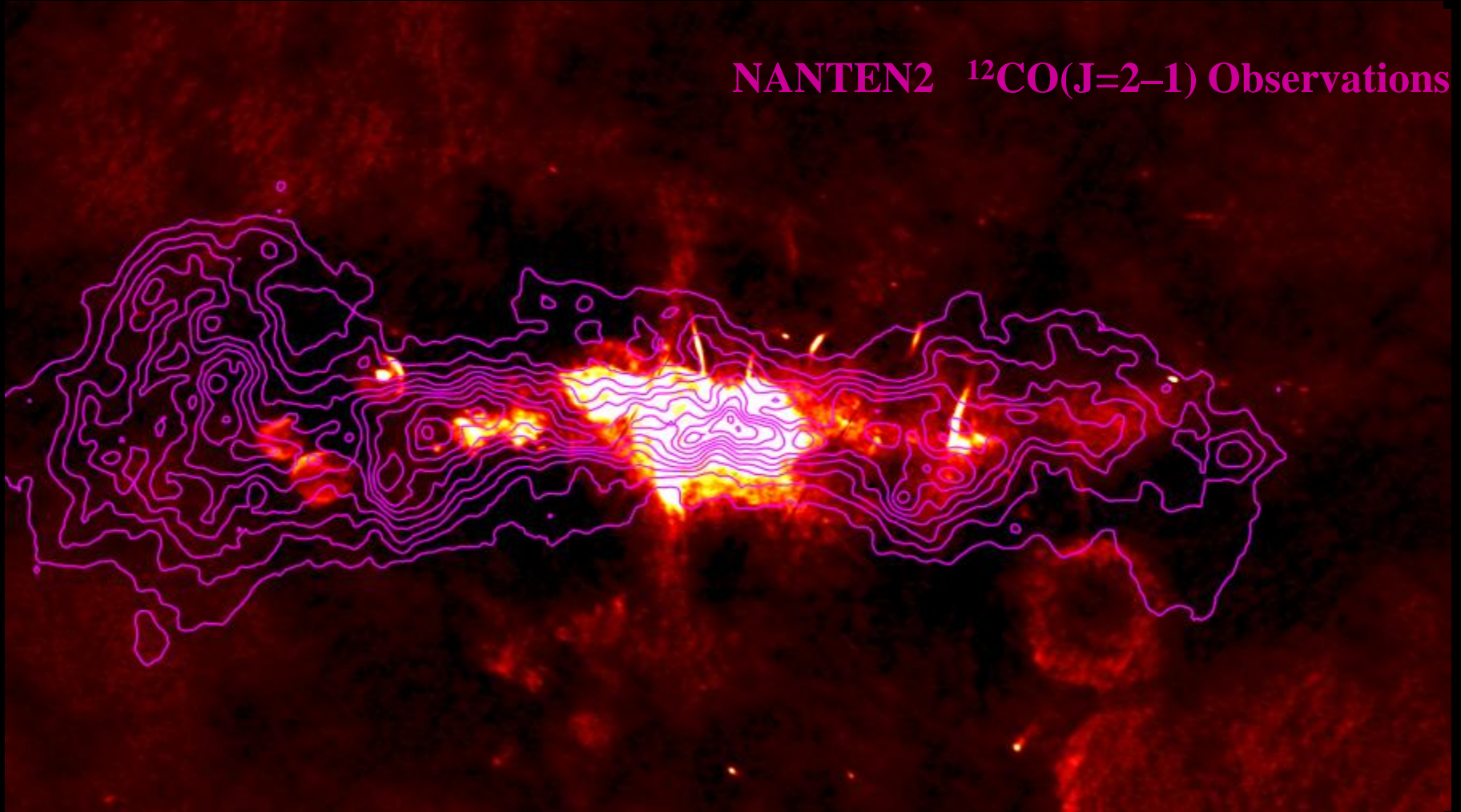


Large scale and high sensitivity multi line CO surveys toward the Galactic Center

Rei Enokiya (Nagoya university)

K. Torii, M. Schultheis, Y. Asahina, R. Matsumoto, K. Dobashi,
A. Ohama, N. Moribe, A. Nishimura, H. Yamamoto, K. Tachihara,
T. Okuda, A. Kawamura, N. Mizuno, T. Onishi, M. R. Morris
and Y. Fukui

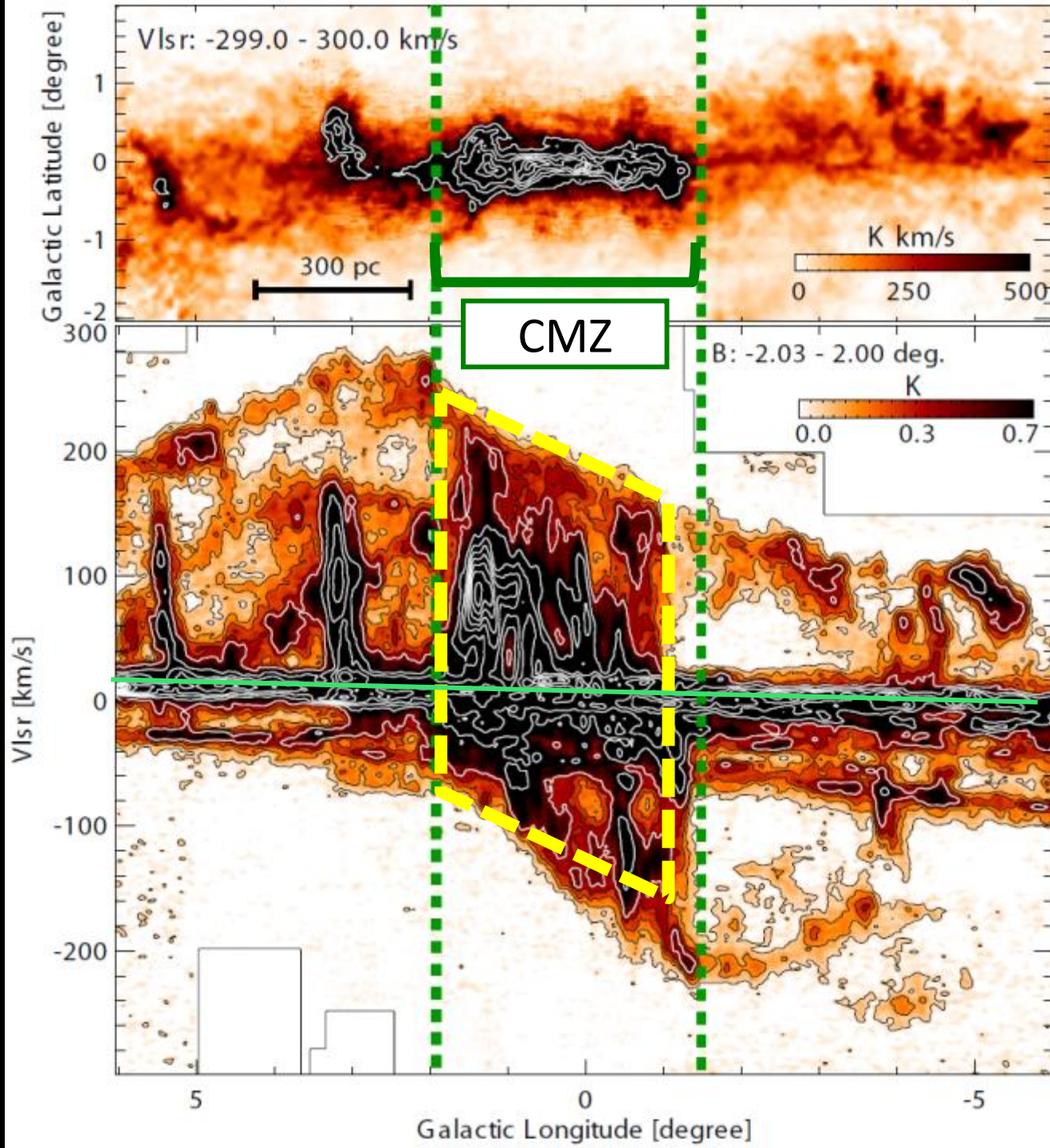
NANTEN2 $^{12}\text{CO}(J=2-1)$ Observations



