	nttp://www.vla.nrao.ed T midnight and is availa	u/operators/logs/ ble from the online a	rchive at: https:/	/archive.nrao.eo	lu				VLA OBS	ERVING	LOG	2018-10-04	_0541	_TD	RW0001
			Wind Speed &	Bar. Pressure								internitient?	d in the court	/	+	
Time (17	Dew Point (C)	Temp. (C)	Direction (avg)	(mbars)	Phase (degs		Rer	narks				- If possible, a spectrum of the RFI should be include	d in the e-mail.	/	+	
040 .+4:10	6.2	17.2	SW at 6.7 m/s	789.7	4.3	Sky cover 2.	Cumuliforr	n clouds.				Inanks very much for your support, this information				
ct 6:46:31	7.0	14.3	SW at 4.7 m/s	789.4	4.7	Sky cover 10%	Stratiform	n clouds.				updated on the construction of the second se	an ida (mada a / fi/			
										040ct 5:41:20		Lastanaas in the Diarray may be shadowed at low eleving	signer If shadowing	,!		
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										040ct 5:41:30		NOTE: The VLA is still recovering from a long nower	utage and these data may	/		
										04001 3.41.50		have unusual artifacts missing antennas or IEs ect	n them NRAO staff will	/	+ +	
					1							evamine the data closely after obsenting to determine	if they meet the criteria for			
												a successful observation	If they meet the chiefta for		<u> </u>	
Number o	f antennas used:	27								040rt 5:41:30	040ct 8:32:51	Antenna(s) 5 (Data: Corrunted):	FRONT END	-959	0.97	165.4
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Start Time	End Time		Comm	nents/Outages		-	Form #	#Ants	(in minutes)	040ct 5:41:30	040ct 8:32:51	Antenna(s) 12 (Data: Corrupted):	C ^D NICS	PM	0.03	5.8
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04Oct 5:41:30		The band(s) used is	s(are): C S.					+		040ct 5:57:00		Your new operator(s) is(are): Sam Gilmore				
04Oct 5:44:06		On source 0137+33	1=3C48 w	ailable antenna	S.			+								
040ct 5:41:30		To access your data	- NRAO an	chive visit:												
		https://science	.edu/facilities/vla/a	archive.			<u> </u>	+								
		All VLA Le data	are processed thro	ough the VLA ca	alibration pipeli	ne. Details		+								
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Choosing the reference antenna(s)

Goal: good 'refant' reduces loss due to flagging/calibration

- Use plotants:
- Baselines/uv-dist not too long
 - Near physical array center
 - Use listobs/vishead
 - e.g. 08-pad: [WNE]08
 - A: inner antennas
 - **D**: no shadowing !
- Good on all calibrator scans
 - Target scans irrelevant
 - No antenna "issues"
 - Clean from RFI

default plotants		set task and default parameters
inp		review parameter options
vis	=	'CASA_MS_directory_ string'
antindex	=	True
inp		review parameter selections
go		execute task



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Choosing the reference antenna(s)

Goal: good 'refant' reduces loss due to flagging/calibration



• Clean from RFI



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Observing structure

Goal: identify key strategy, scan outline, calibration hints

- Use listobs (not vishead):
- Source (i.e. field) names
- Scan intents: calibrators
 - Flux density
 - Bandpass
 - Gain (phase and amp!)
- Correlator setup
 - Spectral windows
 - (IFs/subbands, channels)
 - Polarization
 - Doppler setting

default listobs		set task and default parameters
vis	=	'CASA_MS_directory_string'
inp		review parameter selections
go		execute task
NOTE:		
listfile	=	'output_filename_string'