CMZ (Central Molecular Zone)
Morris & Serabyn 1996

VLA 90 cm Radio continuum
(LaRosa et al. 1999)

NANTEN
 $^{12}\text{CO}(J=1-0)$
Observations



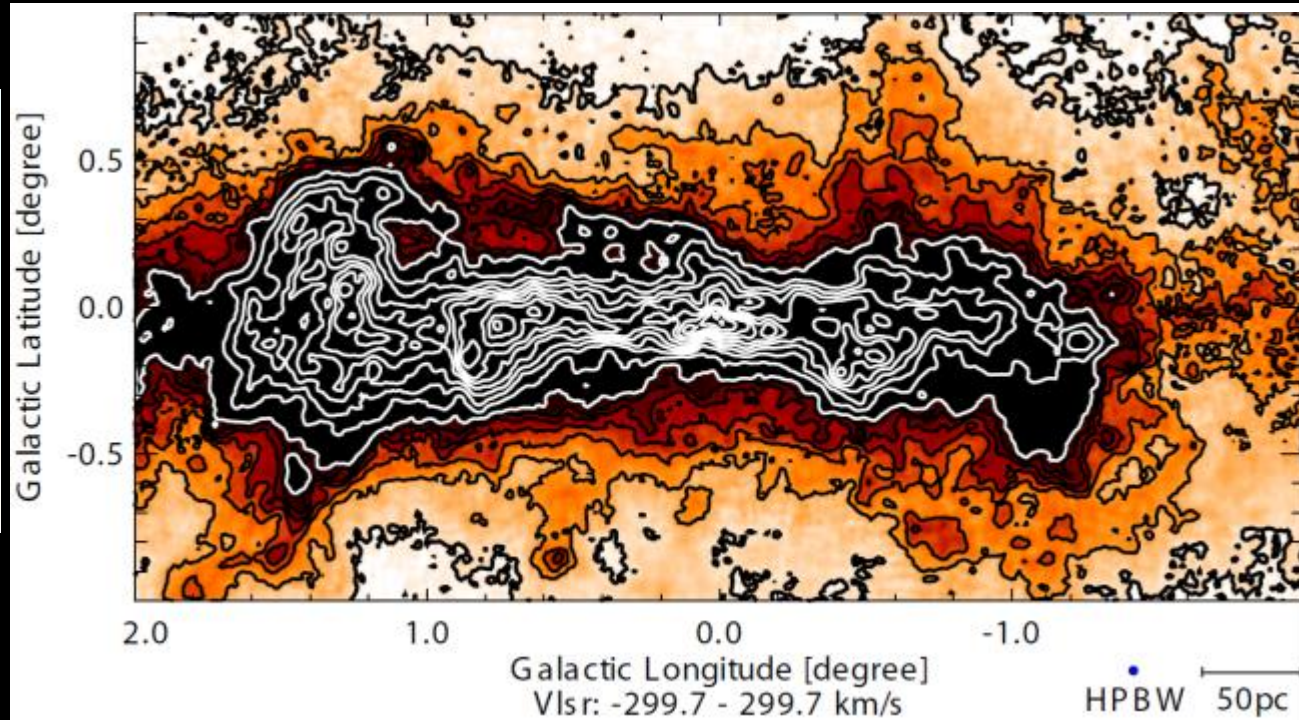
Observation (NANTEN2)

mm / sub-mm telescope

NANTEN2

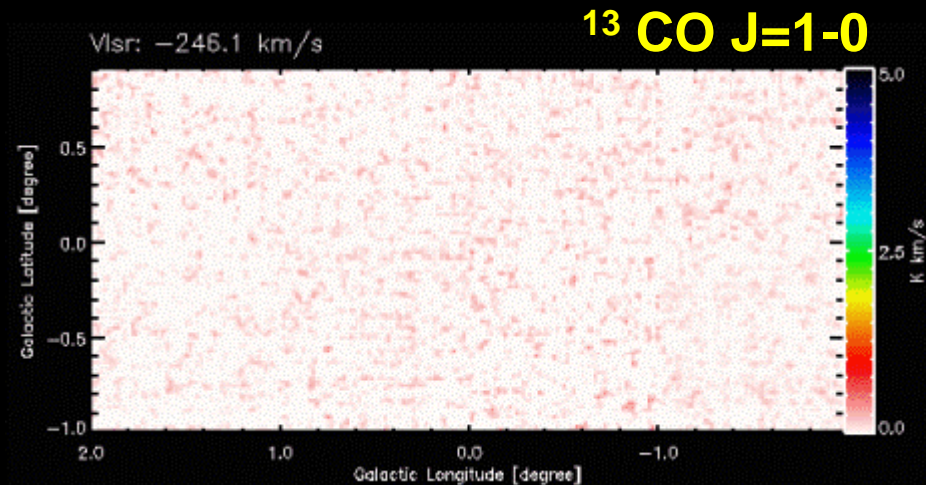
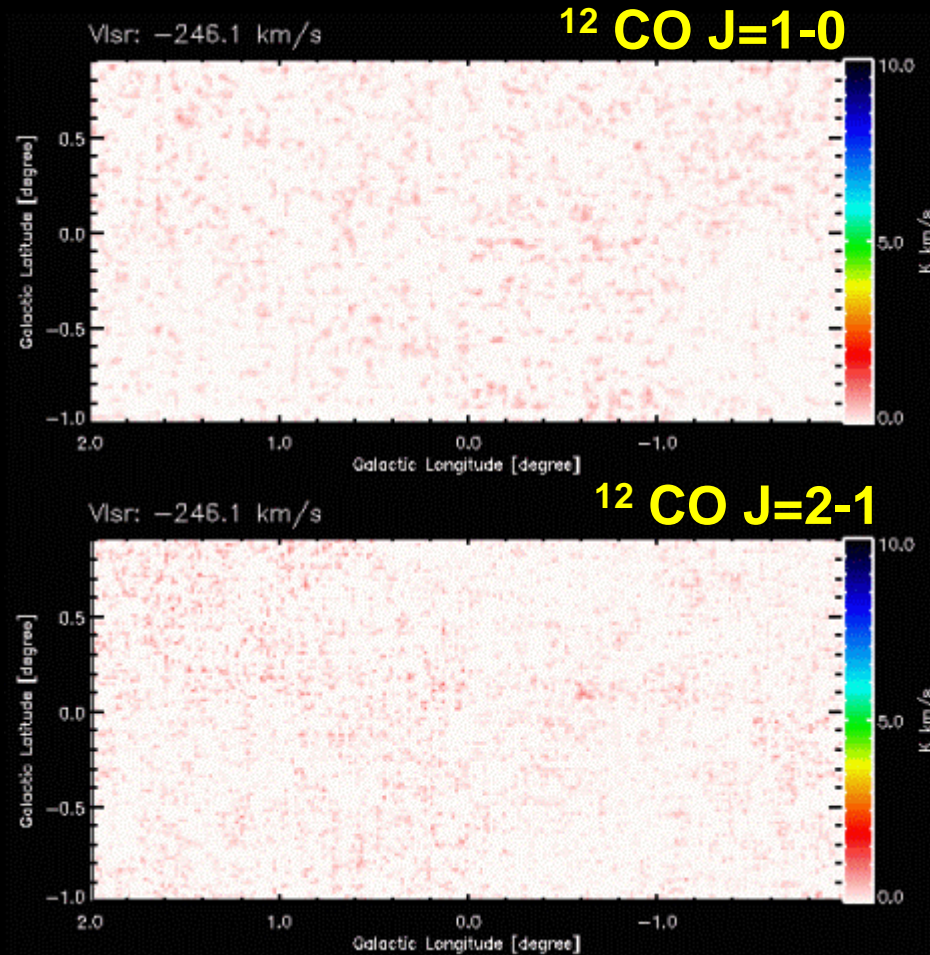


- Atacama, Chile
(Alt. 4865m)



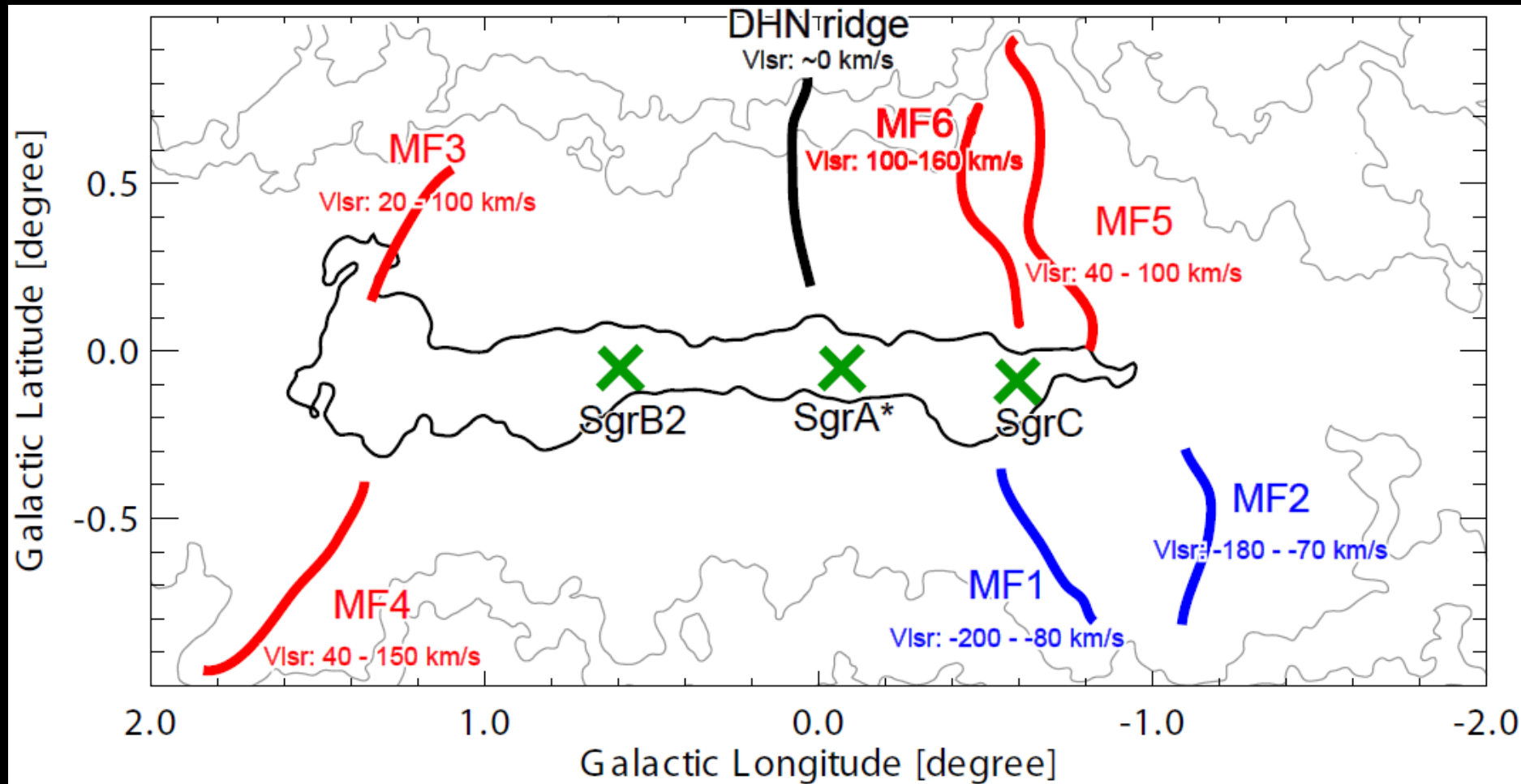
| Line | $^{12}\text{CO}, ^{13}\text{CO} (1-0)$ | $^{12}\text{CO} (2-1)$ |
|----------------------------|--|------------------------|
| Beam size (HPBW) | 180" | 90" |
| Spatial resolution (@8kpc) | 7.0 pc | 3.5 pc |
| Vel. resolution (km/s) | 0.16 | 0.08 |
| Sensitivity (K/ch) | 0.58, 0.43 | 0.25 |
| Covered are (deg) | $L = \pm 10, B = \pm 1$ | $L = \pm 2, B = \pm 1$ |

CO distribution toward the CMZ



Several molecular filaments are vertically distributed to the plane

Molecular filaments & halo



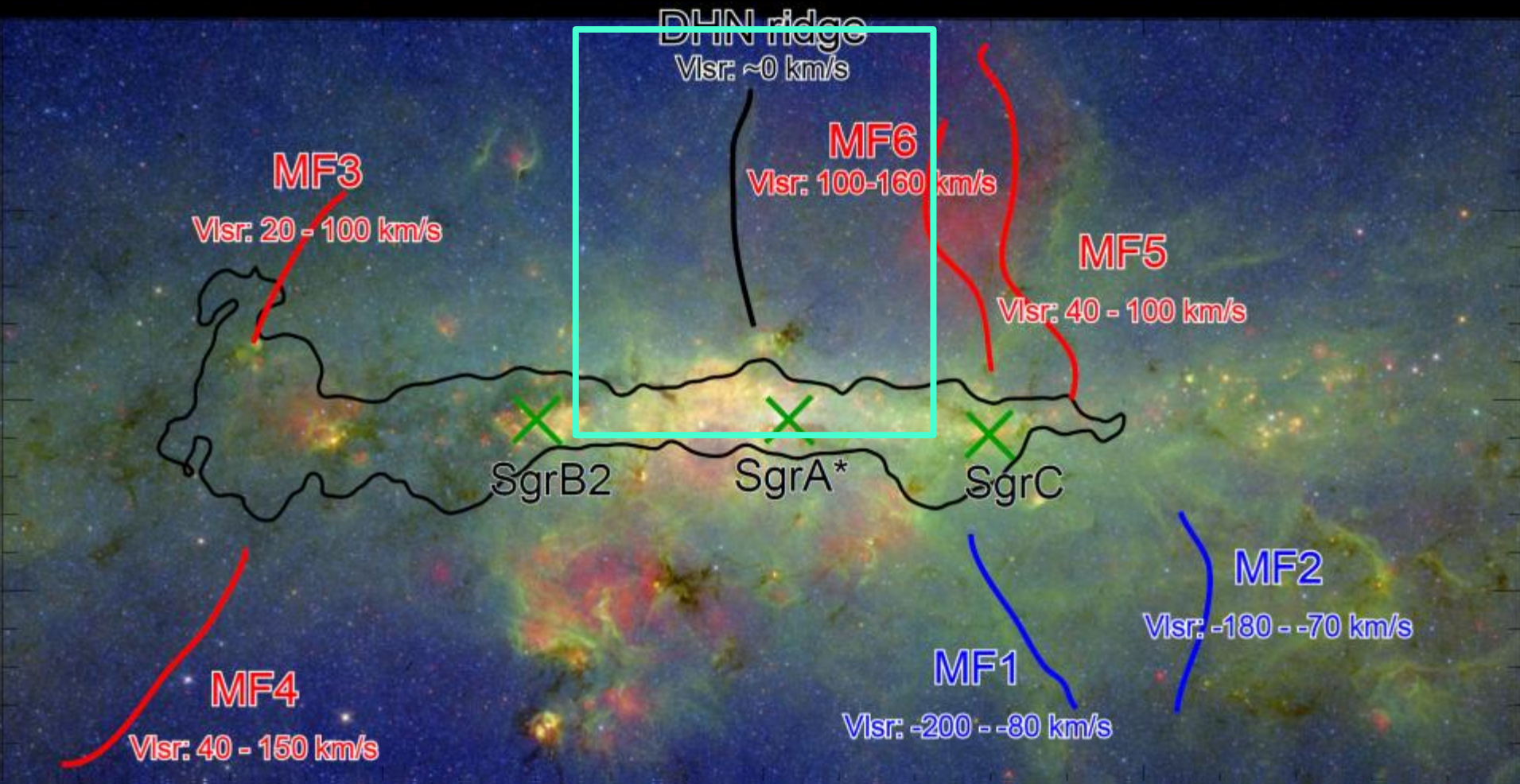
Molecular mass

halo : $\sim 10^6 M_{\odot}$ (typically $N(\text{H}_2) \sim 10^{21}$)

halo + MFs : $\sim 4 \times 10^6 M_{\odot}$ (**10% of the M_{CMZ}**)