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 - Spectral windows

 (IFs/subbands, channels)
 - Polarization
 - Doppler setting

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		06	:20 15.0	0 - 06:3	5:05.0	10	3	3C75	
		06	:35.15.0	0 - 06:3	6:20.0	11	2	J0259+0747	
		06	:3:30.0	0 - 06:5	1:20.0	12	3	3C75	
		06	:5.:25.0	0 - 06:5	2:30.0	13	2	J0259+0747	
		06	: 2:40.0	0 - 07:0	7:30.0	14	3	3C75	
		07	: 7:40.0	0 - 07:0	8:45.0	15	2	J0259+0747	
		07	:(3:50.0	0 - 07:2	3:40.0	16	3	3C75	
		07	:2:50.0	0 - 07:2	6:25.0	17	2	J0259+0747	
		07	:2:35.0	0 - 07:4	1:25.0	18	3	3C75	
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		07	:42:5.0	0 - 07:5	7:35.0	20	3	3C75	
		07	:57:4 (0 - 07:5	8:50.0	21	2	J0259+0747	
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Examine MS visibility data: plotms

Intermezzo: introduction to *plotms* GUI and panels:



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Examine MS visibility data: plotms

Intermezzo: introduction to plotms GUI and panels:



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Examine MS visibility data: plotms

Intermezzo: introduction to plotms GUI and panels:



Control panel parameters

Axes options

- Typical **X Axis** selections (not all):
 - 'time'/'scan'/'field'
 - 'frequency'(freq)/'spw'
 - 'channel'(chan)/'velocity'(vel)
 - 'antennal'(antl)/'antenna2'(ant2)/'baseline'
 - 'UVdist'/'UVwave'
- Typical **Data** (i.e.Y-axis) selections:
 - 'amplitude'(amp)
 - 'phase'
 - 'real'/'imaginary'(imag)
 - 'weight'(wt)/'flag'

Plot	Flag Tools Annotate Options
Data	X Axis: Time
Calibration	Cached: Attach: Top Bottom Range: Automatic 2018/10/04/05:49:07.242 to
Axes	2018/10/04/08:32:42.356
Page	Image Sideband Curve:
Transform	Data: Amp 🔶 Coronne Gata 🗘
Display	Cached: Attach: Range: Automatic
Canvas	Add Y Axis



Control panel parameters

Display options

- Typical **Colorize** selections: For time-sequence-like x-axis
 - Field, scan
 - Antenna I/2, baseline
 - For frequency-order-like x-axis
 - spw
 - Correlation (i.e. polarization)
 - Antenna I/2, baseline

For array-configuration-like x-axis

- Antenna I/2, baseline, spw, field, correlation

Plot	Flag Tools Annotate Options	
Data	✓ Colorize: Feld	•
Calibration	Uninagged Points Symbol None © Default Custom	
Axes	Style: 2 px, autoscaling Fill: 0000ff fill Outline: None O Default	
m Page	Flagged Points Symbol	
Transfo	Style: 2 px, circle Fill: ff0000 fill Outline: None O Default	=
Display	None O Derault	
Canvas	Connect along time axis	



Control panel parameters

Page options

In multi-page mode, when plots are distributed over more than one plot, one can cycle through the plots using the page-selection tool icons:





Hold Drawing

- Typical **multi-page** selections:
 - Antenna
 - Correlation
 - Field
 - Spw

Plot	Flag	Tools	Annotate	Options	
Data	Iteration	1 ———			
s Calibration	Global / Sł Page He	Axis: Axis Sca nared A eader —	None Scan Field Spw Baseline Antenna	□ Y □ Y	
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Examine, assess, act

Goal: remove potentially problematic data, identify other potential problems that may influence calibration/imaging

- Overall sense of calibrators/intents and target visibilities, frequency setup, observing conditions, instrumental response
- Where to expect potentially bad data, note the reference antenna!
 - III-performing antennas/baselines/samplers/polarizations (instrument)
 - In time domain:
 - Start of scans (not yet on-source, settling time)
 - Bad weather/pointing/RFI-bursts (observing conditions)
 - In frequency domain:
 - Bandpass, subband edges and baseband roll-off (channel sensitivity)
 - RFI-frequencies not your line!