Xfactor = 7.0×10^{19}
(Torii et al. 2010b)

Infrared counterparts to the MFs



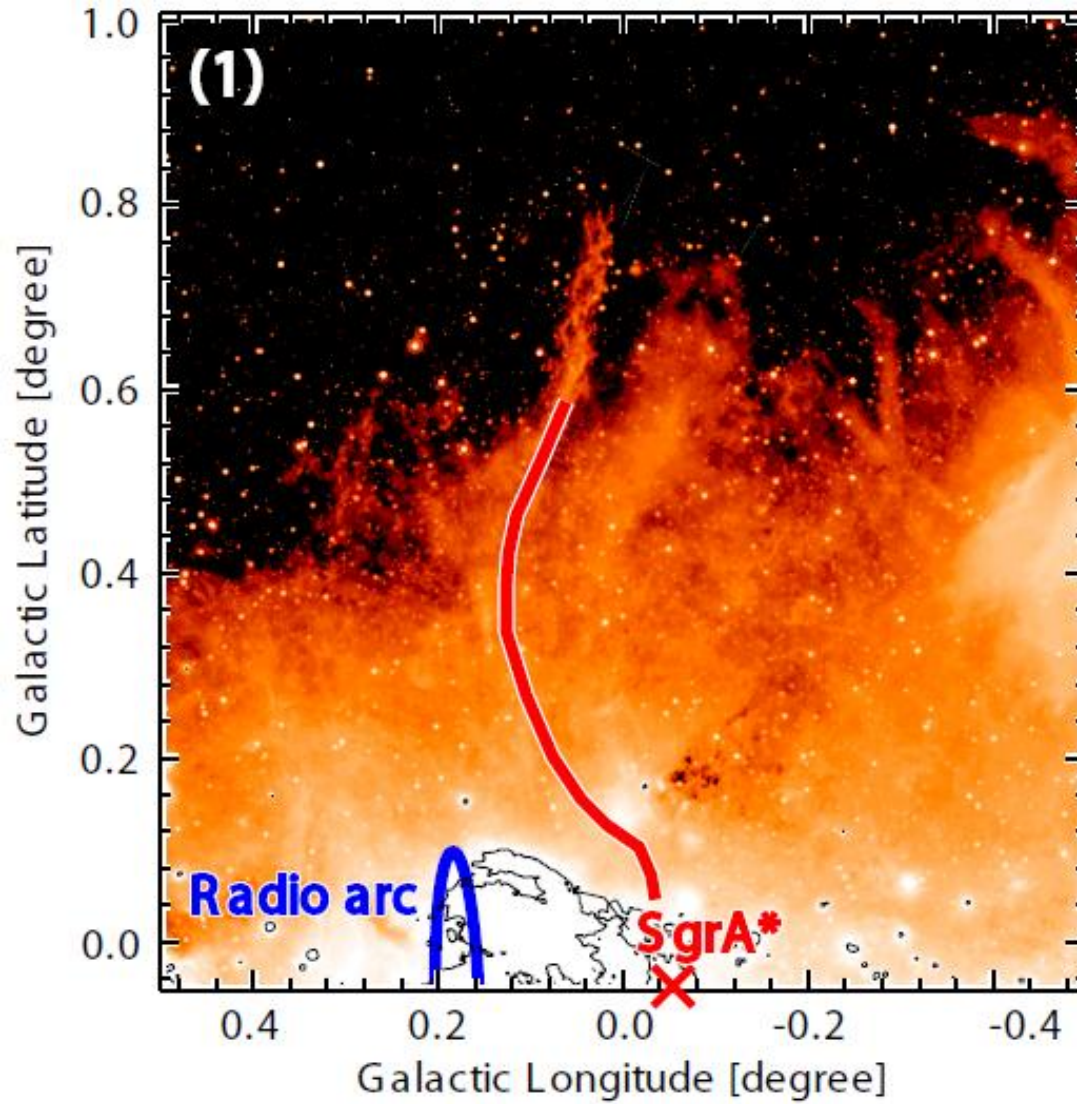
Spitzer IRAC+MIPS combine

Red: 24um, Green: 8um

Double Helix Nebula (DHN)

8

Spitzer 24 μ m



DHN

- ~ 0.7 deg from the GC
- seems to be a magnetic event

(Morris, Uchida & Do. 2006)

Possible origins..

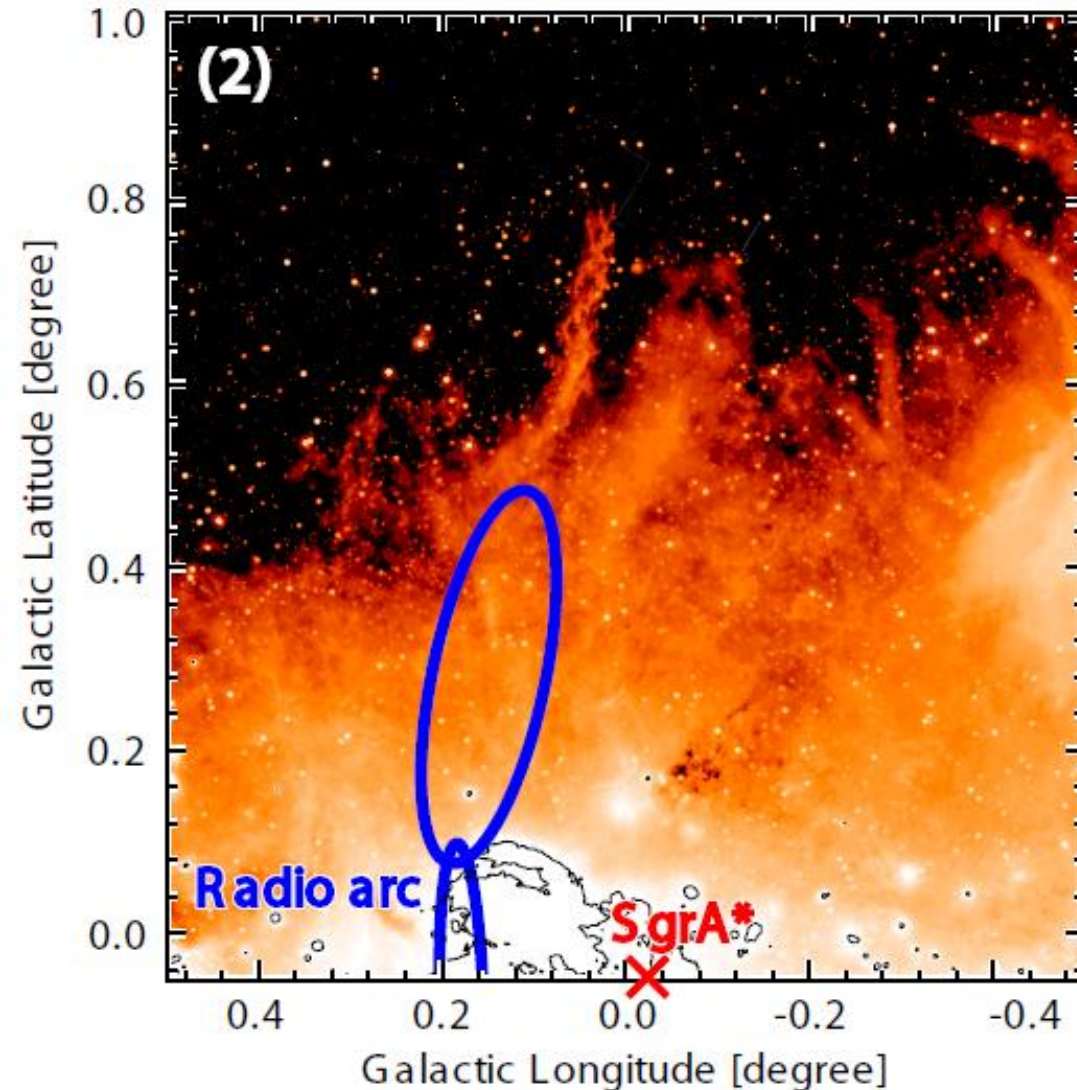
1. A torsional Alfvén wave launched by rotation of the magnetized CND
(Morris, Uchida & Do. 2006)