Time/scan/field dependent issues

Example: xaxis='time', yaxis='amp', coloraxis='field'









Radio frequency interference (RFI)

Intermezzo: RFI awareness

- Observations, in particular in the observing bands 4, P, L, S, C, but also X and Ku, will be affected by time/frequency dependent RFI
- VLA RFI information is available at http://go.nrao.edu/vla-rfi
 - Contains RFI listings per frequency band and array configuration
 - Shows (snapshot) spectra of various RFI sweeps between 1-50 GHz



Frequency/spw dependent issues

Example: xaxis='frequency', yaxis='amp', coloraxis='field'



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Editing (flagging or unflagging) data

Goal: introduce different flagging tasks and methods

- flagdata: All purpose flagging based on selection and algorithms
 In particular automated RFI flagging operations RFLAG and TFCROP
- *flagcmd*: All purpose flagging based on commands
 - See earlier shadowing and pure zero visibility examples
- *plotms*: interactive flagging using graphical user interface
- Note that the VLA operator log has useful information but may not capture all (or too many) possible items to flag, e.g.
 - Antennas not participating in array may need manual flagging
 - "low amplitude"/"weak": pointing/baseband error or recoverable?





Editing (flagging or unflagging) data

Goal: understand flagging and flag-backups in CASA

- Each visibility is either **used** (not flagged) or **flagged** (ignored)
 - A flag-column contains a False (default) or True bit for each visibility
 - "Applying flagging" converts this bit to True, regardless the state it was in before; there is no record keeping of the original value
 - "Unflagging" resets this bit to False, with no knowledge of the original value if it was previously flagged this will be undone too!
- Most flagging tasks but not plotms! allow to save a copy of the previous state, before applying new flags, in a flag-backup file
- *flagmanager* restores/creates a/the previous/current flagging state
 - When planning to flag with *plotms*, maybe make a flag-backup first?
 - Sequential flag-backup files restore to sequential stages in process



Using flagdata

Different modes and options to use in *flagdata*

- manual flag based on specific selection of parameters (default)
- *list* apply a list of flagging commands in a file or list of files/string
- *clip/quack* clip on amplitude or (scan)time range values
- shadow/elevation flag on antenna geometry
- *tfcrop* algorithmic removal of outliers in time-frequency plane
- rflag algorithmic removal of outliers with respect to local RMS
- extend/antint expand flags along different axis (time, freq, ants, ..)
- summary list what is flagged
- unflag reset flag-column bits to False (but be aware and careful!)



Using flagcmd

Different modes and options to use in *flagcmd*

- Table flag items in the FLAG_CMD table (default)
- List apply a list of flagging commands in a file
- Xml apply on-line flags as listed in the Flag.xml file of the MS
- *flagcmd* can make a plot of flags but lacks runtime displays
- Both flagcmd and flagdata allow saving to FLAG_CMD file
- Both *flagcmd* and *flagdata* have unflag/unapply operations
- Both *flagcmd* and *flagdata* also operate on **calibration tables**



Goal: look for outliers, here amplitude versus time

• Two obvious times where probably "something is going on"



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Goal: select outliers with select tool icon

• Use mouse to draw; selections show up as shaded regions



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Goal: **remove outliers with** flag **tool icon** (but wait..!)

• Flag tool may miss low-level bad data – it only flags in the regions



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Goal: list outliers with locate tool icon

• Locate will list visibilities in region in message log (can be many!)



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Goal: identify common identifying parameters with locate

• all channels in spw I, on baseline ea03-ea12 at 7h58m in scan 21

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Message						
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Important final notes

FYI:

- Be careful when using *plotms* for editing data
 - Interactive unflag possibility may not properly unflag
 - No flag backup file (must use flagmanager)
- Editing data with *plotms* perhaps may require extending the flags using the Flag tab:
- Probably best practice:
 - Use the locate feature and message log window in *plotms* to identify bad data
 - 2) Flag the bad data using flagcmd or flagdata
 - 3) Check that all bad data is gone (extended the flags?)
 - 4) Repeat if necessary (new bad data may appear)





Finally!(?)

Summary of **visibility data inspection** achievements

- \checkmark Data structure is understood, reference antenna is picked
- ✓ Calibrators (flux density, bandpass, gain) are identified
- \checkmark Bad antennas and other instrumental flaws are flagged
- ✓ RFI is removed (as much as possible); Hanning smooth?
- Residual bad individual visibilities are flagged (at least on calibrators – bad target data can be flagged before imaging)
- Maybe inspect (some parts) of the data again to make sure More flagging may need to be done during/after calibration steps
- **Ready to start with data calibration** (next lecture)





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