2. Polarized lobe extending to the radio arc
(Law et al. 2008; Tsuboi & Handa 2010)

Double Helix Nebula (DHN)

9

Spitzer 24 μ m



DHN

- ~ 0.7 deg from the GC
- seems to be a magnetic event

(Morris, Uchida & Do. 2006)

Possible origins..

1. A torsional Alfvén wave launched by rotation of the magnetized CND

(Morris, Uchida & Do. 2006)

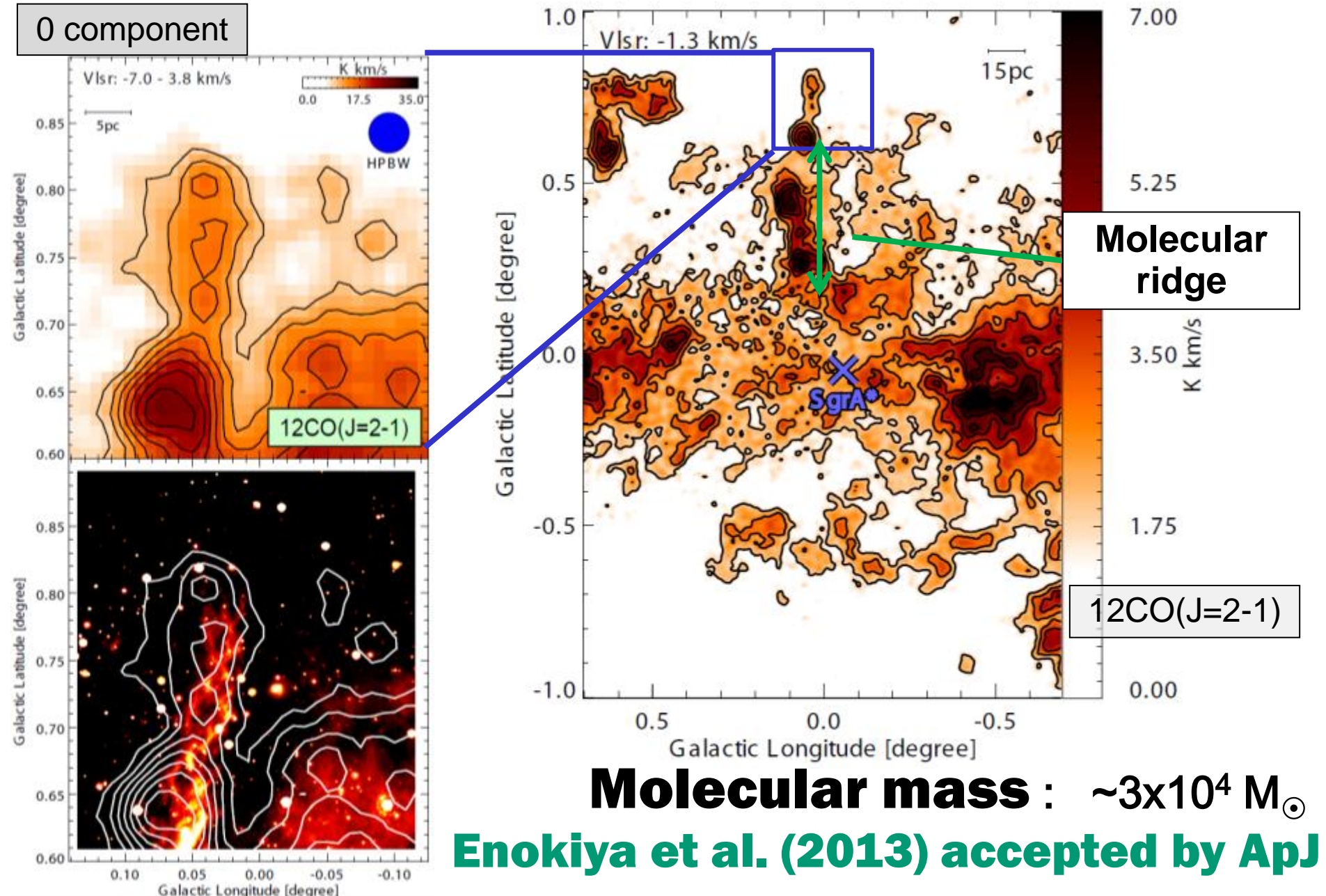
2. Polarized lobe extending to the radio arc

(Law et al. 2008; Tsuboi & Handa 2010)

Discovery of CO counterparts

10

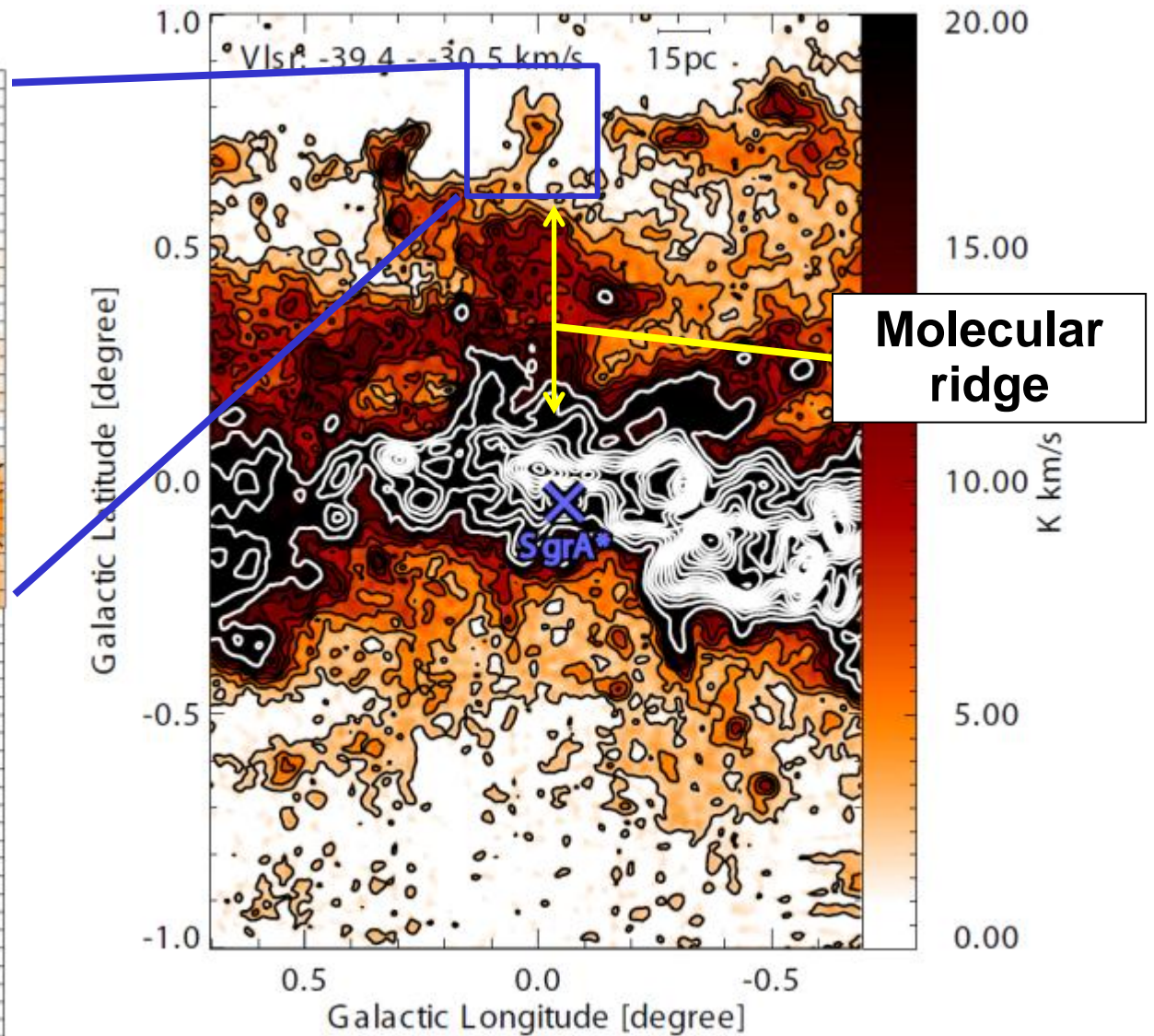
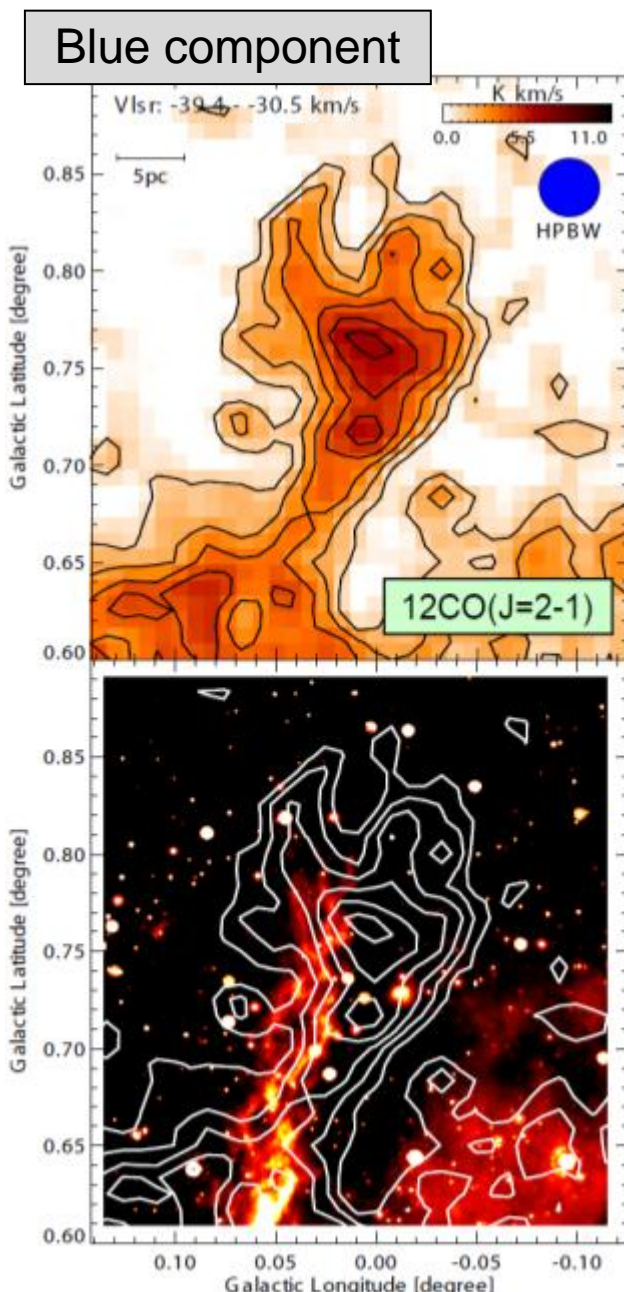
0 component



Discovery of CO counterparts

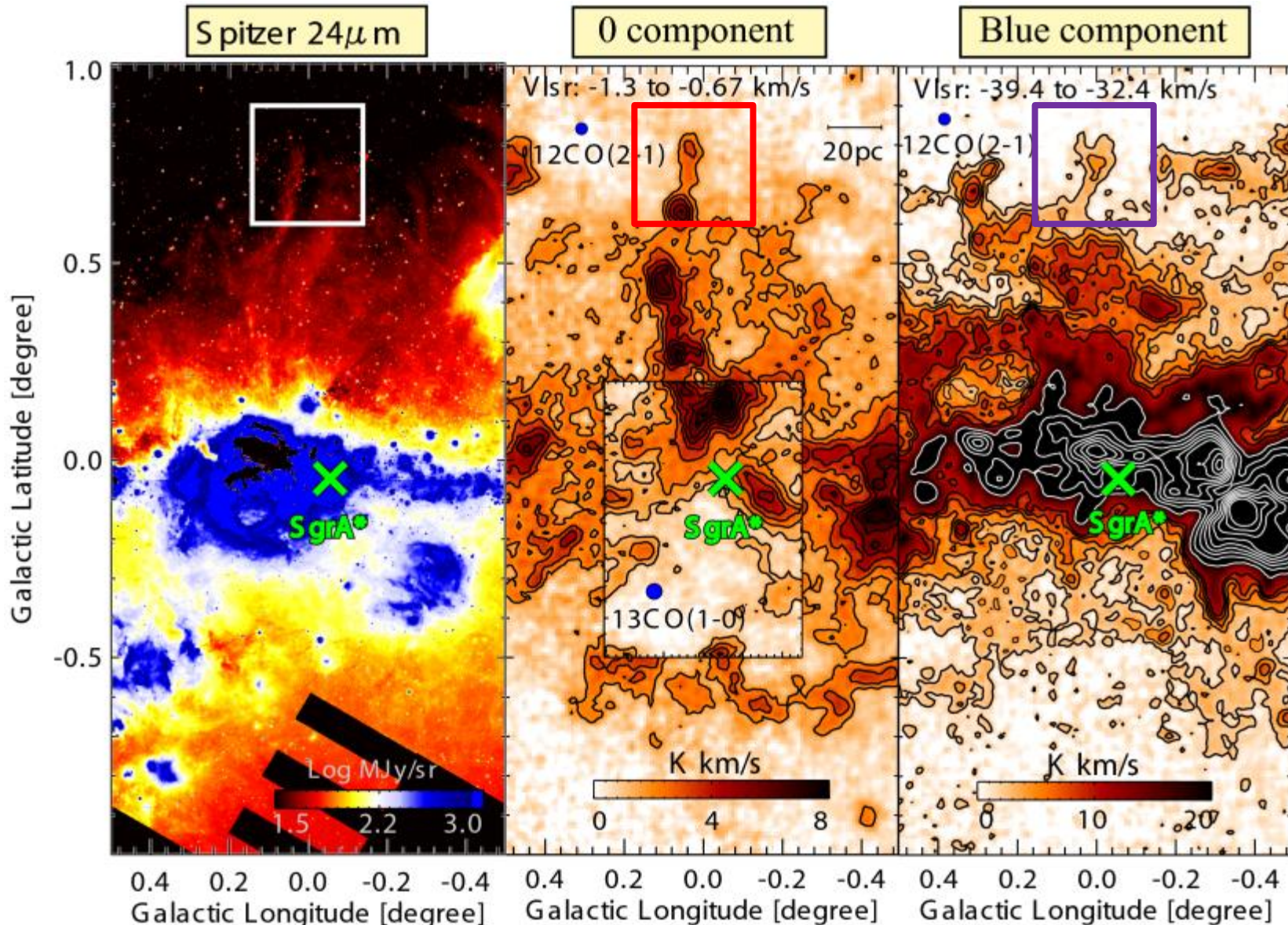
11

Blue component



Molecular mass : $\sim 10^5 M_{\odot}$
Enokiya et al. (2013) accepted by ApJ

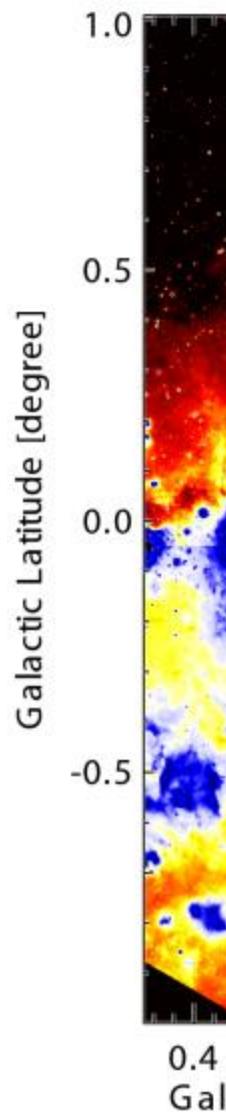
Discovery of the molecular columns 12



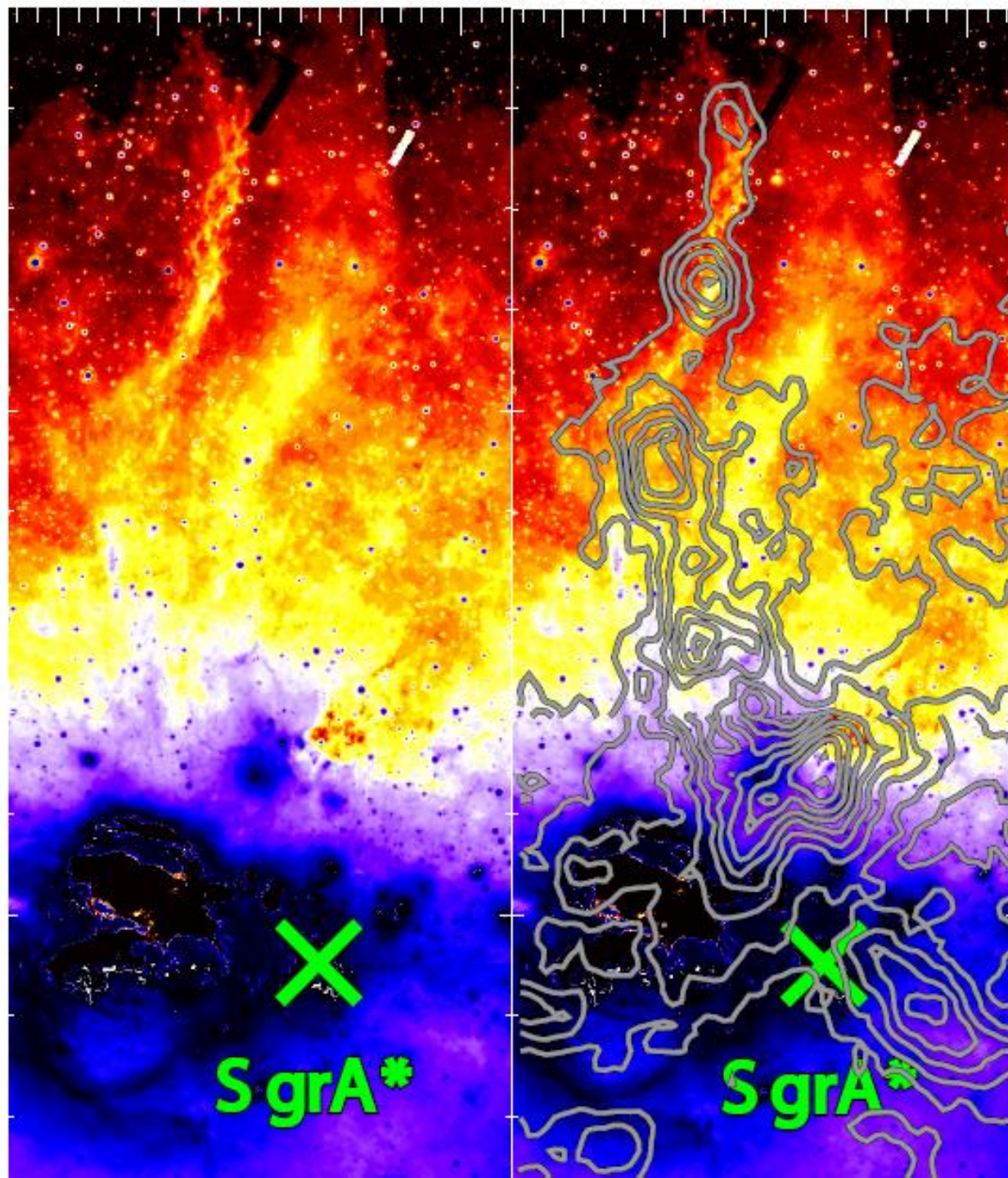
Each counterpart has a molecular column extending from SgrA*

Dis

mns 13

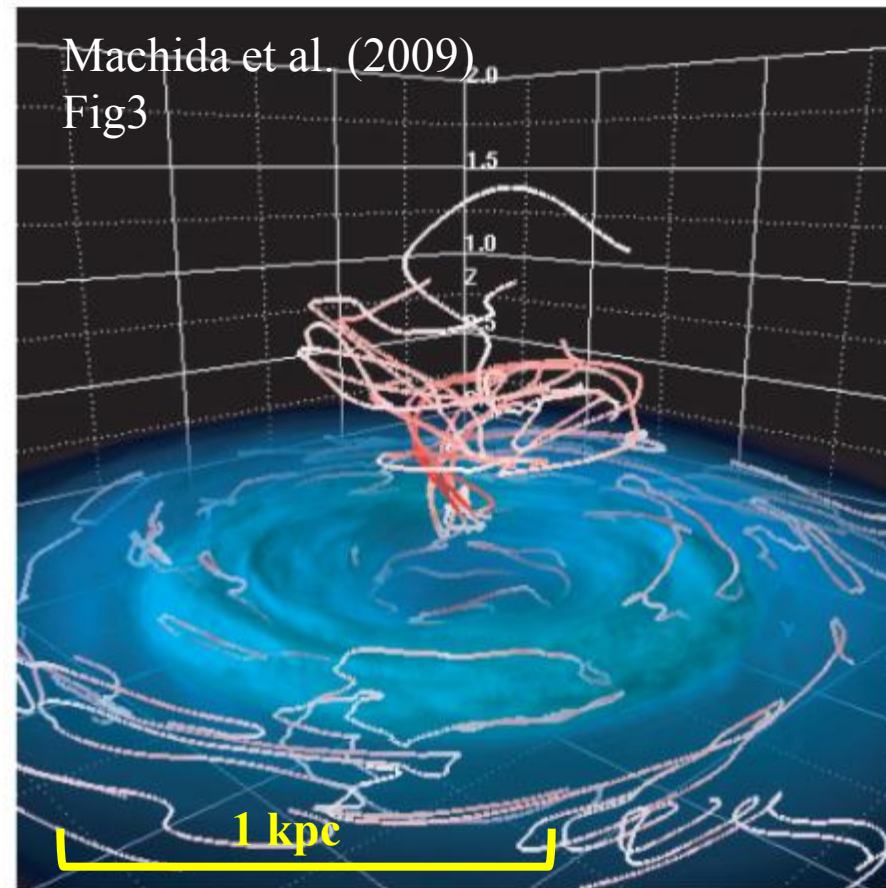
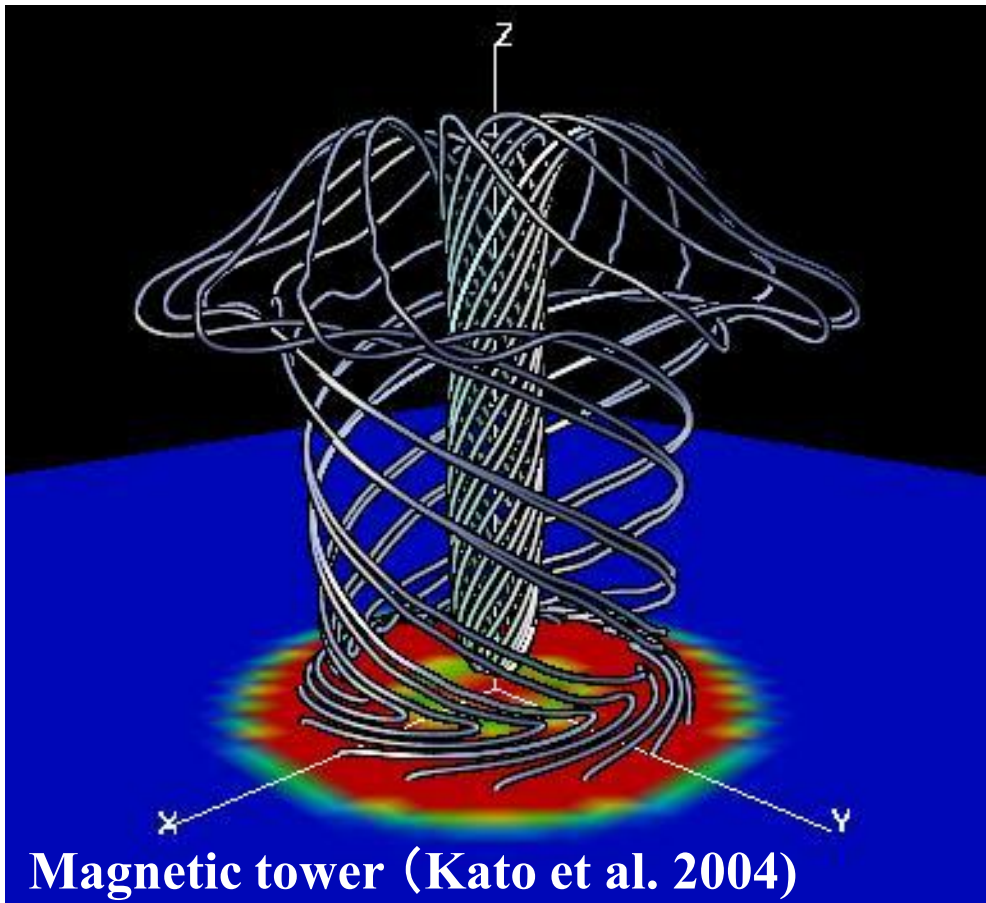


Each c



n SgrA*

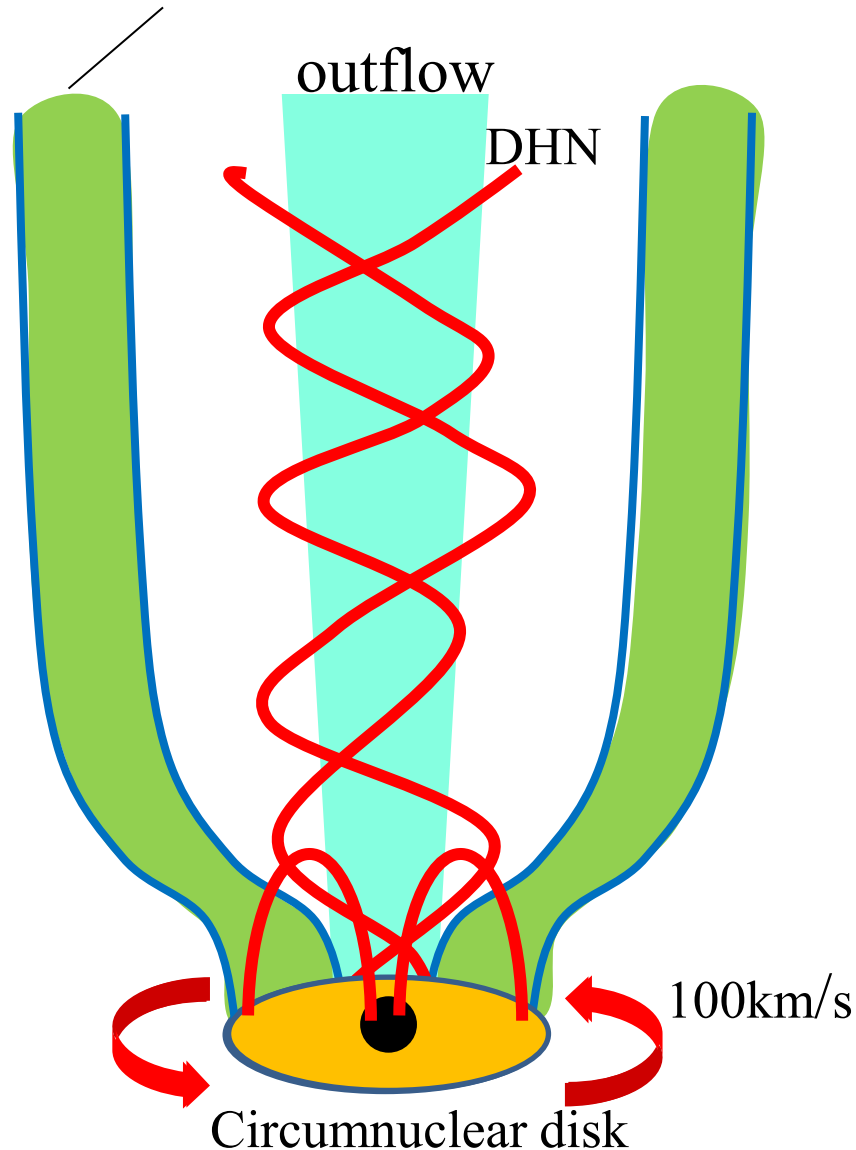
Discussion: Magnetic tower model¹⁴



Magnetic tower model. Rotation of the CND twists and ejects upward loops anchored to the CND

Formation mechanism of the DHN ¹⁵

Molecular column



Magnetic tower model

- The rotating CND twists the loop like magnetic fields and blows up the ISM upward from the Galactic plane
- Dense HI gasses grew up to molecular clouds by the compression of an outflow

There is a possibility that the DHN is a remnant of an outflow from the CND

Order estimation

Given timescale of formation of the DHN is ~ 5 Myr (10 times rotation of the CND), velocity of the ISM floated from the CND is estimated to be $\sim 25 \text{ km s}^{-1}$

| | Mass (M_{\odot}) | Radius (pc) | V_{rot} (km/s) |
|------------|-------------------------------------|-------------------------|--------------------------|
| CND | $\sim 4 \times 10^5$ _[2] | ~ 6 _[2] | ~ 80 _[3] |
| DHN | $\sim 10^5$ _[1] | ~ 2 _[1] | ~ 2 _[1] |

[1] Observational results

[3] Bradford et al.(2004)

[2] Oka et al. (2011)

Order estimation

Given timescale of formation of the DHN is ~ 5 Myr (10 times rotation of the CND), velocity of the ISM floated from the CND is estimated to be ~ 25 km s $^{-1}$

| | Mass (M_{\odot}) | Momentum flux (M_{\odot} km/s pc /yr) | Energy flux (erg/s) |
|-----|----------------------|---|-------------------------|
| CND | $\sim 4 \times 10^5$ | $\sim 5 \times 10^2$ | $\sim 3 \times 10^{34}$ |
| DHN | $\sim 10^5$ | ~ 1 | $\sim 3 \times 10^{31}$ |

➡ The CND origin can be explained by this scenario

Summary

- Newly detected CO features are extending up to $B=1^\circ$ from the CMZ and sorted to two types (Molecular filaments & halo)
- The total mass of them is $\sim 4 \times 10^6 M_\odot$ ($\sim 10\%$ of the M_{CMZ})
- Some MFs have their IR counterpart
- The DHN has two CO counterparts and both are located at the top of molecular ridges and at least 0 component is extending from the direction of SgrA*
- We propose a magnetic tower model from the CND as the origin of the DHN and CO counterparts. The estimated parameters are consistent with the model